

Stand-alone High Resolution Neutron Spectrometer

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Bachelor of Science

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## ABSTRACT

### Stand-alone High Resolution Neutron Spectrometer

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We are interested in improving neutron spectrometric capabilities in the energy range of 0.1 to 3 MeV. Current neutron spectrometry technologies do not have good energy resolution in this energy range without the use of either a complicated or a multiple-detector setup. We have developed a stand-alone detector with improved features in this energy range. The detector uses  $\text{Li}_6\text{Gd}(\text{BO}_3)_3$  : Ce crystal in a thin slab of polyvinyl toluene scintillator for capture-gated neutron detection and a second thin slab of solid plastic scintillator for proton recoil detection. Comparisons of proton recoil pulse area to measured time-of-flight data and theoretical neutron energy spectra indicate a useful correlation with neutron energy. The spectrometer we have developed functions as an initial prototype.

Keywords: neutron detection, neutron spectrometry, proton recoil, capture-gated, LGB, stand-alone

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# Chapter 1

## Introduction

### 1.1 Purpose of Neutron Detection and Spectrometry

The ability to detect neutrons and measure their energy is important for both research and radiation safety applications. Nuclear reactions produce characteristic particles at distinct energies. Studying the particles and their energies provides insight into the nuclear processes that occur during the reaction. Good detection and spectrometric technologies improve our ability to identify and characterize unknown or low-rate reactions. This is useful for studying anomalous results in nuclear research as well as for detecting and inventorying nuclear material. Reliable and accurate detectors are also critical for monitoring radiation in work environments such as reactor or accelerator facilities.

### 1.2 Laboratory Nuclear Astrophysics Research

The Laboratory Nuclear Astrophysics Research (LNAR) group at Brigham Young University is interested in studying stellar-like nuclear reactions, typically using d-d fusion. Typical nuclear astrophysics research is done using reactants with energies much greater than those actually found

in stars. This is done to push the resulting signals above background levels. As the energy of the reactants is increased, the probability of fusion also increases as it becomes easier and easier to tunnel through the Coulomb barrier caused by electrostatic repulsion between positively charged nuclei. Using higher energy reactants creates a fundamental difference between nuclear reactions in stars and nuclear reactions in the laboratory. As detector capabilities have improved in recent decades, studies have been done using reactant energies closer to those actually found in stars. Many of these studies have yielded unexpected results. They have found higher-than-expected rates of fusion in condensed matter environments such as deuterium-loaded metal foils. This effect has in the literature been attributed to electron screening potentials which effectively lower the energy barrier to fusion. The effect varies from material to material and is especially strong in some metals. For a summary of screening potentials measured in several studies, see Table A.1 in Appendix A. Improving neutron spectrometry capabilities enables us to more reliably measure the rate of these reactions in the laboratory as we study potential explanations for the unexpected phenomenon.

### **1.3 Thesis Overview**

My purpose in writing this is both to describe the development of a specific detector and its capabilities and also to provide a brief introduction to many of the concepts and methods used in the LNAR group's neutron detection research. I will provide an overview of common neutron detection and spectrometric techniques and a few existing technologies utilizing those techniques. I will discuss the development of our neutron spectrometer and several important features of the detector. I will also discuss how we acquire and analyze data from the spectrometer. Finally, I will compare the measured energy spectrum from our spectrometer to time-of-flight data and to the test energy spectrum and discuss the detector's usefulness and applications.

# Chapter 2

## Methods

### 2.1 Neutron Detection

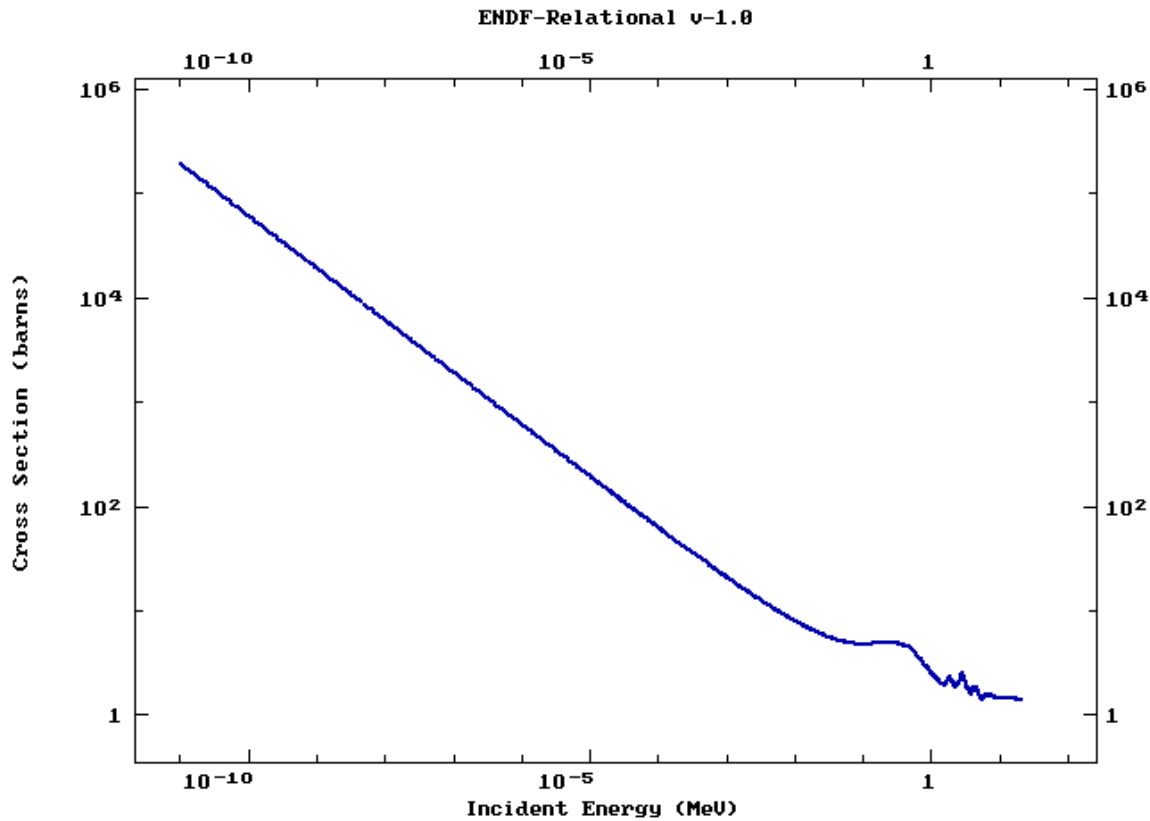
Detectors do not yet exist which are capable of detecting neutrons directly. Instead, it is necessary to observe the effects of neutron interactions, typically through derived charged particles which are directly detectable by current technology (Knoll 2010). Several techniques have been developed over the decades to detect the charged particles produced by neutron interactions. Detection methods differ depending on the energy of the neutrons to be detected. These techniques can be divided into categories by the energy range (also called temperature) of the neutrons they are capable of detecting. There are several commonly used neutron temperature ranges but in this thesis I will address only two of them: fast neutrons and slow neutrons.

#### 2.1.1 Fast Neutrons

Fast neutrons have sufficient energy ( $>1$  MeV) that they do not have large probabilities of interaction, expressible in terms of cross section, with other particles via nuclear reactions. (Cross section is measured in barns, as in the broad side of a barn. Though it must be a rather small barn as 1 barn



is equal to  $10^{-24}$  cm<sup>2</sup>.) For example, <sup>10</sup>B has a very low cross section for neutrons with energy above about 1 MeV (see Fig. 2.1).



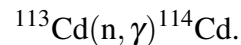
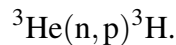
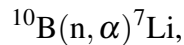
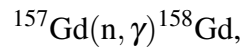
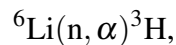
**Figure 2.1** The total neutron cross section of <sup>10</sup>B versus energy (ENDF 2016).

Instead, fast neutrons interact with other particles via elastic collision. When a neutron collides with a heavy (compared to the mass of a neutron) particle or nucleus the neutron will bounce off, imparting little of its energy to the larger particle (think ping-pong ball hitting a bowling ball). However, when a neutron hits a similarly sized particle, such as a proton, it transfers a large portion of its energy to the proton, up to the total energy of the impinging neutron (think ping-pong ball hitting a ping-pong ball). The newly energized proton then moves through scintillating material creating a light output detectable using photomultiplier tubes (PMTs). The PMT then converts the light into an electrical pulse. Since the energy of the neutron was largely transferred to the proton

it is possible to infer energy information about the original neutron from the detected proton recoil pulse. As the neutron transfers energy to protons, it is moderated to a sufficiently slow speed that the probability of interaction with other particles via a nuclear reaction becomes significant.

### 2.1.2 Slow Neutrons

Slow neutrons do not have sufficient energy to transfer to protons for the recoiled protons to produce detectable light in a scintillating material. Instead, they are detected by the production of charged particles via nuclear reactions such as neutron-capture:



(The standard notation  $a(b, c)d$  indicates a particle  $b$  hitting a target  $a$ , producing a particle  $c$  and a product  $d$ .) Many successful neutron detectors exist utilizing these reactions in various configurations including  ${}^3\text{He}$  tubes,  ${}^6\text{Li}$  glass,  $\text{Li}_6\text{Gd}(\text{BO}_3)_3$  crystal (LGB), and  ${}^{113}\text{Cd}$  sheets.

The light output from LGB (the scintillator we use in our detector) neutron capture reactions typically occurs over a longer period of time (hundreds of nanoseconds) than recoil reactions (tens of nanoseconds) resulting in wider pulses. The distinctly wider pulses are useful for capture-gated neutron detectors such as our spectrometer. However, unlike proton recoil pulses, capture pulses contain energy information from the reaction products, rather than the original neutron. Neutron energy information was lost in the moderation process necessary for a capture reaction to occur. A combination of recoil and capture pulses is useful for retaining energy information while simultaneously identifying true neutron detection events.

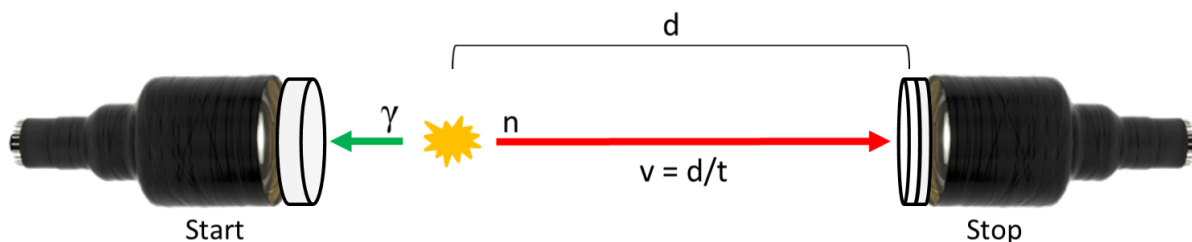
## 2.2 Neutron Spectrometry

In addition to being able to detect neutrons, it is important to be able to measure their energy. This can be difficult as the energy information of the neutron is often lost in the detection process. However, several technologies exist that are capable of neutron energy measurements. Each method has its own strengths and weaknesses regarding resolution, efficiency, portability, and useful energy range.

### 2.2.1 Time-of-flight

A common technique used to measure neutron energy is time-of-flight. The reactions that produce neutrons also produce charged particles such as gammas, alphas, or fission fragments. The time-of-flight technique uses these other products to create a start signal when the neutron-producing reaction has occurred. The most accurate time-of-flight spectrometers use charged particle spark chambers to detect the charged fission fragments to produce a start signal.

The time-of-flight technique requires the use of more than one detector: one near the neutron source to detect the charged particle products and a second at a known distance (typically on the order of meters) to detect the arrival of the neutron, producing a stop signal (see Fig. 2.2).



**Figure 2.2** A diagram showing the important features of a time-of-flight spectrometer.

By measuring the difference in time between the start signal and the stop signal you can calcu-

late the kinetic energy of the neutron using

$$E = m_n c^2 \left( \frac{1}{\sqrt{1 - \frac{(d/t)^2}{c^2}}} - 1 \right), \quad (2.1)$$

where  $m_n$  is the mass of a neutron,  $c$  is the speed of light,  $d$  is the distance from the source to the stop detector, and  $t$  is the difference in time between the start pulse and the stop pulse. Neutrons do not move at the speed of light, but they do travel at sufficiently high speed that it can be necessary to use relativistic mechanics to calculate their kinetic energy (neutrons with energy greater than around 939 MeV are considered relativistic, so the effect for our energy region of interest, 0.1 to 3 MeV, is negligible) (Knoll 2010).

Fission fragments are most commonly used to produce a start signal because they are always produced alongside neutrons, do not attenuate significantly through air, and are easy to detect promptly after they are produced. Additionally, the detection of a fission fragment is a better indicator of a neutron-producing reaction. For example,  $^{252}\text{Cf}$  decays 97% of the time via alpha decay, which does not produce any neutrons. Only 3% of the decays are spontaneous fission, producing neutrons and fission fragments. Gating the start detector on fission fragment detection events immediately removes 97% of the noise that would otherwise be present if gammas were used for the start signal.

Care must be taken that the start detector is not so close to the source that the recovery time of the detector (set by the type of scintillator and photomultiplier tube used) is longer than the time between pulses, effectively blinding the detector. Increasing the distance from the stop detector to the source will increase the resolution of the spectrometer as the longer path length will spread out the distribution of discernible flight times. However, for sources that emit neutrons in every direction equally (isotropic), increasing the path length also decreases solid angle and thus the neutron flux through the detector, requiring longer run times.

Time-of-flight systems are especially useful in permanent installations like reactor or accelerator facilities. Such facilities typically have the necessary space and resources to develop and

maintain a time-of-flight spectrometer as well as the need for a high resolution spectrometric setup. Time-of-flight spectrometers provide very high resolution measurements. Additionally, they can use a variety of detection methods since the two detectors need only be capable of detecting the start particle and the neutrons rather than being able to independently measure their energy. However, time-of-flight is not an acceptable solution where portability is a necessary requirement or where it is not feasible to have a detector at the source.

### 2.2.2 Bonner Sphere

The Bonner sphere spectrometer is more portable than a time-of-flight setup. This spectrometer uses several spheres of moderating material of various diameters (see Fig. 2.3).

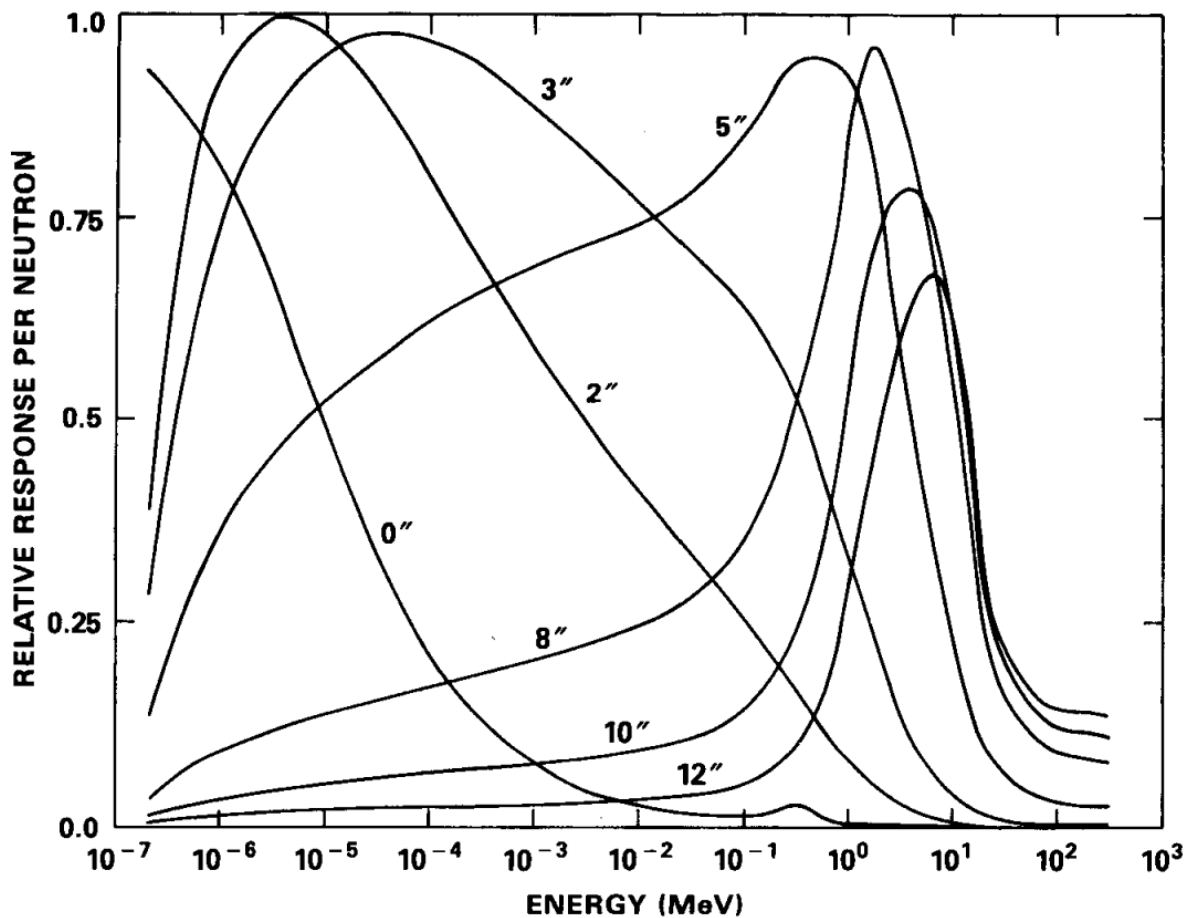


**Figure 2.3** A Bonner sphere neutron spectrometer (ElseNuclear 2016).

When a neutron enters each sphere, it may be moderated to a sufficiently low energy that a detector in the center of the sphere can capture and detect the neutron (typically a  $^3\text{He}$  detector).

By using a variety of sphere diameters, neutrons that fall within different energy windows can be detected by each sphere. For example, a neutron entering a small sphere would be sufficiently moderated and subsequently detected only if it does not have so much energy that it passes completely through the detector before it is sufficiently moderated. Increasing the diameter of the moderating material allows higher and higher energy neutrons to be detected. However, as the diameter of the sphere increases, the probability of the neutrons successfully making it to the detector in the center decreases, necessitating the use of both small and large spheres.

Each sphere will produce an output spectrum which can then be combined with the spectra from the other spheres into a single energy spectrum across the total energy range of the spheres used. (see Fig. 2.4).



**Figure 2.4** Energy spectra from a Bonner sphere spectrometer. Each curve represents the energy spectrum from a different diameter sphere (Knoll 2010).

This process is known as unfolding the spectra. The resolution of the spectrometer is limited by the number of different sizes of spheres used and the capability of the unfolding code to accurately assemble the total neutron energy spectrum. While more portable than a time-of-flight spectrometer, the Bonner sphere spectrometer requires many detectors, each with power supplies, signal processing electronics, and their own opportunities for component failure. A variation of the sphere design has been developed that uses nested cylinders rather than several spheres (Dubeau et al. 2012) (see Fig. 2.5).



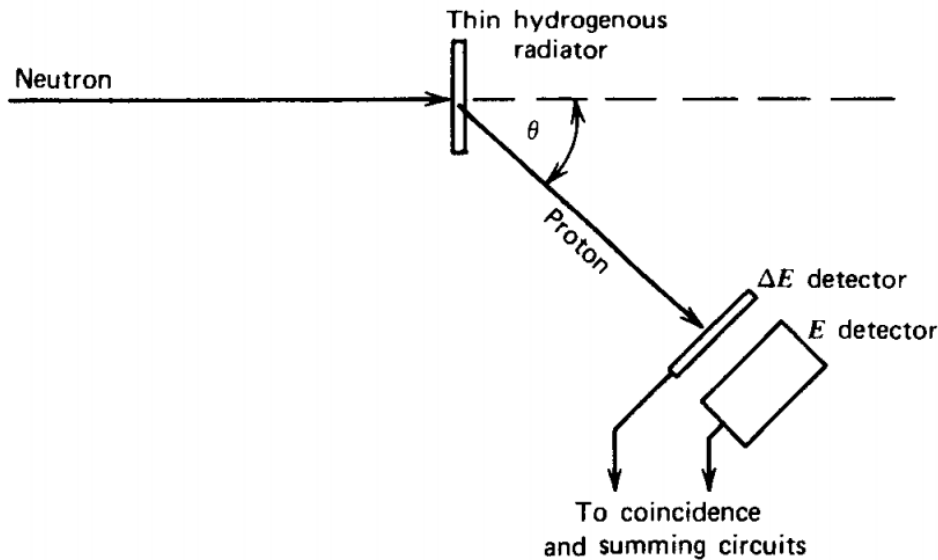
**Figure 2.5** Moderating material from a nested neutron spectrometer (Nicholishiell 2013). Image used under the Creative Commons Attribution-Share Alike 3.0 Unported license.

This reduces the number of detectors and increases the portability substantially, but does not improve the resolution capabilities of the spectrometer.

### 2.2.3 Proton Recoil Telescope

Proton recoil telescopes are capable of performing much higher resolution neutron spectrometry without the need of a second detector, at the cost of efficiency. A collimated beam of neutrons is directed into a very thin hydrogenous radiator. When a neutron collides with a proton, the proton is expelled from the radiator at an angle related to the portion of the energy of the impinging neutron that was transferred to the proton (see Fig. 2.6).





**Figure 2.6** A diagram showing the important features of a proton recoil telescope (Knoll 2010).

By measuring the energy of the recoiled proton, the energy of the original neutron can be calculated using

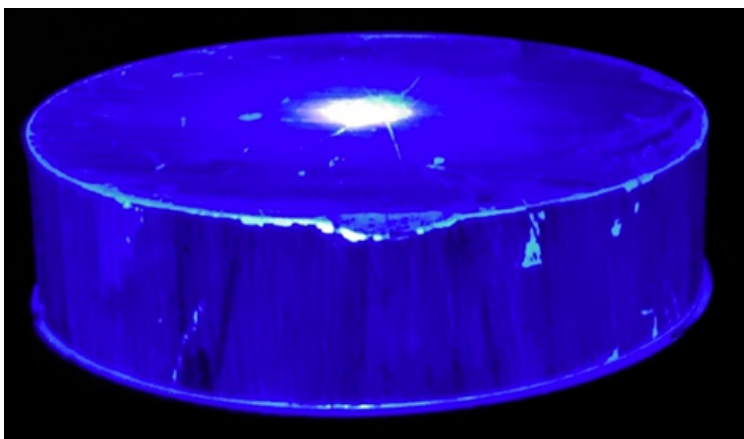
$$E_p = E_n \cos \theta.$$

The resolution of the detector is only limited by the resolution of the proton spectrometer used to detect the recoiled proton and the width of the detector. Using a narrow detector reduces the possible values of  $\theta$  but also reduces the efficiency of the detector. Typical proton recoil telescopes have very poor efficiency, on the order of 1 in  $10^5$  (Knoll 2010). In addition to poor efficiency, this spectrometer requires that the travel path from the radiator to the proton detector be evacuated, creating additional engineering difficulties. Our detector is a sort of proton recoil telescope with greatly improved efficiency and an enforced trigger on events with an angle near zero for protons ejected from the radiator. We also couple the detector directly to the hydrogenous radiator, eliminating the need for an evacuated travel path. More details are given in the following sections.

## 2.3 Detector Development

### 2.3.1 Scintillators

Scintillators are materials that emit light when charged particles move through them, depositing energy and exciting molecules (see Fig. 2.7).



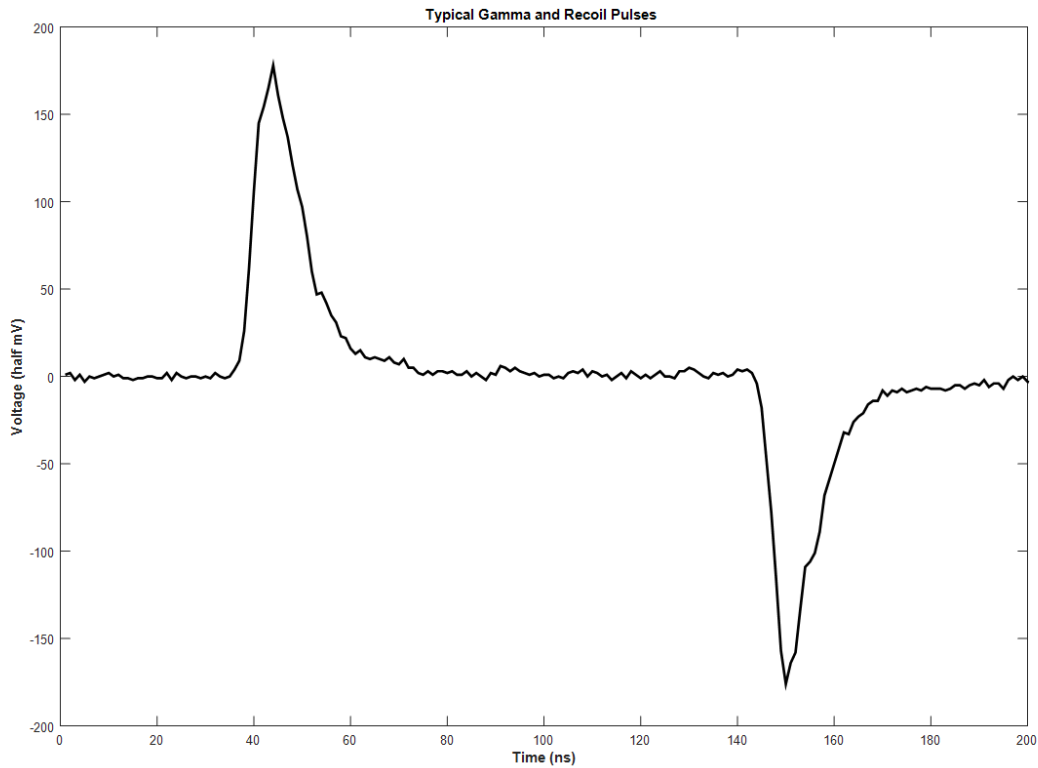
**Figure 2.7** EJ-200 plastic scintillator being excited with an ultraviolet laser.

There are two major types of scintillators: organic and inorganic. We use both types in our spectrometer.

Organic scintillators are composed of hydrocarbons with some kind of scintillating dopant such as anthracene. They may be liquid or solid. Liquid organic scintillators have historically been better suited for pulse shape discrimination (PSD) (Zaitseva et al. 2012). The organic scintillator we use in our spectrometer is EJ-200 solid plastic scintillator (see EJ-200 Plastic Scintillator in Appendix B). This particular plastic scintillator is very fast and thus well-suited for high resolution timing measurements. While liquid organic scintillators are better suited for PSD, solid plastic scintillator is much more practical for use in our spectrometer as it can be fairly easily machined into thin cylindrical shapes. (Care must be taken when machining and cleaning plastic scintillator that the optical surfaces are not scratched or marred in any way as this reduces light output and

consistency. Solvents (including isopropyl alcohol) should never be used to clean plastic scintillator.) Organic scintillators emit light when excited by charged particles or high energy photons moving through them, depositing energy. When the excited scintillator de-excites it emits photons of a wavelength characteristic of the scintillator type.

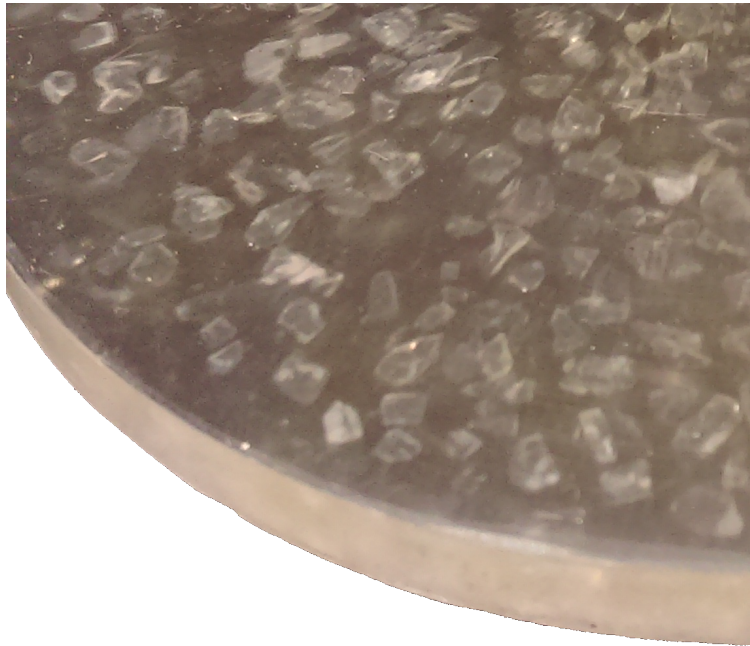
The plastic scintillator in our spectrometer serves two primary purposes. First, the hydrogen in the organic scintillator acts as moderating material to slow impinging neutrons sufficiently that they can be captured in the inorganic crystals in the detector. Second, as the neutrons are slowed through collisions with protons, the newly-energized protons move through the scintillator depositing energy and producing light pulses that can be detected by the photomultiplier tube. These pulses serve as a stop signal for a time-of-flight setup and also retain much of the energy information of the original neutron. This preserved energy information is central to our spectrometer design and will be discussed in more detail later. Although the timing of the plastic scintillator is very fast, pulses resulting from gammas and neutrons are too similarly shaped for reliable pulse shape discrimination (see Fig. 2.8).



**Figure 2.8** Typical gamma and recoil pulses from EJ-200 scintillator. The positive pulse is from a gamma and the negative pulse is from a recoil proton (probably: identified by proximity in time to a capture pulse outside the plotted window).

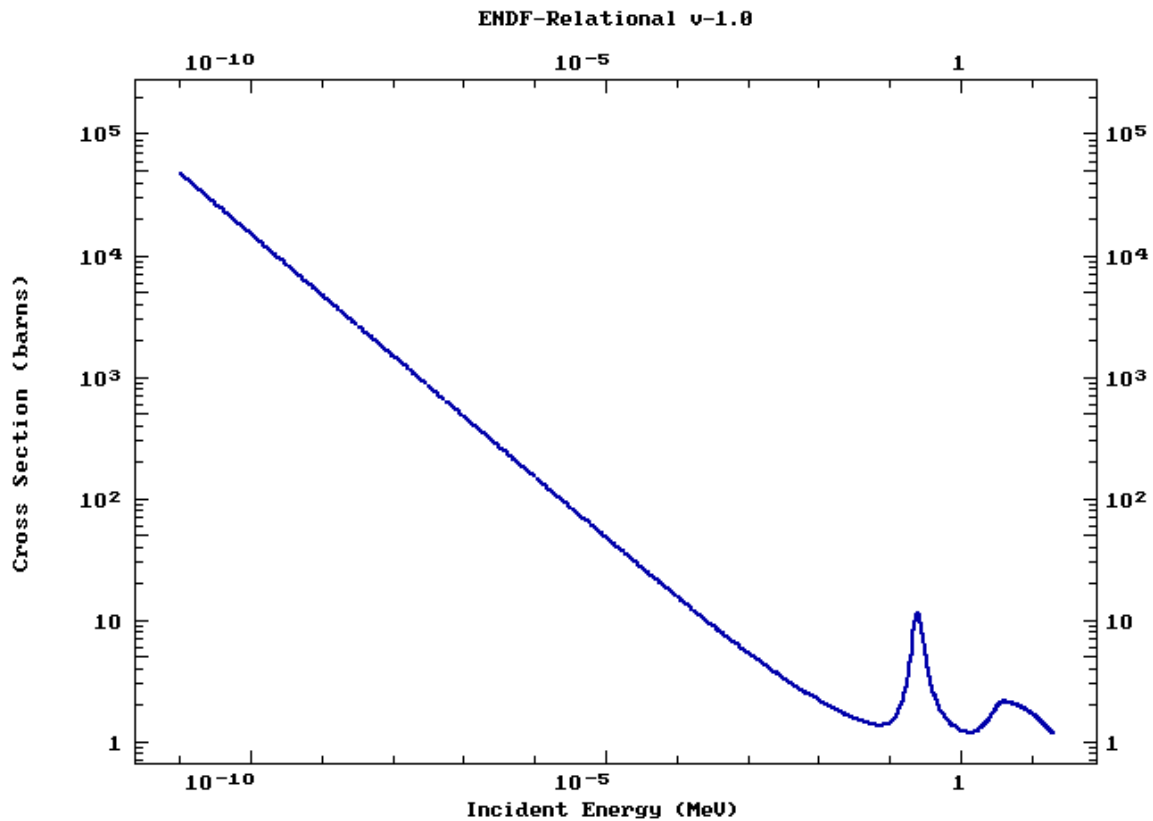
Thus, another method is required to ensure triggering on true neutron detection events.

The second major subtype of scintillators is inorganic scintillators. Inorganic scintillators are typically materials that emit light when a neutron-capture reaction occurs such as in  ${}^6\text{Li}$  glass. The inorganic scintillator used in our detector is  $\text{Li}_6\text{Gd}(\text{BO}_3)_3 : \text{Ce}$  (LGB) in crystalline form where the LGB captures the neutron and the cerium dopant acts as the scintillating material (see Fig. 2.9).



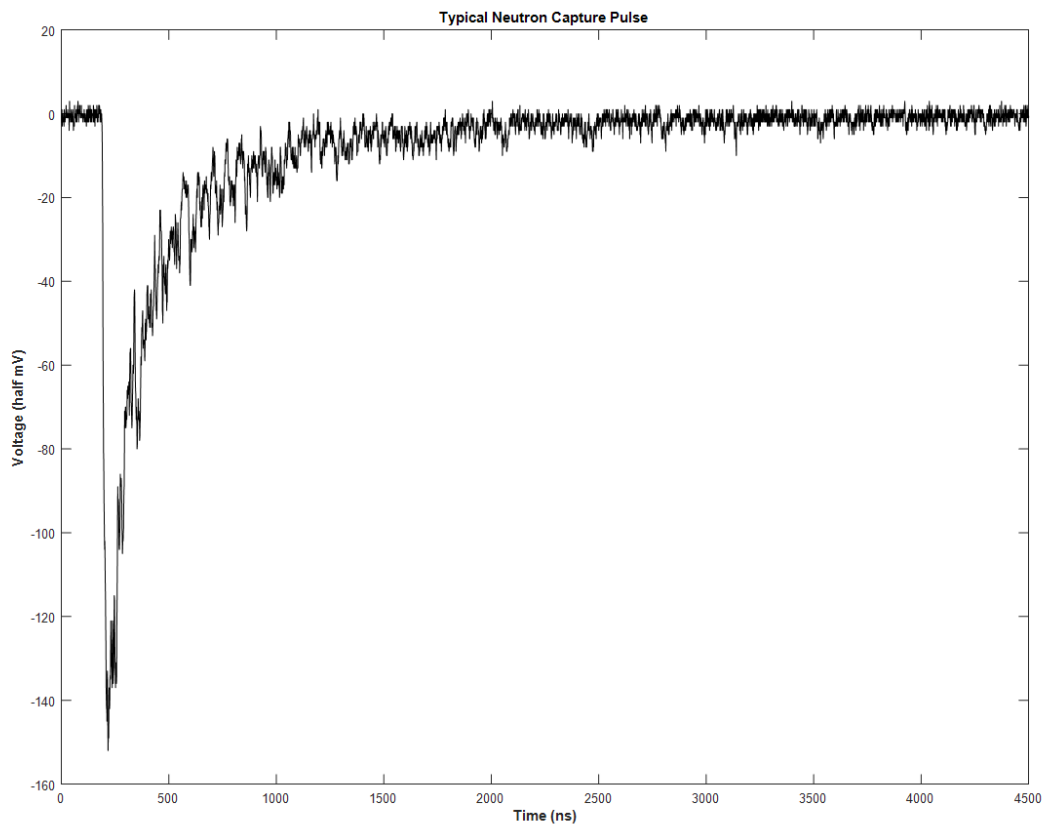
**Figure 2.9** Crystalline LGB inorganic scintillator suspended in EJ-200 plastic organic scintillator.

As neutrons collide with protons in moderating material, such as organic scintillator, their energy decreases and their probability of interaction with inorganic scintillators increases (see Fig. 2.10).



**Figure 2.10** The total neutron cross section of  ${}^6\text{Li}$  versus energy (ENDF 2016).

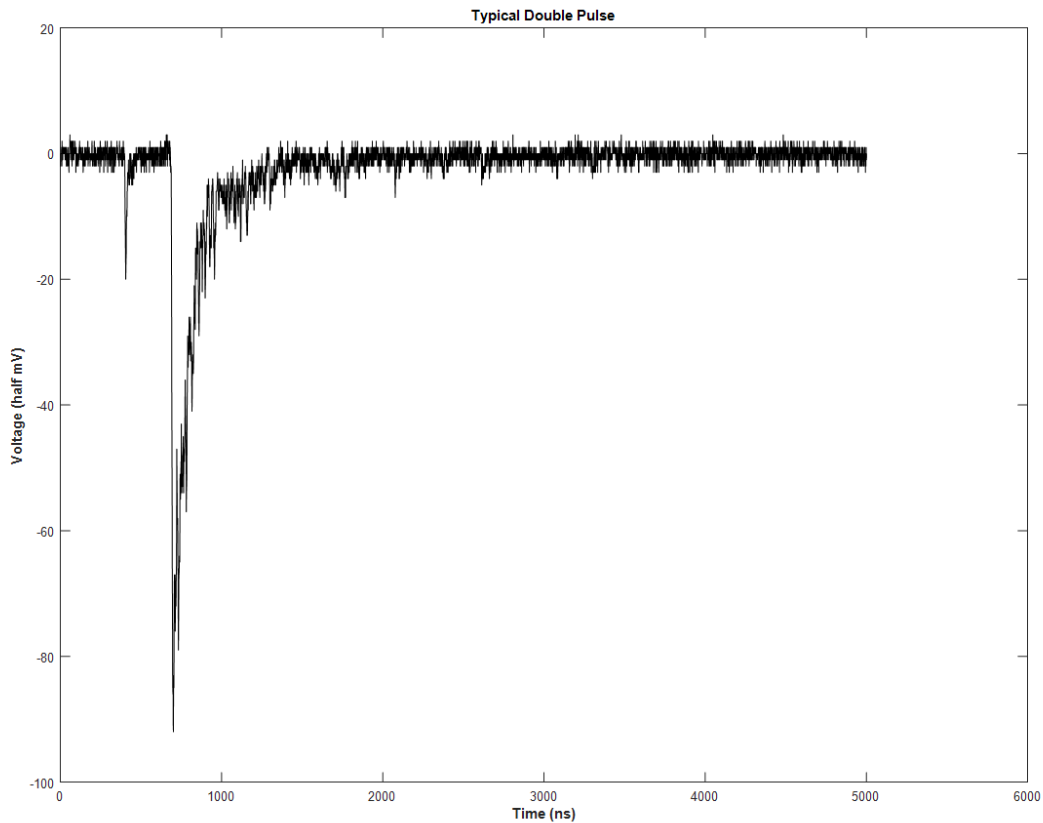
The neutron-capture reactions in  ${}^6\text{Li}$  and  ${}^{10}\text{B}$  in LGB produce very wide and distinct light output pulses (see Fig. 2.11) ( ${}^6\text{Li}$  in glass or plastic produces narrow pulses that are difficult to distinguish from gammas).



**Figure 2.11** A typical neutron capture pulse in LGB.

We gate our spectrometer on these wide pulses in order to discriminate against gammas.

Our detector produces signature double pulses when a neutron is detected: one from proton recoil in the organic scintillator and one from neutron-capture in the inorganic scintillator (see Fig. 2.12).



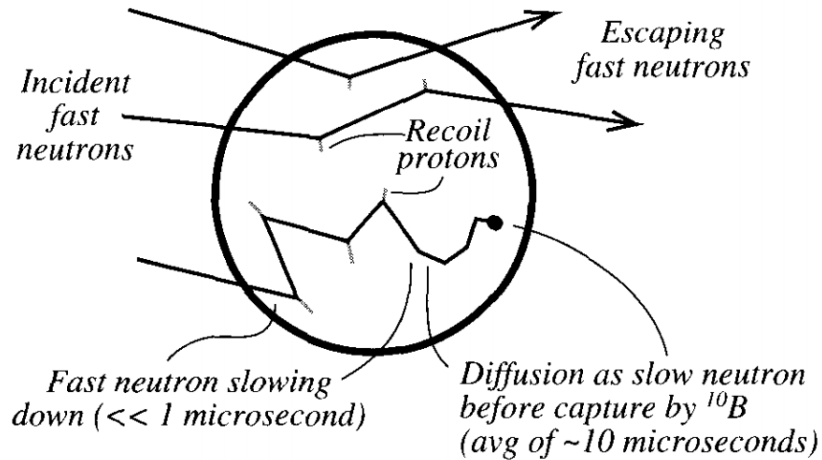
**Figure 2.12** A typical double pulse seen by our detector. The first narrow pulse is from the proton recoil. The second wide pulse is from capture in  ${}^6\text{Li}$  or  ${}^{10}\text{B}$ .

This ensures that we mostly observe valid neutron events and, only occasionally, other coincidental events such as gammas. This significantly reduces the need for pulse shape discrimination. However, because the plastic scintillator is also sensitive to gammas, it is fairly common for pulses from gammas to be misidentified as proton recoil pulses. This can be minimized by enforcing a timing window between the supposed proton recoil and the neutron capture pulses.

In addition to gating our detector on neutron events, we enforce mostly one-to-one energy transfers in the initial proton recoil. On average, neutrons do not impart all of their energy to a proton in a single collision. We enforce one-to-one energy transfers using a combination of



capture-gating and very thin scintillators. When a proton recoils off a neutron, the interaction follows standard elastic scattering physics (see Fig. 2.13).

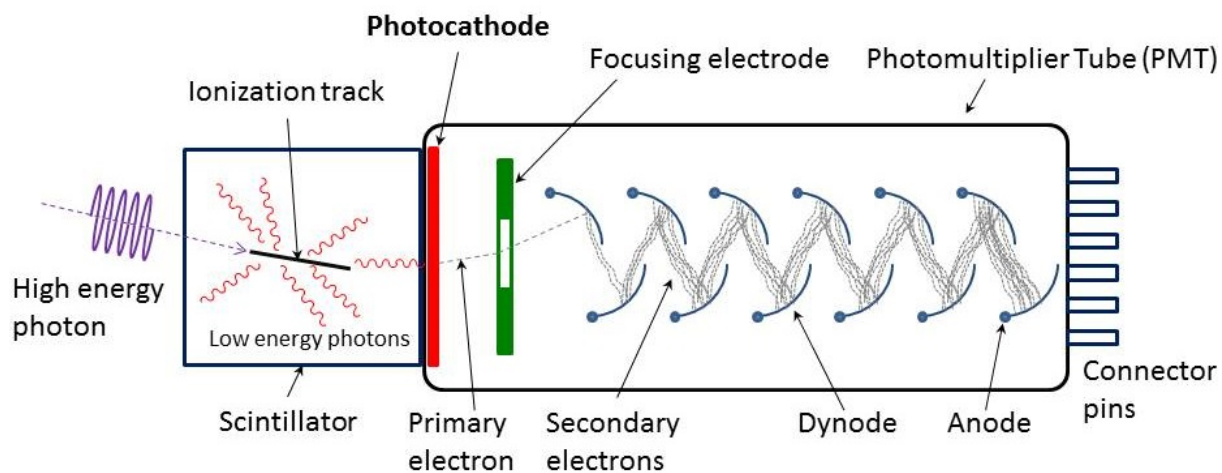


**Figure 2.13** Possible results of proton recoil interactions (Knoll 2010). Slowed neutrons in our detector can also capture in <sup>6</sup>Li in addition to the <sup>10</sup>B shown in the figure. In our detector, the diffusion time is closer to an average of a few hundred nanoseconds.

The closer to a one-to-one energy transfer, the closer the neutron will recoil ninety degrees from its original path. When using very thin scintillators, if the neutron does not recoil ninety degrees it will pass out of the scintillator before it can either recoil again or be captured in the LGB. Using thin scintillators also has the advantage of limiting the light loss due to absorption in the material. Once the light has been produced in the scintillator, it must be detected and converted to an electrical signal for analysis.

### 2.3.2 Photomultiplier Tubes

Photomultiplier tubes (PMT) use the photoelectric effect to convert photons into electrons and then multiply the electrons into larger and larger cascades that can then be detected as an electrical signal (see Fig. 2.14).

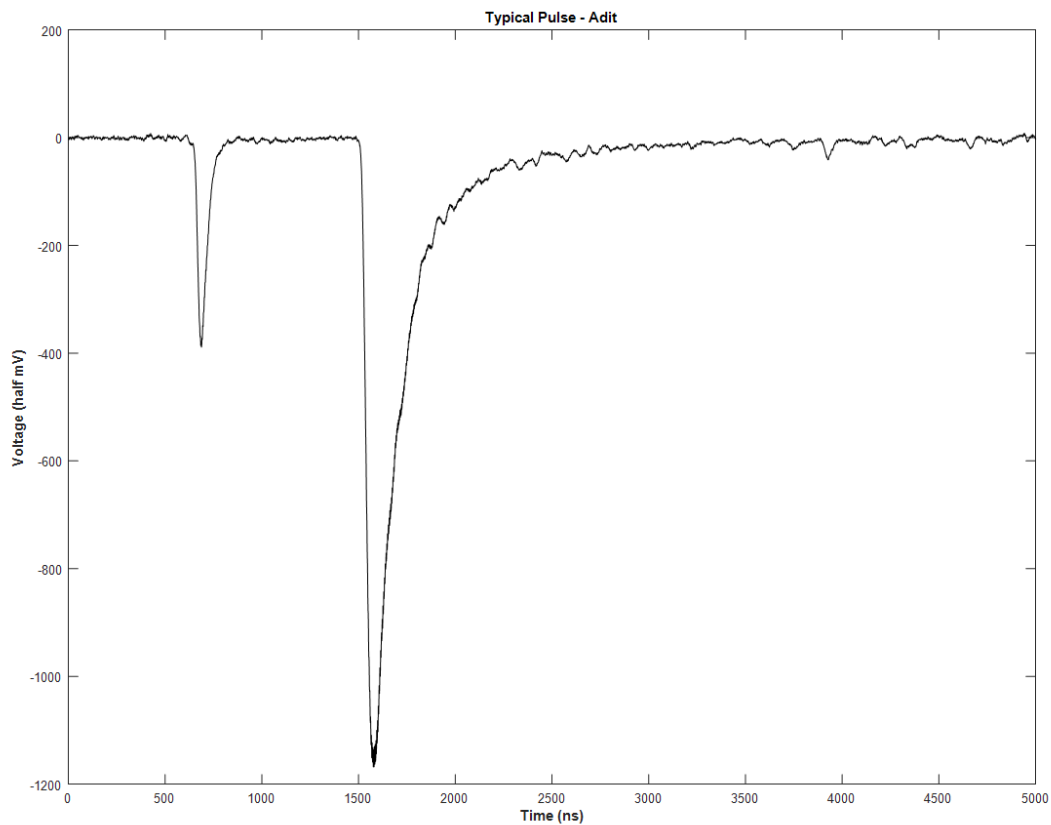


**Figure 2.14** Interactions in a photomultiplier tube (Qwerty123uiop 2013). Image used under the Creative Commons Attribution-Share Alike 3.0 Unported license.

For good photon to electron conversion, the work function of the photoelectric material on the inside face of the PMT must be matched to the output wavelength spectrum of the scintillator used.

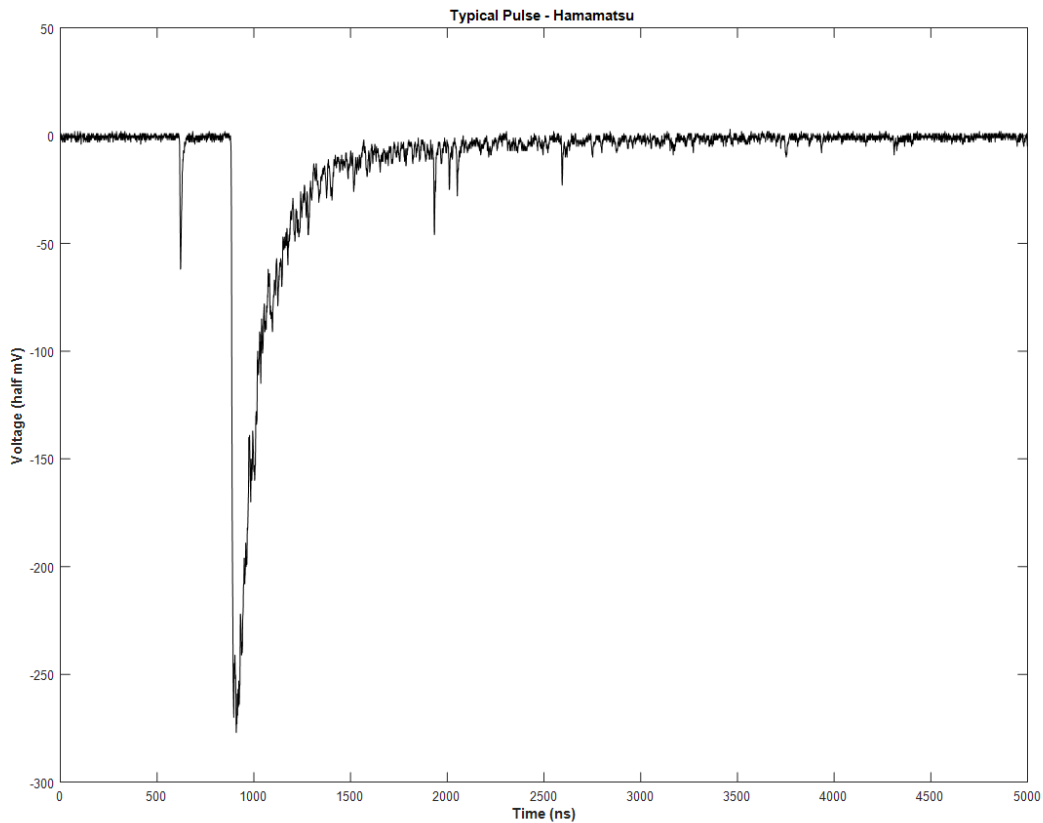
It is important to ensure that power is never applied to the PMT when it is exposed to ambient light as this will cause severe damage. Thus, any detectors using a PMT must be contained in some sort of light tight container. It is useful to do initial tests in a light-tight box, providing easy access for quick changes to the detector. After initial tests are complete it is often more convenient to install the detector into a more portable light-tight can.

We tested two different types of PMTs with our spectrometer. The first was a 10-dynode Adit B133D01 (see B133D01 Photomultiplier Tube in Appendix B). Initial tests showed that this PMT would be good for preserving the pulse shapes resulting from light output from the scintillator. However, additional tests have shown that this is not likely to be true. Additionally, the Adit PMT does not have good timing capabilities (see Fig. 2.15).



**Figure 2.15** A typical waveform from a neutron detection event using an Adit B133D01 PMT and LGB scintillator. The first pulse is the proton recoil pulse and the second, wider pulse is the neutron capture pulse.

The second kind of PMT we used is a Hamamatsu R1250 (see Hamamatsu Photomultiplier Tube R1250 in Appendix B). This 14-dynode PMT has substantially better timing capabilities than the Adit (see Fig. 2.16).

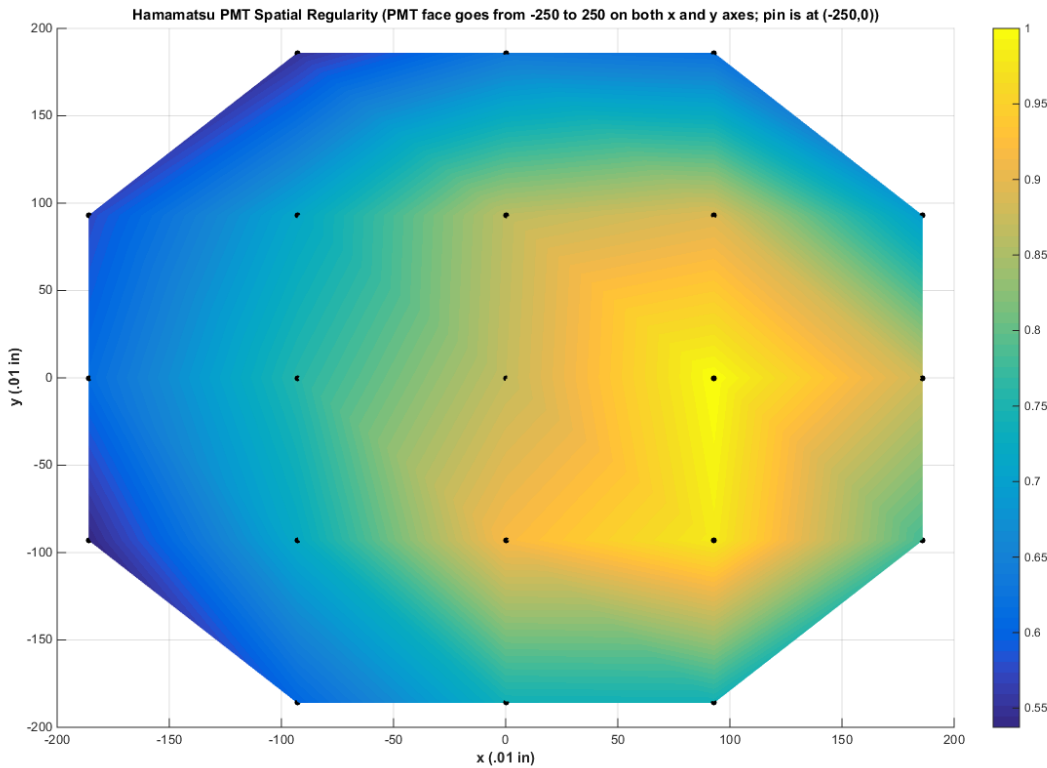


**Figure 2.16** A typical waveform from a neutron detection event using a Hamamatsu R1250 PMT and LGB scintillator. The first pulse is the proton recoil pulse and the second, wider pulse is the neutron capture pulse.

The increased timing resolution in the Hamamatsu tube makes it simpler to separate recoil and capture pulses for small drift times. The significantly narrower recoil pulses also make pulse shape discrimination using pulse area more reliable (this will be discussed in more detail later on). It also reveals more information about the structure of the LGB response to a neutron detection event. The Hamamatsu waveform suggests that some of the light output occurs after the initial peak, rather than decaying from the initial output as the Adit waveform suggests.

Although the Hamamatsu provides better temporal resolution, there is a huge variation (nearly

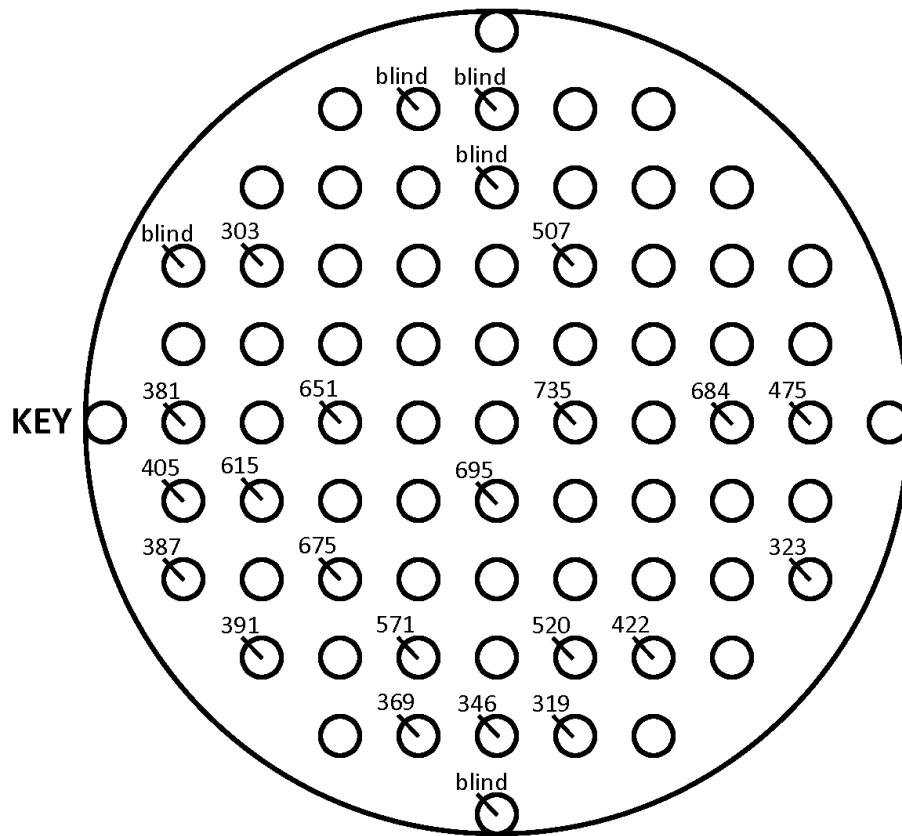
50%!) in signal output across the face of the PMT, making it great for time-of-flight but not good for pulse shape analysis (see Fig. 2.17).



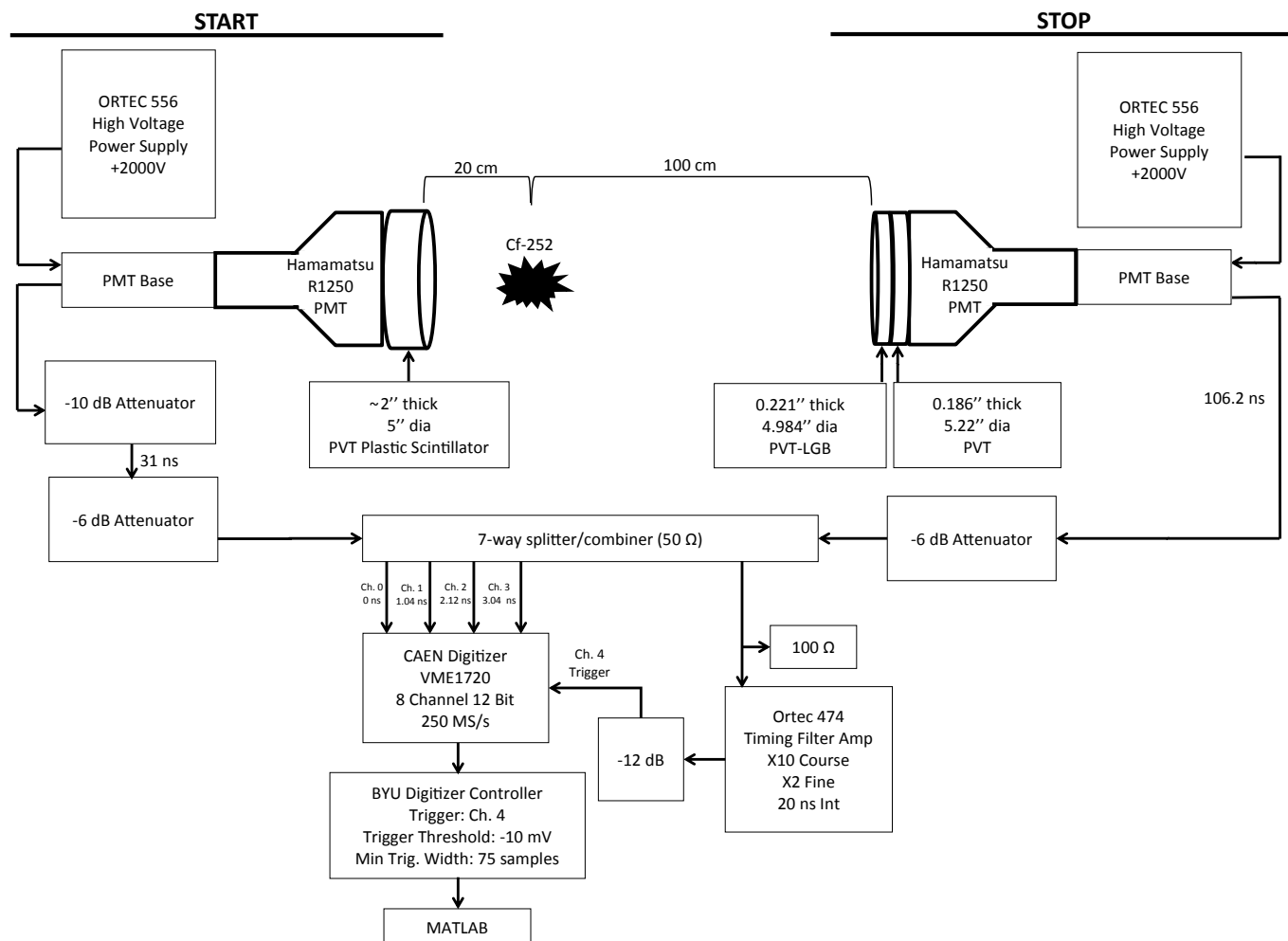
**Figure 2.17** Spatial inconsistency across the face of a Hamamatsu R1250 photomultiplier tube. A small piece of  ${}^6\text{Li}$  glass was placed at each black dot and a pulse height histogram was taken. The color value indicates the relative peak value of each histogram.

We used the Hamamatsu PMT both to get high resolution timing data for comparing our spectrometer to time-of-flight data and also to get high resolution information about the pulse shapes produced by the detector. The final spectrometer was planned to use an Adit PMT to maximize energy information preservation since we would no longer need good timing information for comparison to time-of-flight. Though the slower PMT makes it more difficult to distinguish between occasionally very close start-recoil and recoil-capture pulse pairs, we initially thought the gain in

pulse area consistency across the PMT would be a more significant effect. Additional spatial consistency studies of the Adit PMT indicate that it is even worse than a Hamamatsu for both spatial and timing consistency. 25 randomly selected locations on the face of the PMT were tested using a very small piece of  $^6\text{Li}$  glass. The variation across the tube is at least 60% (see Fig. 2.18).



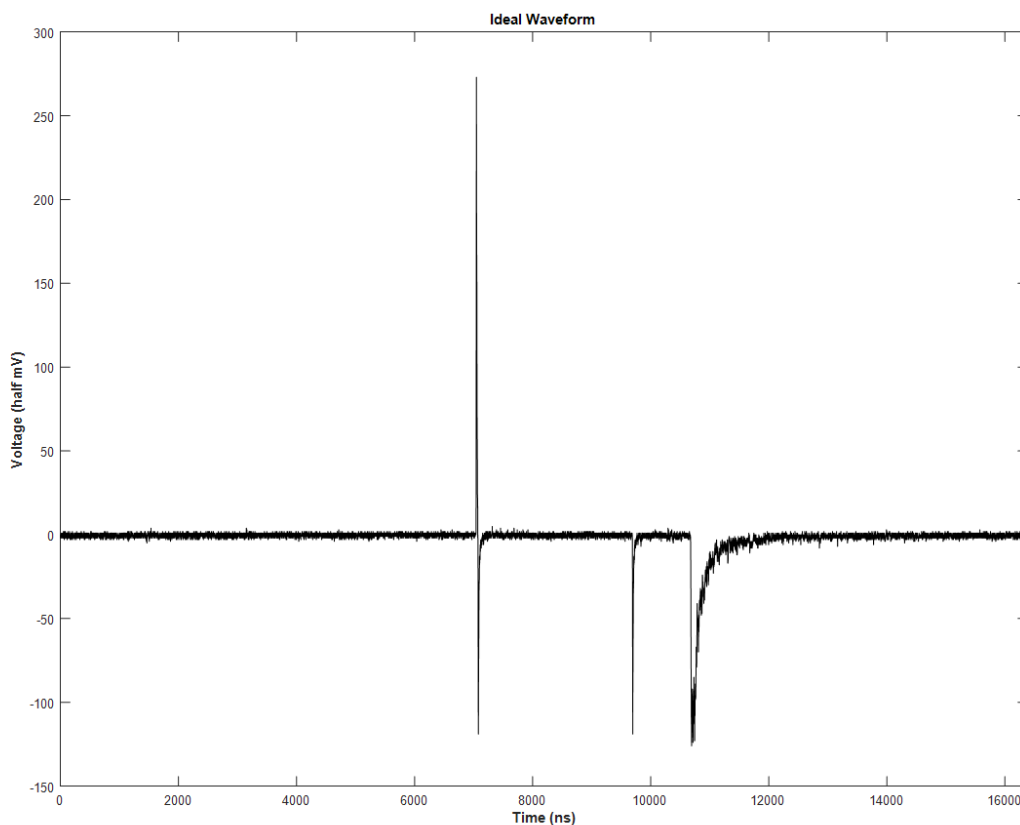
**Figure 2.18** Spatial inconsistency looking down at the face of an Adit B133D01 photomultiplier tube. A small piece of  $^6\text{Li}$  glass was placed in each labeled circle (chosen randomly) and a pulse height histogram was taken. The label indicates the peak value of each histogram. The locations marked "blind" did not produce any triggers at all.



**Figure 2.19** The developmental setup used to test and characterize the spectrometer. The detector is installed in a time-of-flight setup.

### 2.3.3 Developmental Setup

In order to determine the spectrometric capabilities of our detector we installed it as the stop detector in a time-of-flight setup (see Fig. 2.19). The ideal waveform produced by a neutron detection event from this setup would contain four (and only four) distinct pulses (see Fig. 2.20).



**Figure 2.20** An ideal neutron detection event from the developmental setup.

The first pulse (positive) results from a gamma detection event in the start detector. The second pulse (negative) results from a gamma detection event in the stop detector. This would consistently occur within a narrow window of time since gammas all travel at the speed of light. The third pulse (negative) is from a proton recoil detection event and the fourth pulse (negative) from a neutron capture event. In practice, it is very rare that a non-coincidental gamma is detected in the stop detector since the solid angle is so low a meter from the source. Additionally, there are almost always other coincidental pulses scattered throughout the event, complicating the analysis process.

This section will address many of the considerations necessary for getting good time-of-flight data. First, in order for time-of-flight to be accurately measured it is critical to know the timing



characteristics of the system. Electrical signals do not travel instantly from point to point through cables and other components. A good estimate for choosing cable lengths is around 1 ns per foot of RG-58 coaxial cable. The lengths can then be measured by splitting the signal from a pulse generator along two path lengths and measuring the difference in time between the arrival of the signal along the two paths using an oscilloscope. It is useful to start with two paths of equal length and then add the cable to be measured to one path.

It is not necessarily the best choice for the cable lengths from the start and stop detectors to be the same. By delaying one or the other (typically the stop path) you can increase separation between closely correlated pulses making it much easier to resolve the separate pulses. For example, if your setup has a path length of 1 meter you will detect a gamma in the start detector and then a gamma in the stop detector 3.3 ns later, effectively giving you an idea of where 0 is in your time-of-flight plots. If your cable lengths from each detector are the same and you are using a digitizer with a rate of 250 million samples per second (which we do) it will be impossible to resolve the two separate pulses. If the path lengths are different, two separate pulses will be resolved and the timing difference between them can then be accounted for in analysis.

Once all the cable lengths have been measured and the timing characteristics of the setup are known, it is useful to measure the overall timing difference between the two paths, including all cables and electronics. This is done by placing the detectors face to face with only enough space between them to fit a gamma emitter, such as  $^{60}\text{Co}$ . The time-of-flight between the two detectors should now be zero, so any measured variation from zero is a characteristic of the setup and needs to be accounted for in analysis.

In addition to timing considerations, it is critical that signal, cable, and component electrical impedances are matched everywhere in the system. Most of the cabling and nuclear instrumentation modules (NIM) we use have an impedance of 50  $\Omega$ . If any components are added to the system with a different impedance, it will cause characteristic echos in the signal as a result of re-

flection off the impedance mismatch. It will also attenuate the original signal, damaging the pulse shape information critical to the use of the detector as a spectrometer. One common component in our time-of-flight setup that does not have 50  $\Omega$  input impedance is the Ortec 474 timing filter and amplifier. Its 100  $\Omega$  impedance requires a 100  $\Omega$  terminator in parallel with the input to match the signal impedance. When matching impedance it may be necessary to review resistor summing rules:

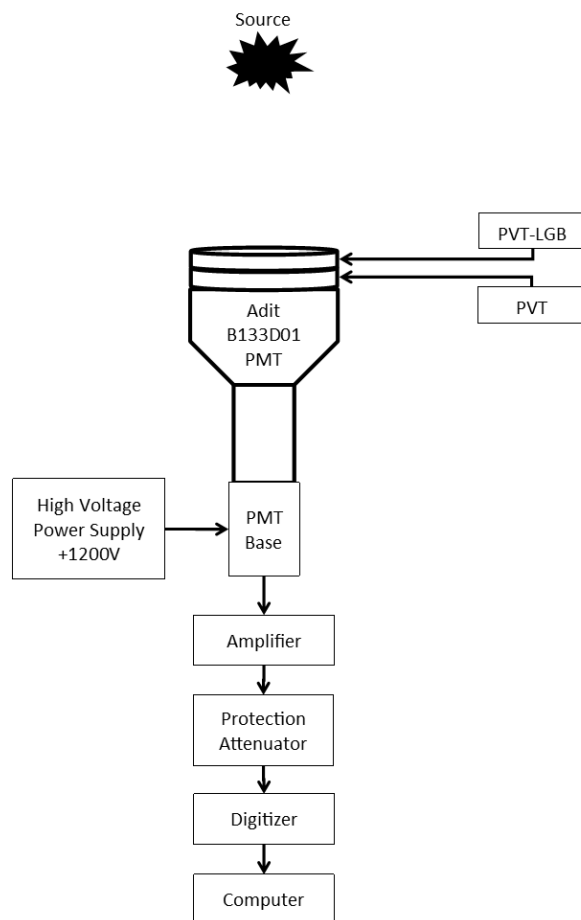
$$\begin{aligned} \text{series: } R &= \sum R_i \\ \text{parallel: } \frac{1}{R} &= \sum \frac{1}{R_i} \end{aligned}$$

The developmental setup uses a CAEN V1720 digitizer with 12 bit voltage resolution over  $\pm 1$  volt and 250 million samples per second per channel (one sample every 4 ns). Having more channels available than needed, we split the signal into four channels, each delayed one nanosecond from the previous. This effectively quadrupled our sample rate to one sample every 1 ns, greatly improving the timing resolution. Rather than do this separately for the start and stop detectors, we summed the two with different polarities; the start waveform with a positive polarity and the stop waveform with a negative polarity (see Fig. 2.20). This provides better resolution for time-of-flight measurements and greater information about the pulse structure from the detector. In combining, splitting, and delaying the two detector signals it was necessary to match impedances at every step along the way. This was done using a 50  $\Omega$  impedance-matched star-configuration 7-way splitter/combiner (one input for each detector, four outputs to the digitizer, and one output to the timing filter to provide an integrated trigger for the digitizer). This trick gives us a voltage resolution of 0.5 mV and a timing resolution of 1 ns.

Another important consideration when building the developmental setup is the voltage requirements and limitations of each component in the system. The digitizer and the timing filter amplifiers have an input limit of only  $\pm 1$  V while the PMTs output up to 30 V from a cosmic ray detection event. Attenuators are necessary to protect the equipment. On the other hand, it is best

to have as much signal over background as possible so we kept the signal voltages near the upper limits of each component in the setup.

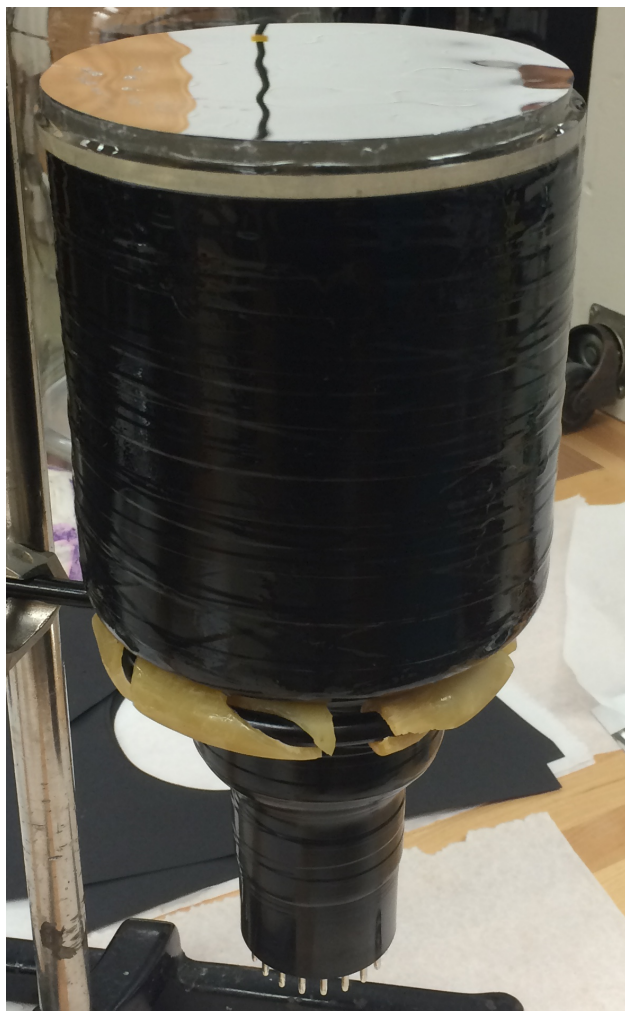
The developmental setup differs from the final stand-alone operation of the spectrometer in several key ways. First, all of the timing considerations are no longer important since the detector does not need to be in a time-of-flight configuration. This greatly simplifies the setup. Additionally, it is not necessary to have such high timing resolution, eliminating the need for a the splitter/combiner and all of the delay cabling. The developmental setup uses a Hamamatsu PMT to increase the resolution of the time-of-flight data. This produces capture pulses that contain many peaks over a long period of time rather than a single wide capture pulse. This pulse shape (see Fig. 2.16) is difficult to trigger on, so we used an integrated waveform (see Fig. 2.19) to trigger the digitizer. The final setup will use a slower Adit PMT, eliminating the need for the integration feature of the Ortec 474 timing filter amp for the integrated trigger. However, the 10-dynode Adit PMT has significantly less gain than the 14-dynode Hamamatsu, so the amplification feature of the 474 may be necessary. Other components necessary for operation of the spectrometer in its intended standalone environment include a power supply for the PMT and a digitizer (see Fig. 2.21).



**Figure 2.21** The necessary components for the spectrometer to operate in its final intended standalone mode.

### 2.3.4 Detector Characterization

The spectrometer consists of three main components: a first scintillating slab containing inorganic LGB crystals in organic polyvinyl toluene (PVT), a second scintillating slab of PVT, and a photo-multiplier tube (see Fig. 2.22).

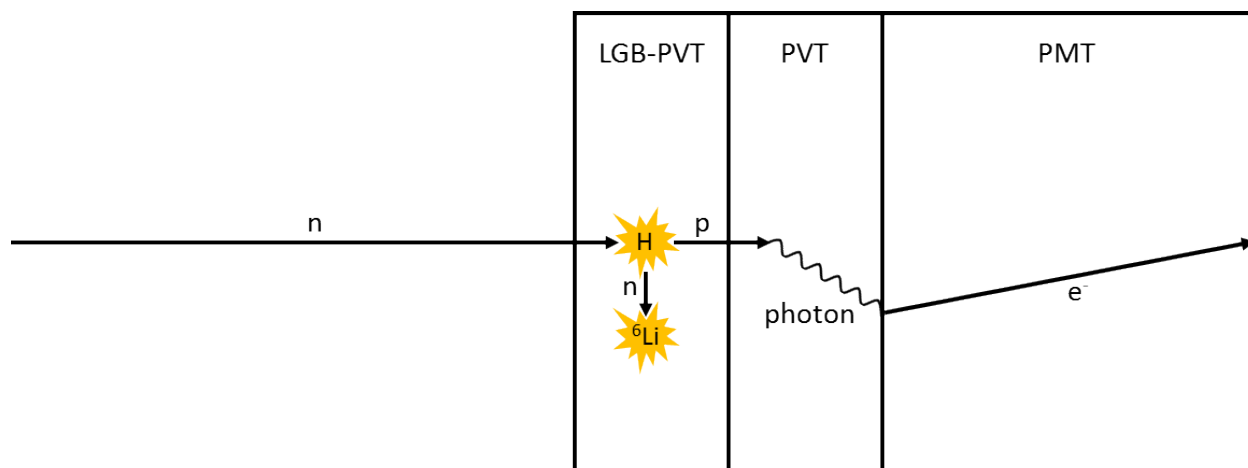


**Figure 2.22** Our neutron spectrometer. The reflective layer on top is aluminized mylar. The first thin slab beneath the mylar is LGB in PVT scintillator. Beneath the LGB is PVT scintillator. Beneath the PVT scintillator is a Hamamatsu R1250 photomultiplier tube.

The LGB slab measures 0.221 inches thick with a diameter of 4.984 inches. The PVT slab measures 0.186 inches thick with a diameter of 5.22 inches. The scintillators are optically coupled to each other and to the face of the PMT with BC-630 silicone grease (see BC-630 Silicone Optical Grease in Appendix B). When run in a time-of-flight setup, the PMT used is a Hamamatsu R1250, chosen for its excellent timing characteristics. When in standalone mode, the detector was planned to use an Adit B133D01 PMT, chosen for its supposedly better energy information preservation.

Additional study has shown that an entirely different photon detector may be necessary. This is discussed in more detail later on.

The detector works in a manner similar to a proton recoil telescope. A neutron collides with a proton in the thin hydrogenous first slab of scintillator, transferring its energy. If all the energy is transferred the neutron will drift ninety degrees to its original path until it either captures in the LGB crystal or exits the detector. The detector is gated on the capture pulse, enforcing neutron detection. If the neutron does not transfer all of its energy to the first collision proton it will move off at an angle of less than ninety degrees and will quickly leave the crystal-filled detection plane. Thus, triggering on captures enforces not only neutron detection events, but also one to one energy transfers. The recoiled proton subsequently excites the scintillator in the second slab of PVT emitting photons that are detected by the PMT (see Fig. 2.23).



**Figure 2.23** A typical neutron detection interaction in the spectrometer. Note that the diagram is not to scale. The scintillators are actually much thinner so that the neutron will escape the detector if the collision does not result in a one-to-one energy transfer.

The number of neutrons successfully detected compared to the total number of neutrons that pass through the detector undetected is referred to as the efficiency of the detector. For example, if 100 neutrons pass through the detector and seven of them are detected, the detector is said to have an efficiency of 7%. In order to calculate efficiency, it is necessary to know how many neutrons

pass through the detector over a given period of time. This can be calculated from the rate of the neutron source and the solid angle of the detector. After the efficiency of the detector has been calculated, it is possible to go in the other direction and calculate unknown source rates from the number of neutrons detected in a given period of time. We tested our detector using a sealed  $^{252}\text{Cf}$  source.  $^{252}\text{Cf}$  undergoes radioactive decay via alpha decay 96.9% of the time:



and by spontaneous fission the other 3.09%, producing various fission fragments and neutrons. Our source had a rate of 88.96  $\mu\text{Ci}$  on 1 August 2012. To calculate the current rate you use

$$A = A_0 2^{-t/T_{1/2}}$$

where  $A$  is the current rate,  $A_0$  is the original rate,  $t$  is the time since the original rate was measured, and  $T_{1/2}$  is the half-life of the material. The half-life of  $^{252}\text{Cf}$  is 2.645 years, giving a current rate of 34.28  $\mu\text{Ci}$  at the time of calculation (with such a short half-life this rate will not stay meaningful for long). A more convenient unit for the next few calculations is the becquerel (Bq), which is one decay per second. 1 curie is  $3.7 \times 10^{10}$  becquerels giving us a current source rate of  $1.27 \times 10^6$  Bq. Since only 3.09% of decays produce neutrons via spontaneous fission and each fission produces on average 3.7675 neutrons that gives us a neutron rate of approximately 147,000 neutrons per second. These neutrons are emitted isotropically (equally in every direction) from the source. The solid angle of the detector is approximately the ratio of the surface area of the face of the detector to the surface area of a sphere with a radius of the distance from the source to the detector (one meter in our setup). This ratio more closely approximates the actual solid angle with increasing distances from the source. Multiplying the number of neutrons emitted from the source each second by this ratio yields the number of neutrons passing through the detector face each second, 148.6 neutrons per second. Over a ten day period, the rate of the spectrometer (triggered on neutron capture pulses) was 1.3 counts per second, yielding an efficiency of 0.87%. Alternatively, assuming there

is a linear relationship between detector efficiency and scintillator thickness, the efficiency of the spectrometer could be calculated from the known efficiency of a similar detector. In our case, one made with four inches of the same type and diameter of scintillator made by Photogenics. Using this method, the efficiency of our spectrometer ought to be around 0.6%. This estimate is in relative agreement with the previous calculation.

## 2.4 Data Acquisition

Data was acquired using a CAEN V1720 digitizer (see V1720 / Digitizers | CAEN in Appendix B). The digitizer was controlled by a C program lovingly referred to as Ugly Controller adapted from digitizer controller software provided by CAEN called CAEN WaveDump. The primary adaptation enabled the digitizer to trigger on pulse width in addition to the standard threshold trigger. Each channel used for data acquisition needs to be DC offset so that the noise level is centered around zero volts in order for predictable trigger behavior and maximum use of the digitizer's  $2 V_{pp}$  range. Each event consisted of 4096 samples (16384 ns) per channel. We used five channels of the available eight: four for signal input and one for the integrated trigger input. The trigger was set to a threshold of 10 mV below the center of the noise level and a width of 75 channels (300 ns). The data was transmitted to the computer via optical link. See Appendix C for a typical configuration file used to configure the digitizer for data acquisition.

## 2.5 Data Analysis

The data was analyzed using software called ToFSpec developed in Matlab. The software reads the binary data saved by the digitizer and processes it in several steps, culminating in several plots of interest. The first step iterates through the data and removes events that do not contain any pulses from the start detector. The second step combines the four sequentially delayed waveforms



from each channel into a single waveform. The third step looks through each waveform and picks out each pulse in the waveform, assigning them labels according to which detector they were produced by. Finally, the fourth step analyzes each individual pulse, measuring various pulse metrics such as height, width, and area. These features are discussed in greater detail in the next section. After analyzing the individual pulses, the software builds several plots displaying the various features compared with each other (for pulse shape discrimination), time-of-flight, and energy to look for useful correlations. The analysis process also aids in calibrating the detector for stand-alone spectrometric measurement.

### 2.5.1 Pulse Metrics

The software characterizes each pulse by recording several pulse metrics: start sample number, peak height, area, early area, time-of-flight (as applicable), and energy (calculated from time-of-flight).

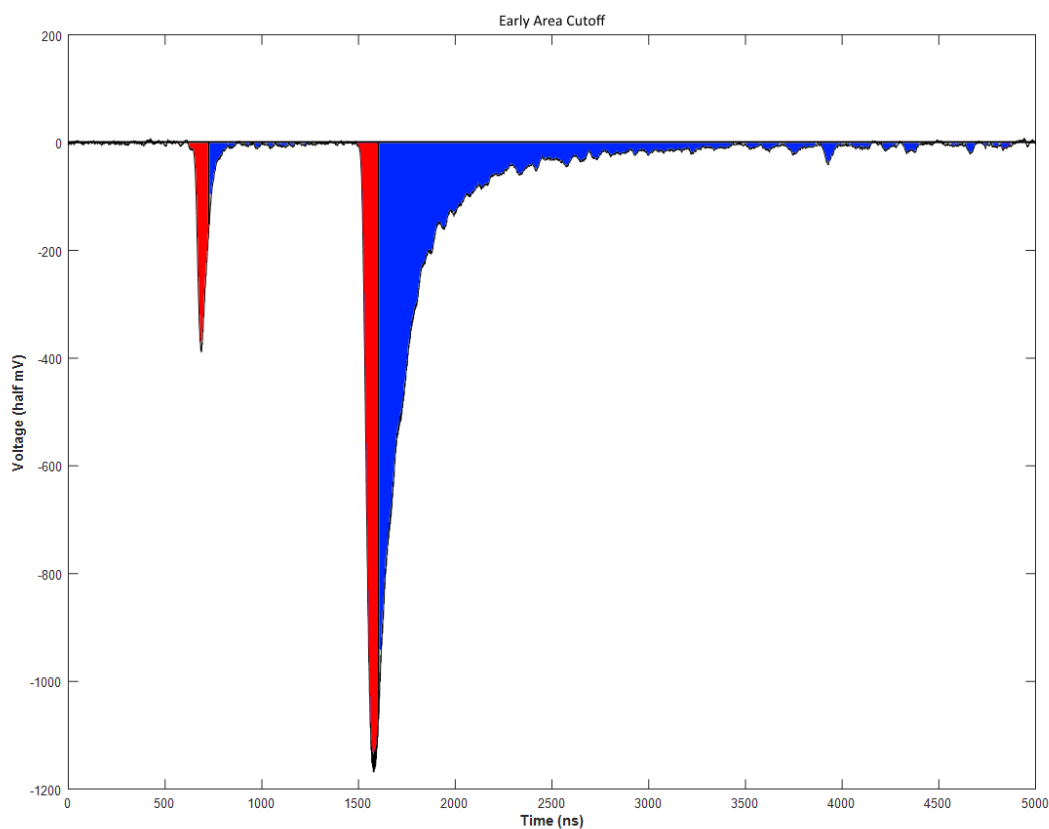
The start sample number is used in time-of-flight calculations and for enforcing timing windows between start and recoil pulses and also between recoil and capture pulses. The windows were chosen to exclude a majority of coincidental events that made it through the event processing stages. For example, neutrons below approximately 0.1 MeV do not have sufficient energy to produce a proton recoil pulse. Therefore, if we see a proton recoil pulse with a time-of-flight longer than what a 0.1 MeV neutron would have, it is likely a coincidental gamma event misidentified as a proton recoil event.

The peak height is related to the energy of the recoil proton that caused the light output in the scintillator by

$$H = kE^{3/2}$$

where  $H$  is the pulse height,  $k$  is a proportionality constant determined by the scintillating material, and  $E$  is the energy of the recoil proton (Knoll 2010). It is assumed that the relationship is the

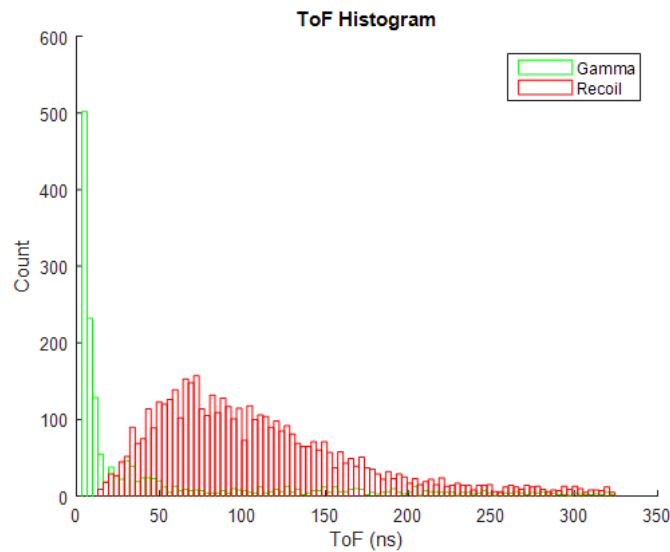
same for pulse area. Pulse area seems to most accurately preserve the energy information among all the pulse metrics. We use the pulse area values to construct an energy spectrum of the original neutrons. The early area is defined as the area of the pulse that occurs before some sample in the pulse and is used for pulse shape discrimination (see Fig. 2.24).



**Figure 2.24** A waveform with an arbitrary early area cutoff marked. The area of the pulse in red would be considered early. The area in red and blue would be considered the total area.

For the data collected with a Hamamatsu PMT, the early area cutoff was set at 15 samples (15 ns) after the start of the pulse. For an Adit PMT, the cutoff should be chosen further into the pulse. The exact number can be chosen by inspecting the pulses in the data.

Time-of-flight data is calculated for all pulses except the first pulse in the waveform. Ideally, this would be actual time-of-flight (start gamma to stop recoil) or drift time (stop recoil to stop capture). However, it could also be the time between coincidental pulses, which is meaningless. Although this would be among the first places to improve the analysis process, time-of-flight histograms indicate that this method works sufficiently well (see Fig. 2.25).

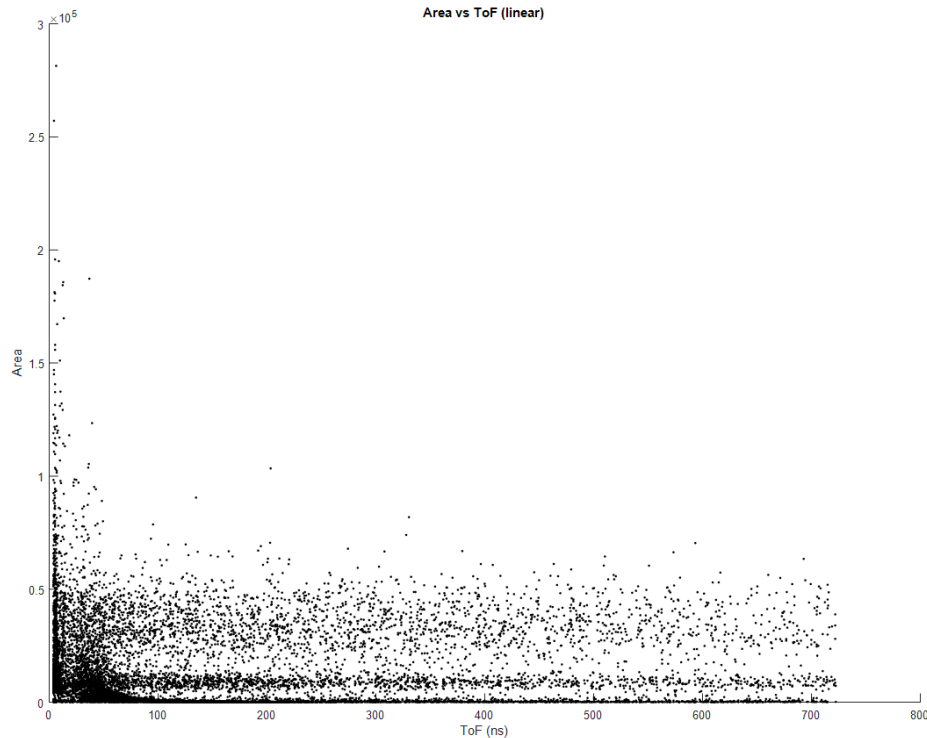


**Figure 2.25** A histogram of time-of-flight values of all pulses identified as either gammas or recoils.

The time-of-flight information is used to do some initial pulse type identification. Gammas only take 3.3 ns to travel 1 m so any pulse with a time-of-flight in this region is either a gamma or a coincidental. The energy is then calculated from the time-of-flight. Time-of-flight and energy are each plotted against pulse area in an effort to find the proportionality constant for our scintillator. This calibrates the spectrometer for taking standalone energy measurements.

## 2.5.2 Pulse Shape Discrimination

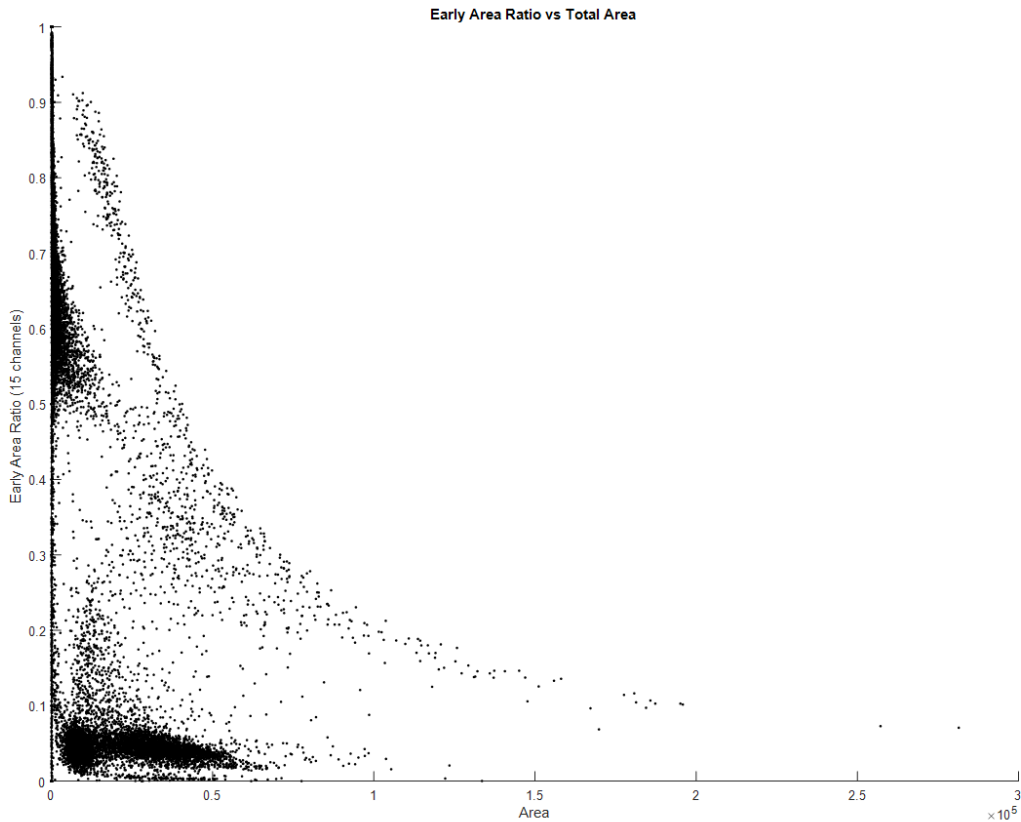
In order to determine the proportionality constant  $k$  it is necessary to find the relationship between the area of pulses caused by recoil events and time-of-flight (see Fig. 2.26).



**Figure 2.26** A plot of the area pulse metric versus time-of-flight. This plot includes all pulses from the stop detector except the first pulse from each event (since the first pulse in each event does not have a time-of-flight value).

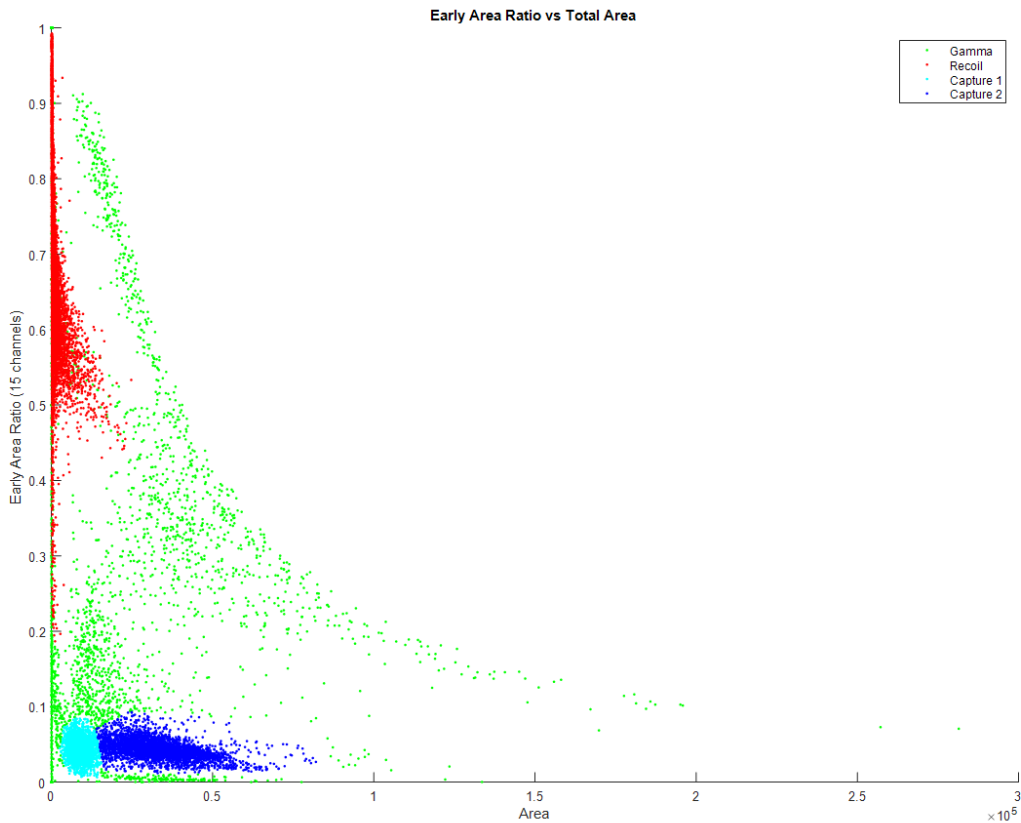
Though some regions seem to appear in the area versus time-of-flight plot, it is difficult to recognize any useful correlations. This is because the plot contains all pulses except the first pulse in each waveform (since the first pulse serves as the start pulse for the first pair of pulses in each waveform). It is necessary to find some way to sort out the pulses caused by recoil events for analysis. This is done using the ratio of early area to total area compared to the total area of each

pulse (see Fig. 2.27).



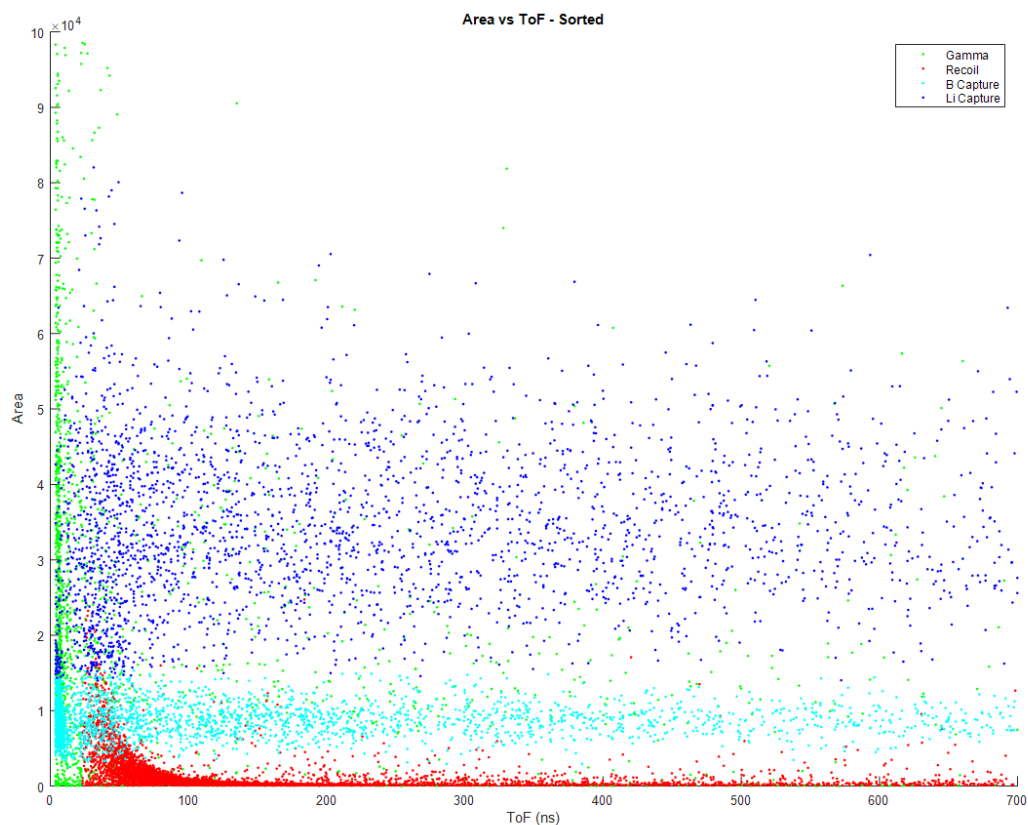
**Figure 2.27** A plot of the ratio of early area to total area versus total area. The regions in this plot are useful for pulse shape discrimination.

The regions in this plot are much better defined and depend on the shape of the individual pulses. Capture pulses are very wide and have very low early area to total area ratios with a variety of total areas. There are two regions that fit these parameters. It is likely that one region contains capture in  ${}^6\text{Li}$  and the other in  ${}^{10}\text{B}$ . Recoils and gammas are very narrow and have much higher early area to total area ratios. Using these features it is possible to determine which types of interactions collect into each region. (see Fig. 2.28).



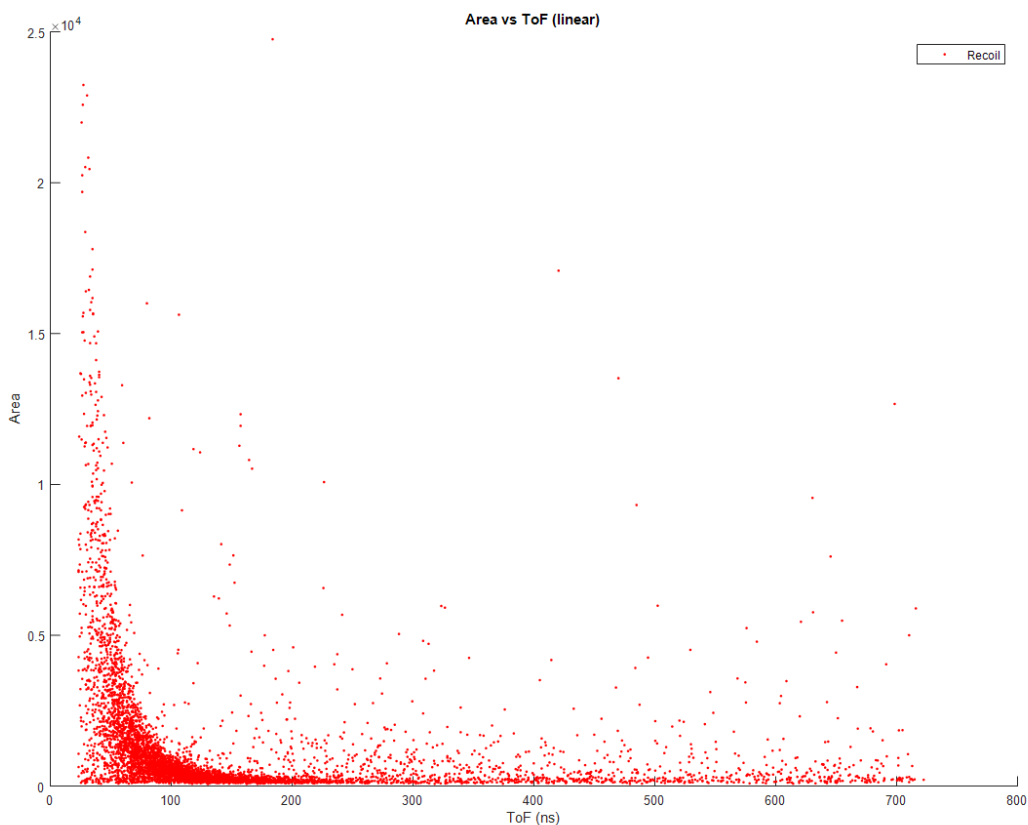
**Figure 2.28** A plot of the ratio of early area to total area versus total area with regions selected. Events marked in red are labeled recoil events, blue are capture in  ${}^6\text{Li}$ , and cyan are capture in  ${}^{10}\text{B}$  (determined by energy considerations). The green is mostly gammas, but also includes other unidentified events such as muon interactions.

It is useful to plot each of the pulse metrics by themselves (in a histogram) after sorting in order to ensure the selected regions make physical sense. Once the pulses have been sorted using pulse shape discrimination, it is possible to sort the area versus time-of-flight plots by pulse type (see Fig. 2.29).



**Figure 2.29** A plot of the area pulse metric versus time-of-flight sorted based on the pulse shape discrimination from the early area ratio plot regions.

Now the originally obscured features of Figure 2.26 become much more apparent. The pulses identified as gammas (green) are clustered together with very low time-of-flight with coincidental gammas scattered across the x-axis. The two capture regions (blue,  ${}^6\text{Li}$ ; and cyan,  ${}^{10}\text{B}$ ) with fairly narrow area distributions along the y-axis are spread across a wide range of time-of-flight, as expected. Finally, the recoils (red) begin to appear under all the other data. They can be easily extracted (see Fig. 2.30).



**Figure 2.30** A plot of the area pulse metric versus time-of-flight containing only pulses labeled as recoil events based on the pulse shape discrimination from the early area ratio plot regions.

Here it becomes very obvious that there is a correlation between recoil pulse area and time-of-flight. Ideally, there should be only a thin, strong line following the top of the curve. We instead see a curve with a large amount of "rain" beneath it and a few scattered pulses above it. The pulses above the curve are likely explained by room-return neutrons.

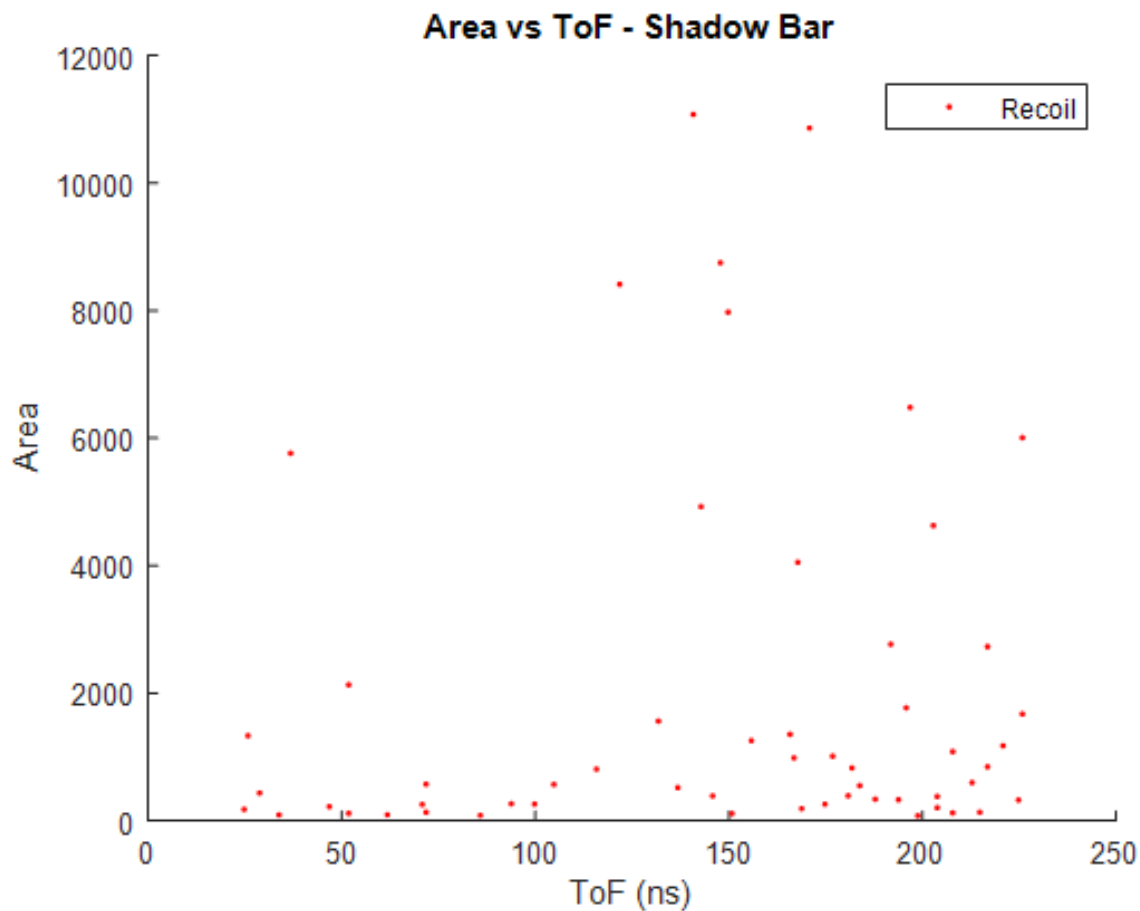
We believe the rain to be a result of the non-linear response of the plastic scintillator at low energies and the spatial inconsistency of the Hamamatsu PMT. For a given energy there would be a distribution of pulse areas up to a maximum value depending on where the photon entered the



face of the PMT. Several tests involving masking off parts of the PMT face did not seem to improve the rain effect at all. It may also be caused by edge effects if the neutrons are being detected around the outside edge of the PMT. We have not yet found a solution to this problem.

### 2.5.3 Room Return

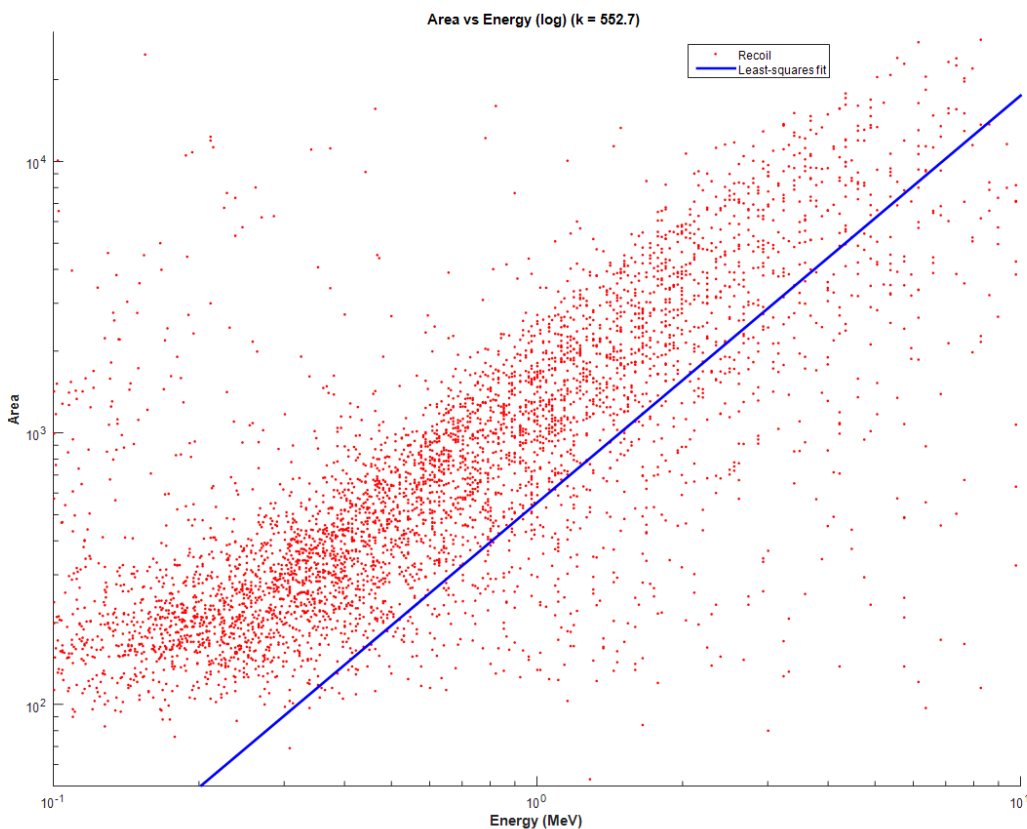
We also made attempts to verify whether the random events were in fact caused by room return. Any neutrons that are not captured, either in a detector or otherwise, will bounce in a random direction, sometimes toward the detector. A neutron that is detected after traveling an indirect path to the detector will have a long time-of-flight that depends on the path it took, rather than its original energy. Additionally, some of its energy is deposited in whatever particle it recoiled off of so that the pulse area when it is finally detected is no longer correlated with its original energy. In order to test whether the randomly distributed points on the plot are caused by room return we took some data with a shadowbar between the source and the detector. The shadow bar is made of a hydrogenous moderator so that any neutrons traveling a direct path to the detector are absorbed before they reach the detector. Any events detected can be attributed to room return. (Note that the shadow bar itself would increase the rate of room return, but not by much.) The shadow bar data supports the explanation of the data above the curve (see Fig. 2.31).



**Figure 2.31** A plot of the area pulse metric versus time-of-flight. This data was taken with a shadow bar in place in an effort to measure only room return neutrons.

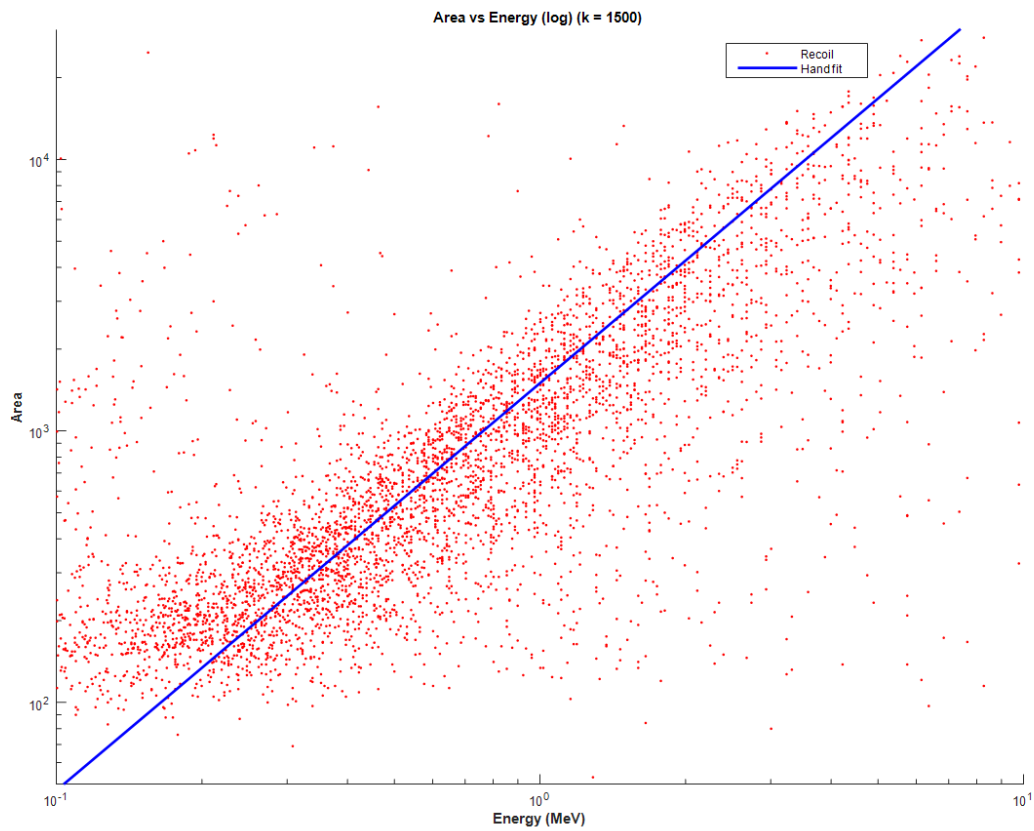
#### 2.5.4 Fitting the Data

With the recoils now identified, it is possible to find the proportionality constants. A least squares regression is used to fit the area to energy relation to the data (see Fig. 2.32).



**Figure 2.32** Log plot of area versus energy of proton recoil events. A least squares regression is used to find an initial value of  $k$ .

This fitting method gives a good initial guess for the value of  $k$  which is then tweaked by hand to more closely approximate the data (the least squares regression is strongly influenced by outliers) (see Fig. 2.33).



**Figure 2.33** Log plot of area versus energy of proton recoil events. The value of  $k$  has been tweaked by hand to better fit the data.

Neutron energies can then be calculated from pulse areas values.

# Chapter 3

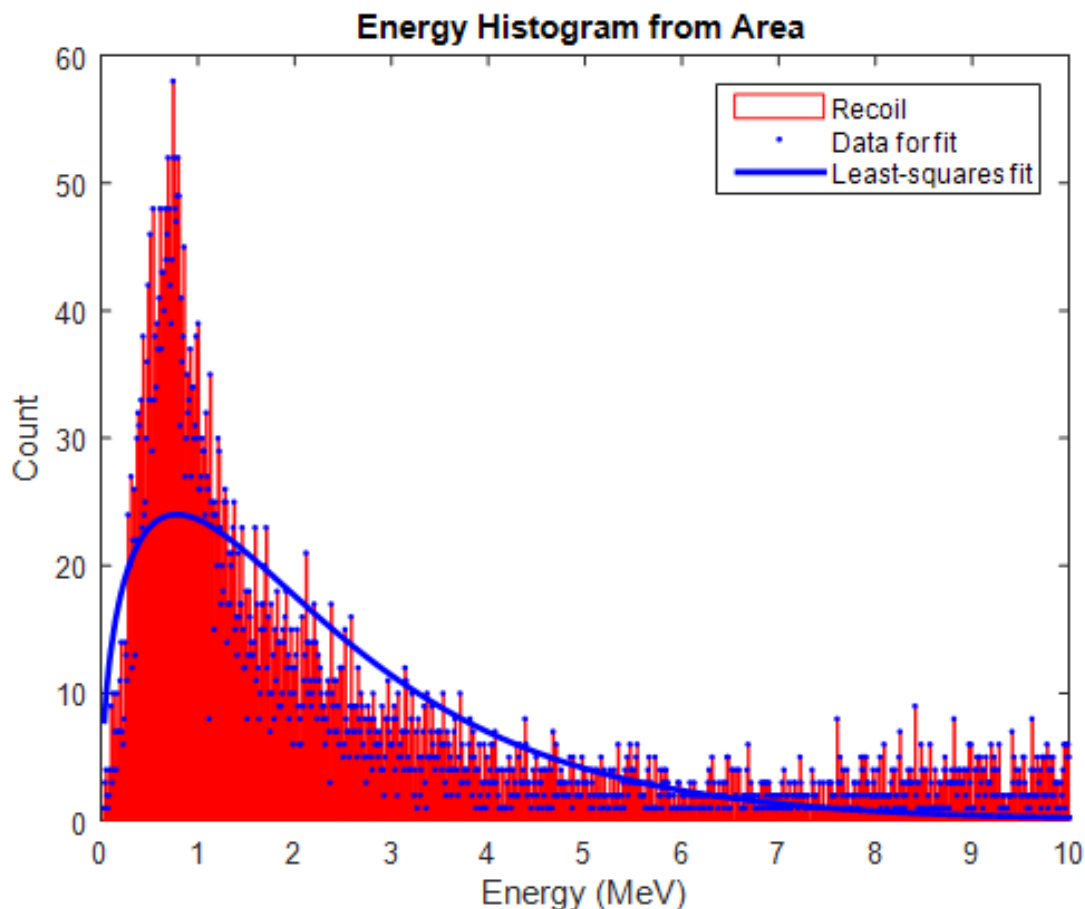
## Results

### 3.1 Energy Spectrum

The neutron energy spectrum calculated from pulse area approximates the spectrum for  $^{252}\text{Cf}$  in the range of around 0.1 MeV to 3 MeV. For comparison, the actual neutron energy spectrum for  $^{252}\text{Cf}$  is very closely approximated by a Watt distribution

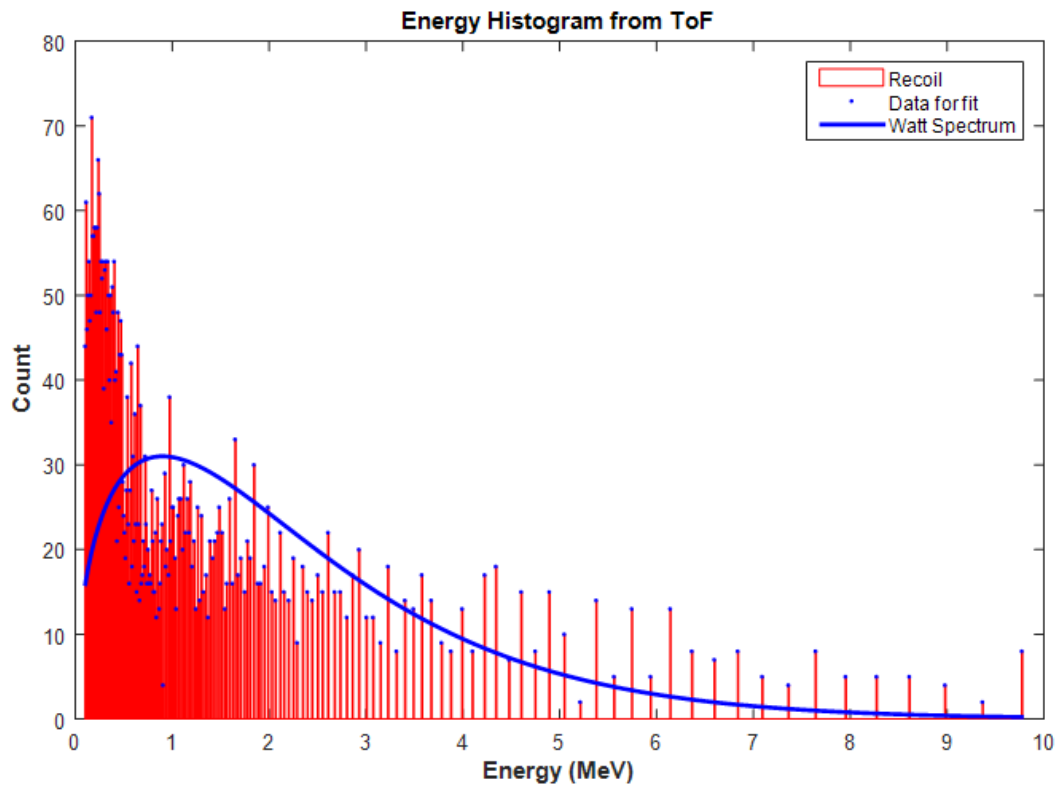
$$f(E) = C \exp(-E/a) \sinh((bE)^{1/2})$$

where  $C$  is a constant of proportionality,  $E$  is neutron energy,  $a$  and  $b$  are constants specified by the isotope. For  $^{252}\text{Cf}$ ,  $a = 1.025$  and  $b = 2.926$  (X-5 Monte Carlo Team 2003) (see Fig. 3.1).



**Figure 3.1** The neutron energy spectrum of  $^{252}\text{Cf}$  calculated from proton recoil pulse areas. The Watt distribution approximating the neutron spectrum has been fit to the histogram using a non-linear least squares fitting method.

The energy distribution measured by the spectrometer has the right general shape and a peak near the most common neutron energy in the energy spectrum, but it does not fit the Watt distribution perfectly, especially for energies higher than 1 or 2 MeV. We believe this is due to limitations in the scintillator and photomultiplier tubes used. Our attempts to explain and correct these effects will be discussed in the next section. Since the detector was installed in a time-of-flight system, we can also compare the energy spectrum calculated from pulse area to the spectrum calculated from time-of-flight (see Fig. 3.2).



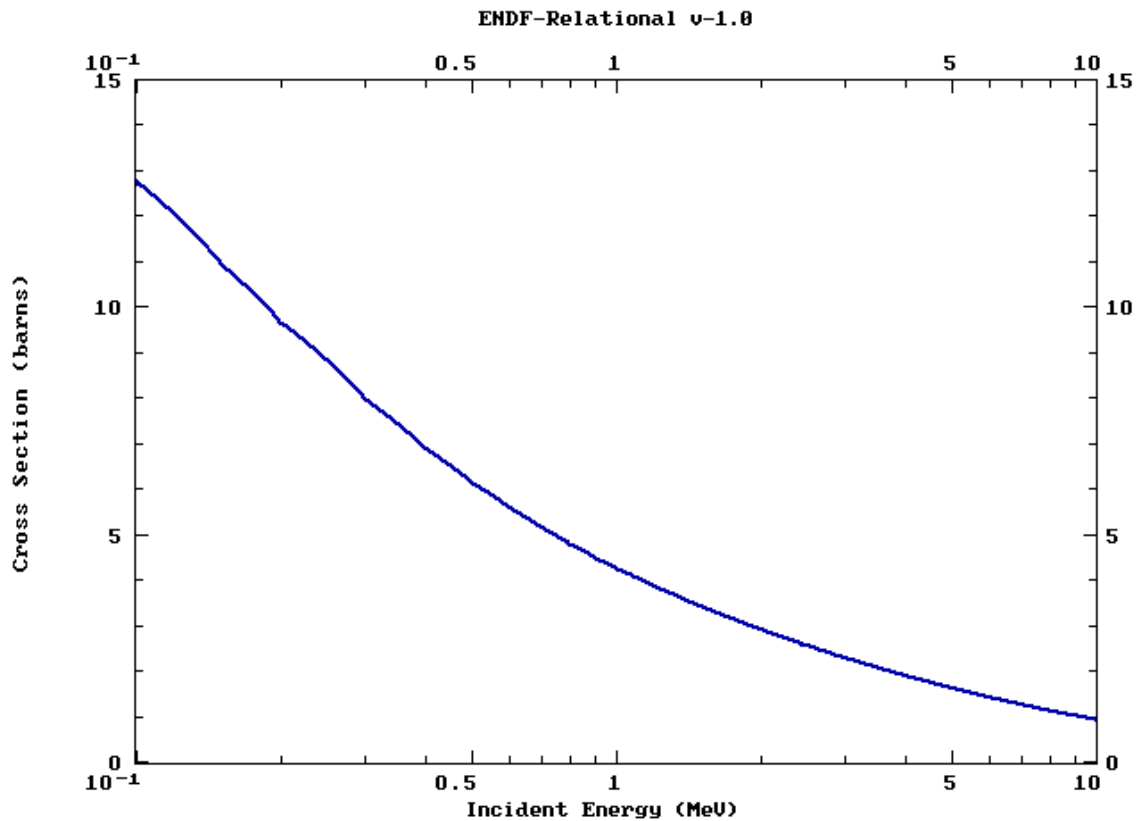
**Figure 3.2** The neutron energy spectrum of  $^{252}\text{Cf}$  calculated from time-of-flight. The Watt distribution approximating the neutron spectrum has been fit to the histogram using a non-linear least squares fitting method.

The energy histogram calculated from area actually seems to fit the predicted spectrum better than the spectrum calculated from time-of-flight. This is almost certainly due to poorly enforced timing windows for the time-of-flight calculations rather than any actual superiority of the spectrometer to time-of-flight data.

### 3.1.1 Corrections to the Energy Spectrum

There are at least two important effects that may explain the discrepancy between the calculated energy spectrum and the Watt distribution. First, the plastic scintillator we used does not have a linear energy to light output relationship. Light output is disproportionately higher for higher en-

ergy protons. However, if the non-linearity can be characterized, a correction can be applied to the energy spectrum. Second, the cross section for elastic scattering in hydrogen changes dramatically over the energy range of interest (see Fig. 3.3).



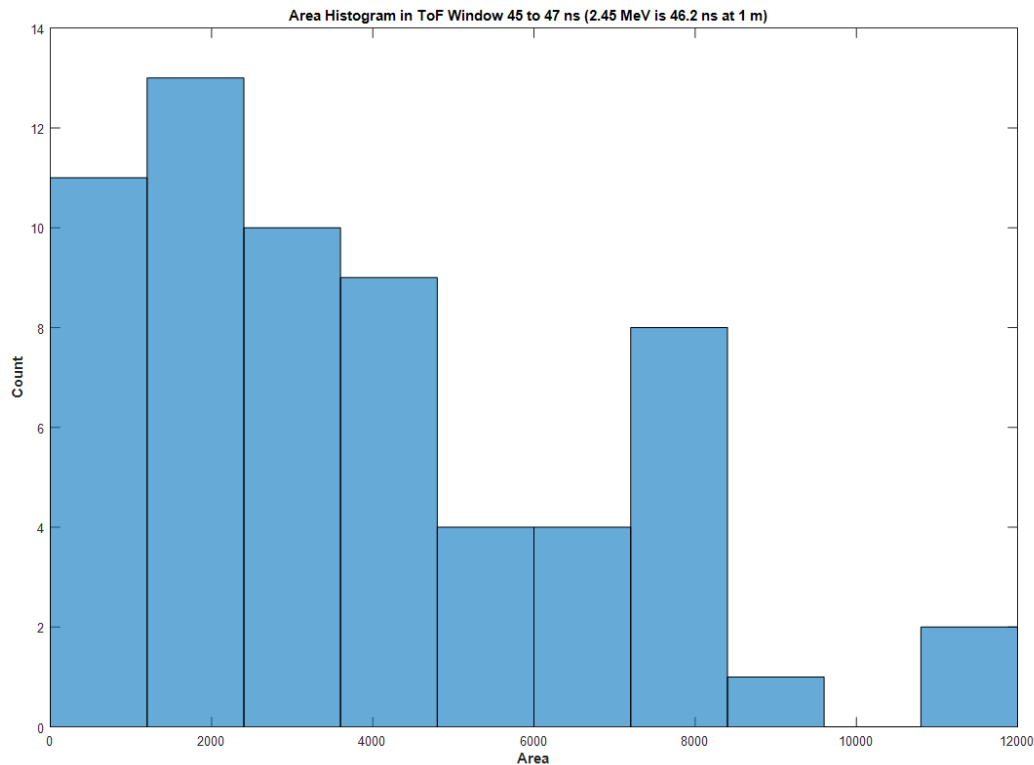
**Figure 3.3** The total neutron cross section of  $^1\text{H}$  versus energy (ENDF 2016).

As before, if the cross section information is known, we can correct the measured energy spectrum accordingly (see the `analyzePulses` function in Appendix D for a rough method of implementing these corrections). These corrections, along with corrections to flatten the PMT response, will hopefully improve the resolution of the spectrometer.



## 3.2 Applications of the Spectrometer

The detector shows some promise as a spectrometer in the energy range of around 0.1 MeV to 3 MeV, but requires additional work to become practically useful. This energy range includes the neutron energy produced by d-d fusion (2.45 MeV) and is thus useful for exploring the increased rates of d-d fusion in condensed matter. However, as stated before, the spread of measured area for a given neutron energy is currently too large for the detector to be a useful spectrometer (see Fig. 3.4).



**Figure 3.4** A histogram of measured pulse area for a narrow time-of-flight window, 45-47 ns. 2.45 MeV corresponds to 46.2 ns at 1 m.

Ideally, the histogram would have a narrow peak near the bin corresponding with the largest pulse area. Other effects, such as inconsistency in the PMT and non-linear light output in the

plastic scintillator spread the peak down and need to be corrected for.

If such corrections can be made, the detector would perform a similar role for neutrons as the NaI detector for gammas. It would be very portable and easy to use at the expense of excellent resolving power. More portable than a time-of-flight system with higher resolution than a Bonner sphere system, this spectrometer would help to improve our neutron spectrometric capabilities.

### 3.3 Future Work

Some additional work is necessary to further improve the resolution and overall utility of the detector. Current photomultiplier tubes do not provide consistent photon to electron conversion across the face of the tube. This effect is significant: as much as a 55% variation for the Hamamatsu R1250 and greater than 60% for the Adit B133D01. If the spatial inconsistency is measured for the tube to be used with the detector, it may be possible to correct the inconsistency with a neutral density filter. It may be necessary to seek an alternative photon detection method.

Another opportunity for future work is improving the pulse shape discrimination capabilities of the detector and the software. Improvements in both scintillator technology and software methods will help to reduce misidentification of gammas as proton recoils and vice versa. The capabilities of this neutron spectrometer will continually improve with future developments in scintillator and photomultiplier tube technology and in software pulse shape discrimination methods.

# Appendix A

## Screening Potential Table

**Table A.1** Screening potentials of d-d fusion in deuterated foils of several metals.

Target	$U_e$ (eV)	Source	Year
SrD <sub>1,0</sub>	350-800	Huke, Phys. Rev. C 78, 015803	2008
Pd	800	Bonomo, Nucl. Phys. A 719, 37	2003
Pd	800	Raoila, Eur. Phys. J. A 19, 283	2004
PtD <sub>x</sub>	730	Bonomo, Nucl. Phys. A 719, 37	2003
Sb	720	Raoila, Eur. Phys. J. A 19, 283	2004
Pt	680	Rolfs, Prog. Theor. Phys. Supplement 154, 373	2004
Pt	675	Raiola, Eur. Phys. J. A 27, 79	2006
Pt	670	Raoila, Eur. Phys. J. A 19, 283	2004
Co	640	Bonomo, Nucl. Phys. A 719, 37	2003
Co	640	Raiola, Eur. Phys. J. A 27, 79	2006
Co	640	Raoila, Eur. Phys. J. A 19, 283	2004
PdO	600	Kasagi, J. Phys. Soc. Jpn. 71, 2881	2002
PdO	600	Kasagi, Surf. Coat. Tech. 201, 8574	2007

Table A.1 cont.

Target	$U_e$ (eV)	Source	Year
Tl	550	Bonomo, Nucl. Phys. A 719, 37	2003
Tl	550	Raoula, Eur. Phys. J. A 19, 283	2004
Li <sub>l</sub>	543	Toriyabe, Phys. Rev. C 85, 054620	2012
Bi	540	Raoula, Eur. Phys. J. A 19, 283	2004
Pt	530	Raiola, Eur. Phys. J. A 27, 79	2006
Pt	530	Raiola, Eur. Phys. J. A 27, 79	2006
Al	520	Bonomo, Nucl. Phys. A 719, 37	2003
Al	520	Raoula, Eur. Phys. J. A 19, 283	2004
In	520	Raoula, Eur. Phys. J. A 19, 283	2004
Ba	490	Raoula, Eur. Phys. J. A 19, 283	2004
Co	480	Raiola, Eur. Phys. J. A 27, 79	2006
Pb	480	Raoula, Eur. Phys. J. A 19, 283	2004
Pt	480	Raiola, Eur. Phys. J. A 27, 79	2006
V	480	Raoula, Eur. Phys. J. A 19, 283	2004
Zn	480	Bonomo, Nucl. Phys. A 719, 37	2003
Zn	480	Raoula, Eur. Phys. J. A 19, 283	2004
Cu	470	Raoula, Eur. Phys. J. A 19, 283	2004
LiF	470	Engstler, Z. Phys. A 342, 471	1992
Nb	470	Raoula, Eur. Phys. J. A 19, 283	2004
Pt	465	Raiola, Eur. Phys. J. A 27, 79	2006
Fe	460	Raoula, Eur. Phys. J. A 19, 283	2004
Fe	450	Bonomo, Nucl. Phys. A 719, 37	2003
Ni	450	Bonomo, Nucl. Phys. A 719, 37	2003

Table A.1 cont.

Target	$U_e$ (eV)	Source	Year
H <sub>2</sub> /D <sub>2g</sub>	440	Engstler, Z. Phys. A 342, 471	1992
Mg	440	Bonomo, Nucl. Phys. A 719, 37	2003
Mg	440	Raouila, Eur. Phys. J. A 19, 283	2004
Pb	440	Bonomo, Nucl. Phys. A 719, 37	2003
Pt	440	Bonomo, Nucl. Phys. A 719, 37	2003
Mo	420	Raouila, Eur. Phys. J. A 19, 283	2004
Re	420	Bonomo, Nucl. Phys. A 719, 37	2003
Nb	400	Bonomo, Nucl. Phys. A 719, 37	2003
Cd	390	Bonomo, Nucl. Phys. A 719, 37	2003
Mn	390	Raouila, Eur. Phys. J. A 19, 283	2004
LiF	380	Engstler, Z. Phys. A 342, 471	1992
Ni	380	Raouila, Eur. Phys. J. A 19, 283	2004
Hf	370	Raiola, Eur. Phys. J. A 27, 79	2006
Cd	360	Raouila, Eur. Phys. J. A 19, 283	2004
Er	360	Raiola, Eur. Phys. J. A 27, 79	2006
Mn	350	Bonomo, Nucl. Phys. A 719, 37	2003
V	350	Bonomo, Nucl. Phys. A 719, 37	2003
Dy	340	Raiola, Eur. Phys. J. A 27, 79	2006
Gd	340	Raiola, Eur. Phys. J. A 27, 79	2006
Ta	340	Bonomo, Nucl. Phys. A 719, 37	2003
TaD <sub>0.13</sub>	340	Bonomo, Nucl. Phys. A 719, 37	2003
Tb	340	Raiola, Eur. Phys. J. A 27, 79	2006
Ag	330	Raouila, Eur. Phys. J. A 19, 283	2004

Table A.1 cont.

Target	$U_e$ (eV)	Source	Year
H <sub>2</sub> /D <sub>2g</sub>	330	Engstler, Z. Phys. A 342, 471	1992
TaD	322	Czerski, Nucl. Instrum. Methods Phys. Res. B 193, 183	2002
TaD <sub>0.9</sub>	322	Huke, Phys. Rev. C 78, 015803	2008
Cr	320	Raiola, Eur. Phys. J. A 19, 283	2004
Sc	320	Raiola, Eur. Phys. J. A 27, 79	2006
Y	320	Bonomo, Nucl. Phys. A 719, 37	2003
ZrD <sub>2</sub>	319	Czerski, J. Phys. G: Nucl. Part. Phys. 35 014012	2008
Sm	314	Raiola, Eur. Phys. J. A 27, 79	2006
Pd <sub>0.3</sub>	313	Huke, Phys. Rev. C 78, 015803	2008
TaD <sub>0.5</sub>	313	Bystritsky, Nuc. Phys. A 889, 93-104	2012
LiF	310	Wang, J. Phys. G: Nucl. Part. Phys. 39, 015201	2012
Pd	310	Kasagi, J. Phys. Soc. Jpn. 71, 2881	2002
Pd	310	Kasagi, Surf. Coat. Tech. 201, 8574	2007
TaD <sub>0.13</sub>	309	Raiola, Eur. Phys. J. A 13, 377	2002
TaD	302	Czerski, Europhys. Lett. 68, 363	2004
H <sub>2</sub> /D <sub>2g</sub>	300	Engstler, Z. Phys. A 342, 471	1992
LiF	300	Engstler, Z. Phys. A 342, 471	1992
ZrD <sub>2</sub>	297	Czerski, Eur. Phys. J. A 27, 83	2006
ZrD <sub>2</sub>	297	Czerski, Nucl. Instrum. Methods Phys. Res. B 193, 183	2002
ZrD <sub>2.1</sub>	297	Huke, Phys. Rev. C 78, 015803	2008
PdD <sub>0.2</sub>	296	Czerski, Europhys. Lett. 68, 363	2004
Ti	295	Raiola, Eur. Phys. J. A 27, 79	2006
TiD <sub>0.23</sub>	295	Raiola, Eur. Phys. J. A 27, 79 (and references therein)	2006

Table A.1 cont.

Target	$U_e$ (eV)	Source	Year
ZrD <sub>2</sub>	295	Czerski, Europhys. Lett. 68, 363	2004
Ti	290	Raiola, Eur. Phys. J. A 27, 79	2006
Au	280	Raouila, Eur. Phys. J. A 19, 283	2004
Ta	270	Raouila, Eur. Phys. J. A 19, 283	2004
TaD <sub>0.13</sub>	270	Raiola, Eur. Phys. J. A 19, 283	2004
Y	270	Raiola, Eur. Phys. J. A 27, 79	2006
Lu	265	Raiola, Eur. Phys. J. A 27, 79	2006
Tm	260	Raiola, Eur. Phys. J. A 27, 79	2006
Ti	250	Raiola, Eur. Phys. J. A 27, 79	2006
TiD <sub>0.26</sub>	250	Raiola, Eur. Phys. J. A 27, 79 (and references therein)	2006
W	250	Raouila, Eur. Phys. J. A 19, 283	2004
La	245	Raiola, Eur. Phys. J. A 27, 79	2006
<sup>6</sup> Li <sub>l</sub>	235	Fang, J. Phys. Soc. Jpn. 80, 084201	2011
Re	230	Raouila, Eur. Phys. J. A 19, 283	2004
Rh	230	Bonomo, Nucl. Phys. A 719, 37	2003
Rh	230	Raouila, Eur. Phys. J. A 19, 283	2004
Cr	220	Bonomo, Nucl. Phys. A 719, 37	2003
Mo	220	Bonomo, Nucl. Phys. A 719, 37	2003
Ru	220	Bonomo, Nucl. Phys. A 719, 37	2003
W	220	Bonomo, Nucl. Phys. A 719, 37	2003
H <sub>2</sub> /D <sub>2g</sub>	218	Wang, J. Phys. G: Nucl. Part. Phys. 39, 015201	2012
Ru	215	Raouila, Eur. Phys. J. A 19, 283	2004
Sr	210	Raouila, Eur. Phys. J. A 19, 283	2004

Table A.1 cont.

Target	$U_e$ (eV)	Source	Year
Zr	205	Raiola, Eur. Phys. J. A 27, 79	2006
ZrD <sub>0.13</sub>	205	Raiola, Eur. Phys. J. A 27, 79	2006
ZrD <sub>2</sub>	205	Bystritsky, Nucl. Phys. A 889 93-104	2012
Ce	200	Raiola, Eur. Phys. J. A 27, 79	2006
Fe	200	Kasagi, J. Phys. Soc. Jpn. 71, 2881	2002
Fe	200	Kasagi, Surf. Coat. Tech. 201, 8574	2007
Ir	200	Bonomo, Nucl. Phys. A 719, 37	2003
Ir	200	Raola, Eur. Phys. J. A 19, 283	2004
Re	200	Kasagi, Surf. Coat. Tech. 201, 8574	2007
Sn	200	Bonomo, Nucl. Phys. A 719, 37	2003
AlD	191	Czerski, Europhys. Lett. 68, 363	2004
AlD	190	Czerski, Nucl. Instrum. Methods Phys. Res. B 193, 183	2002
AlD <sub>0.8</sub>	190	Huke, Phys. Rev. C 78, 015803	2008
Nd	190	Raiola, Eur. Phys. J. A 27, 79	2006
Be	180	Bonomo, Nucl. Phys. A 719, 37	2003
Be	180	Raola, Eur. Phys. J. A 19, 283	2004
Ho	165	Raiola, Eur. Phys. J. A 27, 79	2006
Li	150	Huke, Phys. Rev. C 78, 015803	2008
<sup>7</sup> Li <sub>l</sub>	140	Fang, J. Phys. Soc. Jpn. 80, 084201	2011
TaD	136	Czerski, Eur. Phys. J. A 27, 83	2006
Sn	130	Raola, Eur. Phys. J. A 19, 283	2004
TiD <sub>2</sub>	125	Bystritsky, Nuc. Phys. A 889, 93-104	2012
Cu	120	Kasagi, Surf. Coat. Tech. 201, 8574	2007



Table A.1 cont.

Target	$U_e$ (eV)	Source	Year
Eu	120	Raiola, Eur. Phys. J. A 27, 79	2006
ZrD <sub>2</sub>	112	Czerski, Eur. Phys. J. A 27, 83	2006
Yb	110	Raiola, Eur. Phys. J. A 27, 79	2006
TiD <sub>1.3</sub>	100	Czerski, Eur. Phys. J. A 27, 83	2006
HfD <sub>x</sub>	87	Bonomo, Nucl. Phys. A 719, 37	2003
ZrD <sub>x</sub>	83	Bonomo, Nucl. Phys. A 719, 37	2003
Yb	81	Yuki, J. Phys. G: Nucl. Part. Phys. 23, 1459	1997
Ge	80	Raouila, Eur. Phys. J. A 19, 283	2004
Ni	80	Kasagi, Surf. Coat. Tech. 201, 8574	2007
Yb	80	Kasagi, Surf. Coat. Tech. 201, 8574	2007
PrD <sub>x</sub>	78	Bonomo, Nucl. Phys. A 719, 37	2003
Au	70	Kasagi, J. Phys. Soc. Jpn. 71, 2881	2002
Au	70	Kasagi, Surf. Coat. Tech. 201, 8574	2007
Ho	70	Raouila, Eur. Phys. J. A 19, 283	2004
Pr	70	Raouila, Eur. Phys. J. A 19, 283	2004
Tm	70	Raouila, Eur. Phys. J. A 19, 283	2004
Y	70	Raouila, Eur. Phys. J. A 19, 283	2004
TiD <sub>3.76</sub>	66	Kasagi, J. Phys. Soc. Jpn. 71, 2881	2002
Ti	65	Kasagi, J. Phys. Soc. Jpn. 71, 2881	2002
Ti	65	Kasagi, Surf. Coat. Tech. 201, 8574	2007
Au	61	Bonomo, Nucl. Phys. A 719, 37	2003
C	60	Raouila, Eur. Phys. J. A 19, 283	2004
Ge	60	Bonomo, Nucl. Phys. A 719, 37	2003

Table A.1 cont.

Target	$U_e$ (eV)	Source	Year
La	60	Raoula, Eur. Phys. J. A 19, 283	2004
Si	60	Raoula, Eur. Phys. J. A 19, 283	2004
C	52	Bonomo, Nucl. Phys. A 719, 37	2003
C	50	Raiola, Eur. Phys. J. A 27, 79	2006
CaO <sub>2</sub>	50	Raoula, Eur. Phys. J. A 19, 283	2004
DyD <sub>x</sub>	50	Bonomo, Nucl. Phys. A 719, 37	2003
Er	50	Raoula, Eur. Phys. J. A 19, 283	2004
ErD <sub>x</sub>	50	Bonomo, Nucl. Phys. A 719, 37	2003
Eu	50	Raoula, Eur. Phys. J. A 19, 283	2004
Gd	50	Raoula, Eur. Phys. J. A 19, 283	2004
Ti	50	Raiola, Eur. Phys. J. A 27, 79	2006
TiD <sub>1.1</sub>	50	Raiola, Eur. Phys. J. A 27, 79 (and references therein)	2006
Si	45	Bonomo, Nucl. Phys. A 719, 37	2003
Cu	43	Bonomo, Nucl. Phys. A 719, 37	2003
Lu	40	Raoula, Eur. Phys. J. A 19, 283	2004
Yb	40	Raoula, Eur. Phys. J. A 19, 283	2004
YbD <sub>x</sub>	40	Bonomo, Nucl. Phys. A 719, 37	2003
Zr	40	Raoula, Eur. Phys. J. A 19, 283	2004
ZrD <sub>1.1</sub>	40	Raiola, Eur. Phys. J. A 19, 283	2004
Al <sub>2</sub> O <sub>3</sub>	30	Bonomo, Nucl. Phys. A 719, 37	2003
Al <sub>2</sub> O <sub>3</sub>	30	Raoula, Eur. Phys. J. A 19, 283	2004
B	30	Bonomo, Nucl. Phys. A 719, 37	2003
B	30	Raoula, Eur. Phys. J. A 19, 283	2004

Table A.1 cont.

Target	$U_e$ (eV)	Source	Year
BeO	30	Bonomo, Nucl. Phys. A 719, 37	2003
BeO	30	Raouila, Eur. Phys. J. A 19, 283	2004
Ce	30	Raouila, Eur. Phys. J. A 19, 283	2004
Dy	30	Raouila, Eur. Phys. J. A 19, 283	2004
Hf	30	Raouila, Eur. Phys. J. A 19, 283	2004
Nd	30	Raouila, Eur. Phys. J. A 19, 283	2004
Sc	30	Raouila, Eur. Phys. J. A 19, 283	2004
ScD <sub>x</sub>	30	Bonomo, Nucl. Phys. A 719, 37	2003
Sm	30	Raouila, Eur. Phys. J. A 19, 283	2004
SmD <sub>x</sub>	30	Bonomo, Nucl. Phys. A 719, 37	2003
Tb	30	Raouila, Eur. Phys. J. A 19, 283	2004
Ti	30	Raiola, Eur. Phys. J. A 27, 79	2006
Ti	30	Raouila, Eur. Phys. J. A 19, 283	2004
TiD <sub>1.3</sub>	30	Raiola, Eur. Phys. J. A 19, 283	2004
TiD <sub>x</sub>	30	Bonomo, Nucl. Phys. A 719, 37	2003
Ag	23	Bonomo, Nucl. Phys. A 719, 37	2003
Ti	19	Yuki, J. Phys. G: Nucl. Part. Phys. 23, 1459	1997
C	0	Huke, Phys. Rev. C 78, 015803	2008
CD	-20	Czerski, Nucl. Instrum. Methods Phys. Res. B 193, 183	2002

End of Table A.1

# **Appendix B**

## **Datasheets**

## EJ-200 PLASTIC SCINTILLATOR

This plastic scintillator combines the two important properties of long optical attenuation length and fast timing and is therefore particularly useful for time-of-flight systems using scintillators greater than one meter long. Typical measurements of 4 meter optical attenuation length are achieved in strips of cast sheet in which a representative size is 2 cm x 20 cm x 300 cm.

The combination of long attenuation length, high light output and an emission spectrum well matched to the common photomultipliers recommends EJ-200 as the detector of choice for many industrial applications such as gauging and environmental protection where high sensitivity of signal uniformity are critical operating requirements.

### Physical and Scintillation Constants:

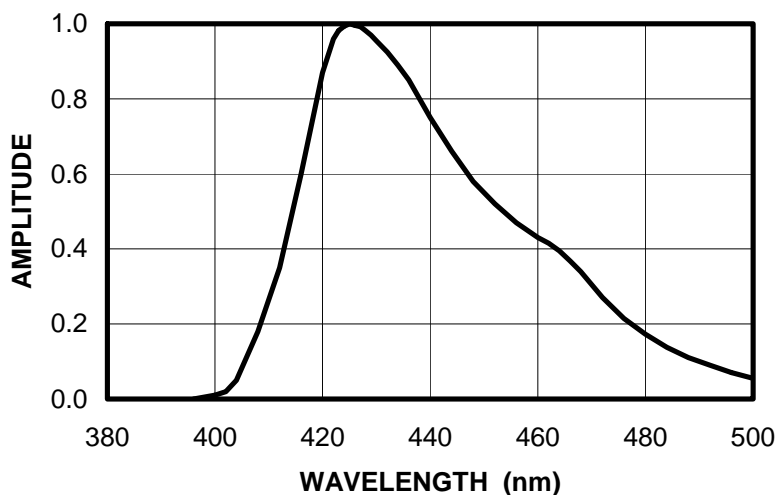
Light Output, % Anthracene .....	64
Scintillation Efficiency, photons/1 MeV e <sup>-</sup> .....	10,000
Wavelength of Max. Emission, nm .....	425
Rise Time, ns .....	0.9
Decay Time, ns .....	2.1
Pulse Width, FWHM, ns .....	~2.5
No. of H Atoms per cm <sup>3</sup> , x 10 <sup>22</sup> .....	5.17
No. of C Atoms per cm <sup>3</sup> , x 10 <sup>22</sup> .....	4.69
No. of Electrons per cm <sup>3</sup> , x 10 <sup>23</sup> .....	3.33
Density, g/cc: .....	1.023

**Polymer Base:** ..... Polyvinyltoluene  
**Refractive Index:** ..... 1.58  
**Vapor Pressure:** ..... Is vacuum-compatible  
**Coefficient of Linear Expansion:** .....  $7.8 \times 10^{-5}$  below +67°C

**Light Output vs. Temperature:**  
At +60°C, L.O. = 95% of that at +20°C  
No change from +20°C to -60°C

**Chemical Compatibility:** Is attacked by aromatic solvents, chlorinated solvents, ketones, solvent bonding cements, etc. It is stable in water, dilute acids and alkalis, lower alcohols and silicone greases. It is safe to use most epoxies and "super glues" with EJ-200.

### EJ-200 EMISSION SPECTRUM



**ELJEN TECHNOLOGY**  
1300 W. Broadway  
Sweetwater TX 79556 USA

Tel: (325) 235-4276 or (888) 800-8771  
Fax: (325) 235-0701  
Website: [www.eljentechnology.com](http://www.eljentechnology.com)

# B133D01 Photomultiplier Tube

The B133D01 is a 5" diameter 10-stage end-on photomultiplier with extended sensitivity in the blue, green and red. Designed for scintillation counting and other applications where high quantum efficiency, low dark current, good collection efficiency, and gain stability are of paramount importance.

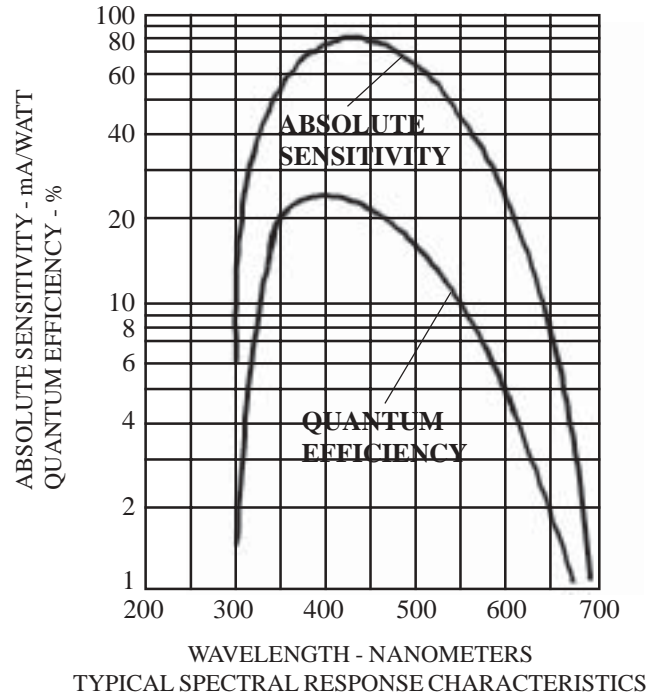
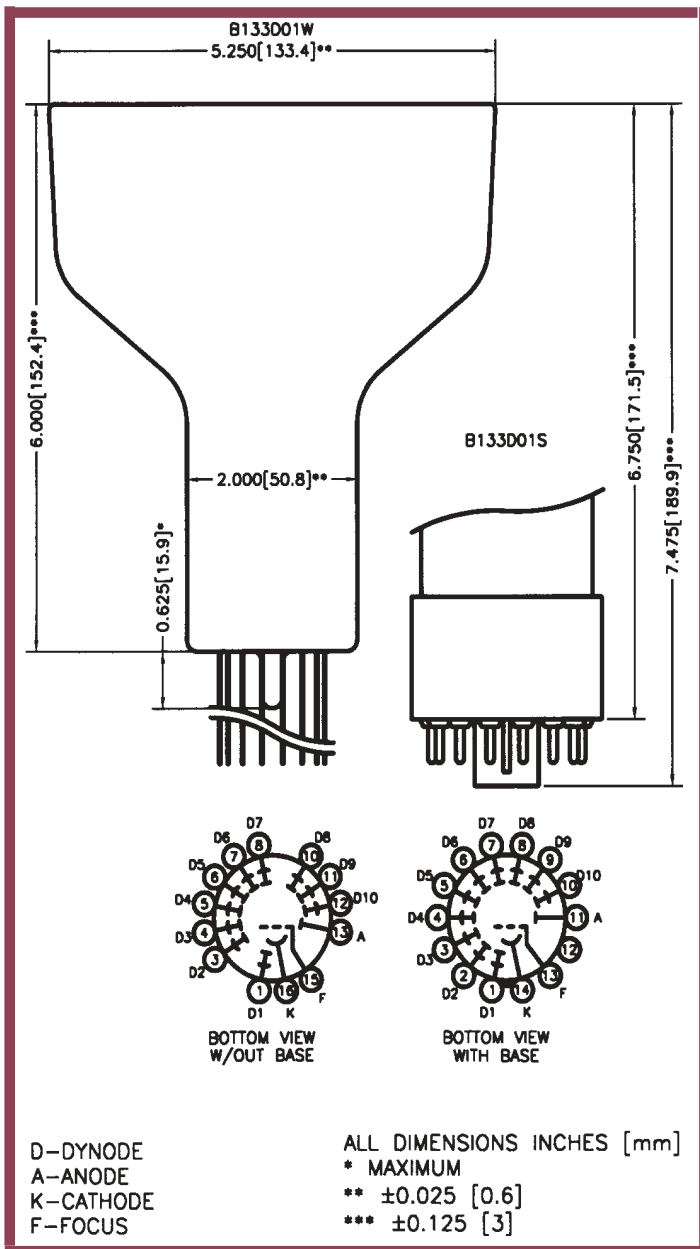


FIGURE 1

Photocathode:	Semitransparent Extended Bialkali
Spectral Response	See Figure 1
Wavelength of maximum response	470 ± 50 nm
Minimum diameter	119.38 mm
Window shape	plano-plano, circular
Window index of refraction @ 436 nm	1.523
Dynodes	BeCu, Box & Grid
Capacitance (anode to all electrodes)	9.5 pF
Operating position	Any
Weight	655 grams

# B133D01 Photomultiplier Tube

## ELECTRICAL OPERATING RATINGS

	MINIMUM	TYPICAL	MAXIMUM <sup>(5)</sup>	UNITS
Cathode to dynode No. 1 voltage	40	150	300	VDC
Cathode to anode voltage		1100	1500	VDC
Voltage between consecutive dynodes			100	VDC
Ambient storage temperature		23	60	°C
Anode current, average over 30 sec.			1.0	μA
Cathode current		1	5	μA
Cathode luminous sensitivity: <sup>(1)</sup> With 2854° K tungsten source	80	120	180	μA/lm
With blue light source <sup>(2)</sup>	5	12	15	μA/lm(B)
With red light source <sup>(3)</sup>	5	10	15	μA/lm(R)
Quantum efficiency @ 420 nm		25		%
Cathode radiant sensitivity @ 420 nm		97		mA/W
@ 540 nm		45		
@ 600 nm		25		
@ 680 nm		4		
Anode luminous sensitivity 1100 VDC: With 2854° K tungsten source of 1 x 10 <sup>-3</sup> lm	3	20	50	A/lm
Current amplification @ 1100 VDC		1 X 10 <sup>6</sup>		
Anode dark current <sup>(4)</sup> @ 22° C	1	10	20	nA

(1) With 150 VDC between cathode and all other elements connected as anode.

(2) This measurement is made with a blue filter (Corning CS-5-58, 1/2 stock thickness) interposed between a calibrated 2854° K tungsten light source and the photocathode. The (B) appearing in the units signifies that the measurement is made with the blue filter in place.

(3) This measurement is made with a red filter (Corning CS-2-62) interposed between a calibrated 2854° K tungsten light source and the photocathode. The (R) appearing in the units signifies that the measurement is made with the red filter in place.

(4) Measured at the supply voltage which gives an anode sensitivity of 20 A/lm

(5) Recommended operating maximums.

**NOTE:** When ordering one of the following basing options must be added, i.e. B133D01S

**BASING OPTIONS:** L - Long Base S - Short Base W - Wire Leads (No Base)

Voltage dividers available made to customer specifications.



**For High Energy Physics, Fast Time Response, High Pulse Linearity  
127 mm (5 Inch) Diameter, Bialkali Photocathode, 14-Stage, Head-on Type**

**GENERAL**

Parameter		Description	Unit
Spectral Response		300 to 650	nm
Wavelength of Maximum Response		420	nm
Photocathode	Material	Bialkali	—
	Minimum Effective Area	φ120	mm
Window Material		Borosilicate glass	—
Dynode	Structure	Linear focused	—
	Number of Stages	14	—
Operating Ambient Temperature		-30 to +50	°C
Storage Temperature		-30 to +50	°C
Base		20-pin base	—
Suitable Socket		E678-20B (supplied)	—

**MAXIMUM RATINGS (Absolute Maximum Values)**

Parameter		Value	Unit
Supply Voltage	Between Anode and Cathode	3000	V
	Between Anode and Last Dynode	500	V
Average Anode Current		0.2	mA

**CHARACTERISTICS (at 25 °C)**

Parameter		Min.	Typ.	Max.	Unit
Cathode Sensitivity	Luminous (2856 K)	55	70	—	μA/lm
	Blue Sensitivity Index (CS 5-58)	7.0	9.0	—	—
	Quantum Efficiency at 390 nm	—	22	—	%
Anode Sensitivity	Luminous (2856 K)	300	1000	—	A/lm
	Blue Sensitivity Index (CS 5-58)	—	130	—	—
Gain		—	1.4 × 10 <sup>7</sup>	—	—
Anode Dark Current (after 30 min storage in darkness)		—	50	300	nA
Time Response	Anode Pulse Rise Time	—	2.5	—	ns
	Electron Transit Time	—	54	—	ns
	Transit Time Spread	—	1.2	—	ns
Pulse Height Resolution with <sup>137</sup> Cs		—	8.3	—	%
Gain Deviation	Long Term	—	1.0	—	%
	Short Term	—	1.0	—	%
Pulse Linearity *	2 % Deviation	—	160	—	mA
	5 % Deviation	—	250	—	mA

**NOTE:** Measured with special voltage distribution ratios shown in the Table 2.

**Table 1: VOLTAGE DISTRIBUTION RATIO AND SUPPLY VOLTAGE**

Electrode	K	G1	G2	Dy1	Dy2	Dy3	Dy4	Dy5	Dy6	Dy7	Dy8	Dy9	Dy10	Dy11	Dy12	Dy13	Dy14	P
Ratio	2.5	7.5	0	1.2	1.8	1	1	1	1	1	1	1	1	1.5	1.5	3	2.5	

Supply Voltage: 2000 Vdc, K: Cathode, Dy: Dynode, P: Anode, G: Grid

**Table 2: SPECIAL VOLTAGE DISTRIBUTION RATIO AND SUPPLY VOLTAGE FOR PULSE LINEARITY MEASUREMENT**

Electrode	K	G1	G2	Dy1	Dy2	Dy3	Dy4	Dy5	Dy6	Dy7	Dy8	Dy9	Dy10	Dy11	Dy12	Dy13	Dy14	P	
Ratio	2.5	7.5	0	1.2	1.8	1	1	1	1	1.2	1.5	2	2.8	4	5.7	8	5		
Capacitors in μF												0.01	0.01	0.02	0.02	0.02	0.04	0.06	

Supply Voltage: 2500 Vdc, K: Cathode, Dy: Dynode, P: Anode, G: Grid

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# PHOTOMULTIPLIER TUBE R1250

Figure 1: Typical Spectral Response

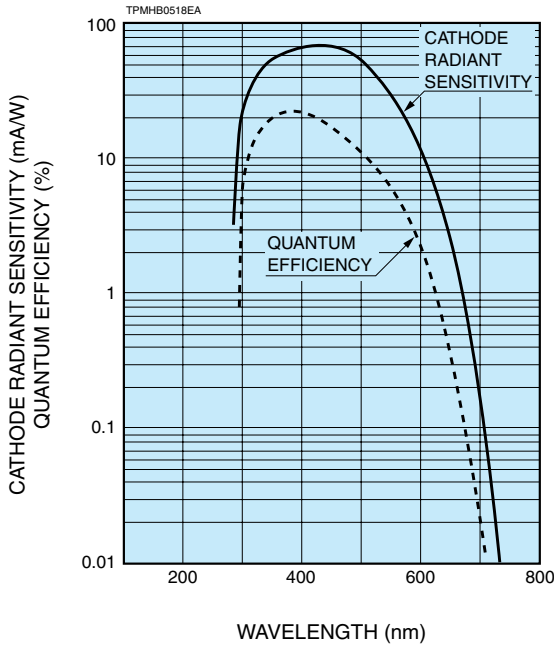


Figure 2: Typical Gain Characteristics

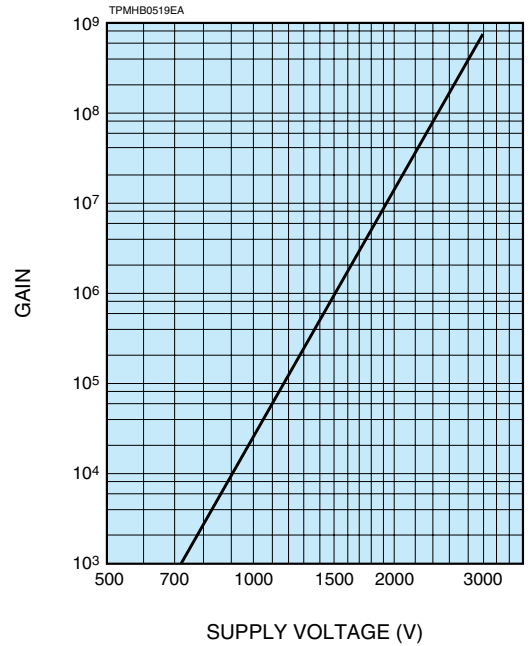
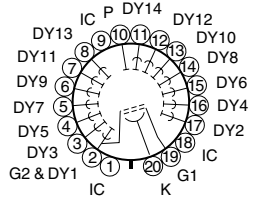
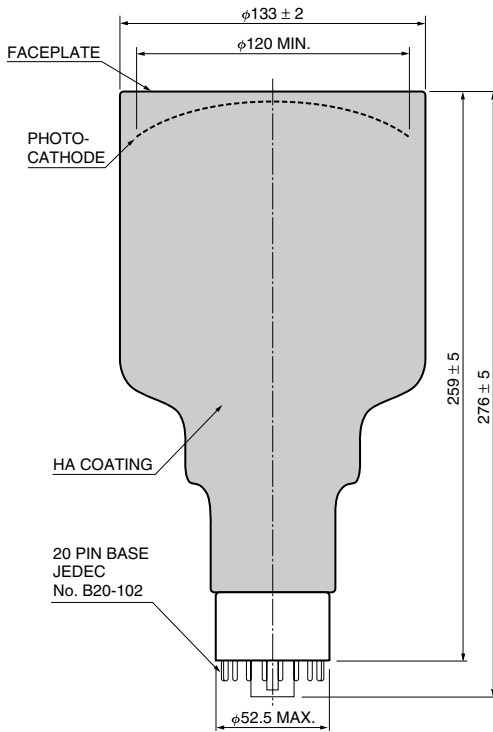
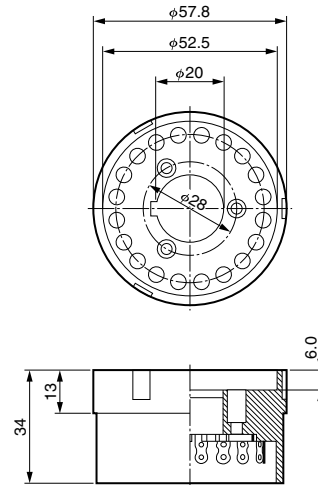


Figure 3: Dimensional Outline and Basing Diagram (Unit: mm)



Socket E678-20B (Supplied)



TACCA0309EA

TPMHA0018EB

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TPMH1213E03  
DEC. 2010 IP

# Detector Assembly Materials

Saint-Gobain Crystals can provide you with various detector assembly materials. For more detailed specifications on BC-600, BC-620 and BC-622A, individual data sheets are available.

## BC-600 Optical Cement –

BC-600 optical cement is a clear epoxy, which sets at room temperature and has a refractive index close to that of SGC plastic scintillators. It is therefore ideal for optically cementing plastic scintillators to light pipes or optical windows. It is not recommended for coupling scintillators to photomultiplier tubes. For that application, we recommend BC-634A or BC-630.

## BC-620 Reflector Paint for Plastic Scintillators –

BC-620 is a highly efficient reflector employing a special grade of titanium dioxide in a water soluble binder. It is applied directly onto plastic scintillators, acrylic light guides, glass and metals. It is not intended for direct contact with liquid scintillators. It is a diffuse reflector and, therefore, should not be applied to sheets of scintillator or light guide material where the length is much longer than the thickness.

BC-620 can be removed with warm water.

BC-620 Consists of	%
Anatase Titanium Dioxide	40
Acrylic Emulsion Resin	24
Water	32
Glycol Coalescent	2.8
Surfactants & Thickeners	1.2

## BC-622A Reflector Paint for Liquid Scintillator Tanks –

BC-622A reflector paint is intended for use with liquid scintillators. It is particularly useful in large steel or aluminum tanks, which require application of the paint at the research site. It is a diffuse reflector and, therefore, should not be used on the major surfaces of long, narrow tanks (total internal reflector and employed in these).

Can be removed from metal by submersing in Methol alcohol.

## BC-630 Silicone Optical Grease –

BC-630 is a clear, colorless, silicone, optical coupling compound that features excellent light transmission and low evaporation and bleed at 25°C. It has a specific gravity of 1.06, an Index of Refraction of 1.465 and has a very flat transmission of approximately 95% for wavelengths between 280nm and 700nm. There is a sharp fall off below 280nm. Transmission at 270 and below is about zero. We supply this single component formulation in 60ml and 500ml jars.

## Assembly Materials Available –

- Optical Cement
- Reflector Paint for Plastic Scintillators
- Reflector Paint for Liquid Scintillator Tanks
- Silicone Optical Grease
- Optical Interface Pads
- Black Wrapping Tape
- Plastic Masking Paper
- PTFE Reflector Tape



Pictured are a variety of BC-634A sizes



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## Detector Assembly Materials

### BC-634A Optical Interface Pad –

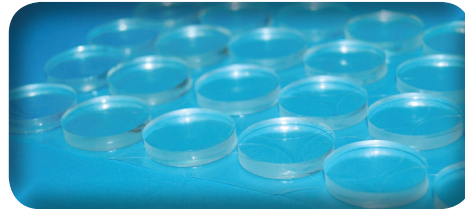
BC-634A is a self-wetting, flexible pad just hard enough to resist tearing while handling.

It is formulated for use within the temperature range of -10 to +60°C, has an index of refraction of 1.42 and an internal transmission >98% around 400nm.

If you cannot maintain sufficient interface pressure, apply a thin film of coupling grease to both sides of the interface pad.

### BC-637 High Temperature Optical Interface Pad –

BC-637 interface pads are placed between the plastic and photomultiplier tube. BC-637 is rated at 200°C.



### BC-638 Black Wrapping Tape –

BC-638 is a black adhesive tape 2" (50.8mm) wide by .008" (0.2mm) thick. Wrapping a plastic scintillator in one layer will give you a light-tight seal. We provide BC-638 in 36 yard (32.9m) rolls.

### BC-640 Plastic Masking Paper –

This material is an adhesive-backed masking paper routinely used for protecting the surfaces of plastic scintillator during handling or storage. We supply BC-640 in rolls 12" (30.5cm) wide by 300' (91.4m) long.

### BC-642 PTFE Reflector Tape –

BC-642 is a .003" (0.08 mm) thick (normal) Teflon tape and is frequently used as a reflecting material for non-hygroscopic scintillators. Three layers give you optimum reflectivity. It comes in rolls 2" (50.8mm) wide by 540" (13.7m) long.

*Manufacturer reserves the right to alter specifications.*

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(06-14)

## V1720

### 8 Channel 12bit 250 MS/s Digitizer



- o 12 bit 250 MS/s ADC
- o FPGA for real time Digital Pulse Processing:
  - o [Charge Integration \(DPP-CI\)](#)
  - o [Pulse Shape Discrimination \(DPP-PSD\)](#)
  - o Zero Suppression (Standard Firmware)
- o 8 channels
- o 2 Vpp input range (single ended or differential)
- o 16-bit programmable DC offset adjustment:  $\pm 1$  V
- o Trigger Time stamps
- o Memory buffer: 1.25 or 10 MS/ch, up to 1024 events
- o Programmable event size and pre-post trigger adjustment
- o Analog Sum/Majority and digital over/under threshold flags for Global Trigger logic
- o Front panel clock In/Out available for multiboard synchronisation (direct feed through or PLL based synthesis)
- o 16 programmable LVDS I/Os
- o Optical Link interface (CAEN proprietary protocol)
- o VME64X compliant interface
- o Firmware upgradeable via VME/Optical Link
- o A2818(PCI) / A3818 (PCIe) Controller available for handling up to 8/32 modules Daisy chained via Optical Link
- o Libraries, Demos (C and LabView) and Software tools for Windows and Linux

The **V1720** is a 1-unit wide VME 6U module housing a **8 Channel 12 bit 250 MS/s** Flash ADC Waveform Digitizer and featuring 2 Vpp single ended input dynamics. Versions with 2 Vpp differential input full scale range are also available. The DC offset adjustment (range  $\pm 1$  V) by programmable 16bit DACs (one for each channel) on single ended input versions allows a right sampling of a bipolar ( $V_{in} = \pm 1$  V) up to a full positive ( $V_{in} = 0 + +2$  V) or negative ( $V_{in} = 0 + -2$  V) analog input swing without losing dynamic resolution.

#### FPGA for real time Digital Pulse Processing (\*)

The model is available in 7 versions with different FPGA densities: **V1720/V1720B/V1720C/V1720D** equipped with Cyclone EP1C4 (4.000 LEs) and **V1720E/V1720F/V1720G** equipped with Cyclone EP1C20 (20.000 LEs).

The module features front panel Clock Input and Output as well as a PLL for clock synthesis from internal/external references. The data stream is continuously written in a circular memory buffer. When the trigger occurs, the FPGA writes further N samples for the post trigger and freezes the buffer that can be read either by VMEbus or Optical Link. The acquisition can continue without dead time in a new buffer.

Each channel has a **SRAM Multi-Event Buffer** divisible into **1 + 1024** buffers of programmable size. Two sizes of the channel digital memory are available by ordering options: **1.25 MS/ch (mod. V1720/V1720C/V1720E/V1720F)** and **10 MS/ch (mod. V1720B/V1720D/V1720G)**. 'Zero suppression' and 'data reduction' algorithms allow substantial savings in data amount readout and processing, rejecting samples smaller than programmable threshold. V1720 supports multi-board synchronization allowing all ADCs to be synchronized to a common clock source and ensuring Trigger time stamp alignment. Once synchronized, all data will be aligned and coherent across multiple V1720 boards.

The trigger signal can be provided externally via the front panel Trigger Input as well as via the software, but it can also be generated internally thanks to threshold self-trigger capability. The trigger from one board can be propagated to the other boards through the front panel Trigger Output.

An Analog Output is available with four operating modes supported:

- o **Waveform Generator:** 1 Vpp ramp generator
- o **Majority:** output signal is proportional to the number of ch. under/over threshold (1 step = 125 mV)
- o **Buffer Occupancy:** output signal is proportional to the Multi Event Buffer Occupancy: 1 buffer ~ 1 mV
- o **Voltage level:** output signal is a programmable voltage level

V1720 houses **VME** (VME64X compliant) and **Optical Link** interfaces. The VME interface allows data transfers of **60 MB/s** (MBLT64), **100 MB/s** (2eVME), **160 MB/s** (2eSST). The Optical Link supports transfer rate of **80 MB/s** and offers Daisy chain capability. Therefore, it is possible to connect up to 8/32 ADC modules to a single Optical Link controller (Mod. A2818/A3818).

#### Software available (Windows and Linux):

CAEN provides drivers for all the different types of physical communication channels, a set of C and LabView libraries (**CAENComm** and **CAENDigitizer**), demo applications and utilities:

- o **CAENSCOPE:** fully graphical program that implements a simple oscilloscope.
- o **CAENUpgrader:** tool that allows the user to update the firmware of the digitizers, change the PLL settings, load, when requested, the license for the pay firmware and other utilities.
- o **CAEN WaveDump:** software console application that can be used to configure and readout event data from any model of the CAEN digitizer family and save the data into a memory buffer allocated for this purpose.

CAEN provides also for this model two **Digital Pulse Processing firmware** for Physics Applications. This feature allows to perform on-line processing on detector signal directly digitized:

- **DPP-CI Digital Pulse Processing for the Charge Integration**  
x720 digitizer running DPP-CI firmware is well suited for data acquisition and processing of signals from scintillators/photomultipliers or SIPM detectors, implementing a digital version of the traditional QDC (Charge-to-Digital Converter).
- **DPP-PSD Digital Pulse Processing for Pulse Shape Discrimination**  
x720(\*) and x751 digitizers running DPP-PSD firmware accept signals directly from the detector and implement a digital replacement of dual gate QDC, discriminator and gate generator.

(\*) The DPP-PSD firmware runs only on V1720E/N1720F/N1720G

<b>Package</b>	1-unit wide VME 6U module
<b>Analog Input</b>	8 channels (MCX 50 Ohm) Single ended or differential Input range: 2 V <sub>pp</sub> Bandwidth: 125 MHz Programmable DAC for Offset Adjustment x channel (single ended only): ±1 V
<b>Digital Conversion</b>	Resolution: 12 bit Sampling rate: 31.25 to 250 MS/s simultaneously on each channel
<b>System Performance</b>	ENOB: 10.14 (64 kS Buffer)      SFDR: 82.0 dB SINAD: 62.85 dB                  SIGMA: 0.95 LSB rms (64 kS Buffer, open input) THD: 74.1 dB
<b>ADC Sampling Clock generation</b>	Three operating modes: - PLL mode: internal reference (50 MHz loc. oscillator) - PLL mode: external reference on CLK_IN - PLL Bypass mode: ext. clock on CLK_IN drives directly ADC clocks (Freq.: 31.25 ÷ 250 MHz)
<b>Digital I/O</b>	CLK_IN (AMP Modu II): - AC coupled differential input clock LVDS, ECL, PECL, LVPECL, CML (single ended NIM/TTL available by custom cable) - Jitter < 100 ppm CLK_OUT (AMP Modu II): - DC coupled differential LVDS clock output locked at ADC sampling clock (Freq.: 31.25 - 250 MHz) TRG_IN (NIM/TTL, Z <sub>in</sub> = 50 Ohm): external trigger input TRG_OUT (NIM/TTL, R <sub>t</sub> = 50 Ohm): local trigger output S_IN (NIM/TTL, Z <sub>in</sub> = 50 Ohm): SYNC/SAMPLE/START front panel input
<b>Memory Buffer</b>	1.25 MS/ch or 10 MS/ch Multi Event Buffer Programmable event size and pre-post trigger Divisible into 1 ÷ 1024 buffers
<b>Trigger</b>	Common Trigger - External (signal on TRG_IN) - Software (by VMEbus or Optical Link) - Self trigger (internal threshold self-trigger)
<b>Trigger Time Stamp</b>	31-bit counter - 16 ns resolution - 17 s range
<b>Multi Modules Synchronization</b>	Clock propagation: by Daisy chain or Fan Out Trigger propagation: by Daisy chain or Fan Out Time stamp synchronization
<b>ADC and Memory controller FPGA</b>	One Altera Cyclone EP1C4 or EP1C20 per channel
<b>Analog Monitor</b>	12 bit/100 MHz DAC FPGA controlled output with four operating modes: - Test Waveform: 1 V <sub>pp</sub> test ramp generator - Majority: MON/Σ output signal is proportional to the number of channels (enabled) under/over threshold (1 step = 125 mV) - Buffer Occupancy: MON/Σ output signal is proportional to the Multi Event Buffer Occupancy - Voltage level: MON/Σ output signal is a programmable voltage level
<b>LVDS I/O</b>	16 general purpose LVDS I/Os controlled by FPGA Busy, Data Ready, Memory full, Individual Trig-Out and other functions can be programmed An Input Pattern from the LVDS I/Os can be associated to each trigger as an event marker
<b>VME interface</b>	VME64X compliant Data modes: D32, BLT32, MBLT64, CBLT32/64, 2eVME, 2eSST, Multi Cast Cycles Transfer rate: 60 MB/s (MBLT64), 100 MB/s (2eVME), 160 MB/s (2eSST) Sequential and random access to the data of the Multi Event Buffer The Chained readout allows to read one event from all the boards in a VME crate with a BLT access
<b>Optical Link</b>	CAEN proprietary protocol, up to 80 MB/s transfer rate Daisy chainable: it is possible to connect up to 8/32 ADC modules to a single Optical Link Controller (Mod. A2818/A3818)
<b>Upgrade</b>	Firmware can be upgraded via VMEbus or Optical Link
<b>Software</b>	General purpose C and LabView Libraries Demo and Software Tools for Windows and Linux

# Appendix C

## Data Acquisition Code - Ugly Controller

A sample config file used to set up the digitizer for data acquisition. This particular config file was one of the actual ones used to take the data presented in this thesis.

```
1 % BYU Pulse Waveform Recorder Configuration File
2 % -----
3 % Settings common to all channels
4 % -----
5 [COMMON]
6
7 % OPEN: open the digitizer
8 % options: USB 0 0      Desktop/NIM digitizer through USB
9 %           USB 0 BA    VME digitizer through USB-V1718 (BA = BaseAddress of
           the VME board, 32 bit hex)
10 %           PCI 0 0 0   Desktop/NIM/VME through CONET (optical link)
11 %           PCI 0 0 BA  VME digitizer through V2718 (BA = BaseAddress of the
           VME board, 32 bit hex)
12 %OPEN USB 0 0
13 %OPEN USB 0 32100000
14 %OPEN PCI 0 0 0
15 OPEN PCI 0 0 32100000
16
17 % DAC_SLOPE: scales the dc offset for the dac
18 % Range 0.5 to 2.0
19 % DAC_SLOPE 1.0 % DT5720 s\n 31
20 DAC_SLOPE 1.19 % DT5720 s\n 79
21
22
23 % DAC_INTERCEPT: dac offset for calibrating the dc offset
24 % Range -1000 to 1000
25 % DAC_INTERCEPT 0 % DT5720 s\n 31
```

```
26 DAC_INTERCEPT    610 % DT5720 s\n 79
27
28 % DIRECTORY: directory to write data files to (full path, no slash at the end)
    . 100 characters max length
29 DIRECTORY D:\Data\PVT-LGBPan5\2016-02-10
30
31 % EXTENSION: data file extension (include the dot). 10 characters max length
32 EXTENSION .dat
33
34 % TOTAL_SECONDS: number of seconds to stop digitizing after. Choose 0 for
    continuous run
35 TOTAL_SECONDS 0
36
37 % TOTAL_EVENTS: number of collected events to stop digitizing after. Choose 0
    for continuous run
38 TOTAL_EVENTS 0
39
40 % EVENTS_PER_FILE: number of events to be written to file before creating a
    new file to write to
41 EVENTS_PER_FILE 10000
42
43 % POST_TRIGGER: post trigger size in percent of the whole acquisition window
44 % options: 0 to 100
45 % On models 742 there is a delay of about 35nsec on signal Fast Trigger TR;
    the post trigger is added to this delay
46 POST_TRIGGER 30
47
48 % RECORD_LENGTH = number of samples in the acquisition window
49 RECORD_LENGTH 4096
50
51 % EXTERNAL_TRIGGER: external trigger input settings. When enabled, the ext.
    trg. can be either
52 % propagated (ACQUISITION_AND_TRGOUT) or not (ACQUISITION_ONLY) through the
    TRGOUT
53 % options: DISABLED, ACQUISITION_ONLY, ACQUISITION_AND_TRGOUT
54 EXTERNAL_TRIGGER DISABLED
55
56 % ENABLE_ZLE: enable zero length encoding (set individual channel thresholds
    below)
57 % options: YES, NO
58 ENABLE_ZLE NO
59
60 % MAX_NUM_EVENTS_BLT: maximum number of events to read out in one Block
    Transfer. High values corresponds to
61 % options: 1 to 1023
62 MAX_NUM_EVENTS_BLT 50
63
64 % USE_INTERRUPT: number of events that must be ready for the readout when the
    IRQ is asserted.
65 % Zero means that the interrupts are not used (readout runs continuously)
66 USE_INTERRUPT 1
67
68 % MEM_BUFFERS: number by which CAEN internal memory is to be divided.
```

---

```
69 % Options: 0 to 10, where 0x00 is 1, 0x01 is 2, 0x02 is 4, 0x03 is 8, ... 0x0A
    is 1024.
70 MEM_BUFFERS 0
71
72 % FPIO_LEVEL: type of the front panel I/O LEMO connectors
73 % options: NIM, TTL
74 FPIO_LEVEL NIM
75
76 % TEST_PATTERN: if enabled, data from ADC are replaced by test pattern (
    triangular wave)
77 % options: YES, NO
78 TEST_PATTERN NO
79
80 % -----
81 % Individual Settings
82 % -----
83
84 % The following settings are applied specific to each channel.
85 %
86 % ENABLE_INPUT: enable/disable one channel
87 % Options: YES, NO
88 %
89 % DC_OFFSET: DC offset is in increments of approximately 2/65536 volts %
90 % Options: -2047.00 to 2047.00.
91 %
92 % TRIGGER_EDGE: rising or falling edge for trigger
93 % Options: RISING or FALLING
94 %
95 % TRIGGER_THRESHOLD: threshold for the channel auto trigger (ADC counts)
96 % are expressed in increments of 2/4096 volt. To specify zero volts use
97 % 2047. For 1.0 volt use 4095 and -1.0 volt use 0.
98 % Options: -2047 to 2047.
99 %
100 % CHANNEL_TRIGGER: sets the channel trigger to one of three modes when enabled
    :
101 % options: DISABLED, ACQUISITION_ONLY, ACQUISITION_AND_TRGOUT
102 %
103 % ZLE_THRESHOLD: zero length encoding voltage threshold. All samples not
    exceeding
104 % threshold voltage are discarded (except Pretrigger and Posttrigger samples).
105 % options -2047 to 2047.
106 %
107 % ZLE_LOGIC: POSITIVE keeps samples over ZLE_THRESHOLD and NEGATIVE keeps
    samples
108 % under ZLE_THRESHOLD.
109 % Options: POSITIVE or NEGATIVE.
110 %
111 % ZLE_PRE_THRESHOLD: number of samples to keep before ZLE_THRESHOLD is true.
112 % Options: 0 to 65535.
113 %
114 % ZLE_POST_THRESHOLD: number of samples to keep after ZLE_THRESHOLD goes
    false.
115 % Options: 0 to 65535.
116
```



---

```
117 [0]
118 ENABLE_INPUT          YES
119 CHANNEL_TRIGGER       DISABLED
120 DC_OFFSET             134
121 TRIGGER_EDGE          FALLING
122 TRIGGER_THRESHOLD     -200
123 MIN_TRIG_WIDTH        10
124 ZLE_THRESHOLD         -100
125 ZLE_POLARITY          NEGATIVE
126 ZLE_PRE_THRESHOLD     10
127 ZLE_POST_THRESHOLD   10
128
129 [1]
130 ENABLE_INPUT          YES
131 CHANNEL_TRIGGER       DISABLED
132 DC_OFFSET             157
133 TRIGGER_EDGE          FALLING
134 TRIGGER_THRESHOLD     -250
135 MIN_TRIG_WIDTH        4
136 ZLE_THRESHOLD         2040
137 ZLE_POLARITY          NEGATIVE
138 ZLE_PRE_THRESHOLD     10
139 ZLE_POST_THRESHOLD   10
140
141 [2]
142 ENABLE_INPUT          YES
143 CHANNEL_TRIGGER       DISABLED
144 DC_OFFSET             106
145 TRIGGER_EDGE          FALLING
146 TRIGGER_THRESHOLD     250
147 MIN_TRIG_WIDTH        4
148 ZLE_THRESHOLD         2040
149 ZLE_POLARITY          NEGATIVE
150 ZLE_PRE_THRESHOLD     10
151 ZLE_POST_THRESHOLD   10
152
153 [3]
154 ENABLE_INPUT          YES
155 CHANNEL_TRIGGER       DISABLED
156 DC_OFFSET             154
157 TRIGGER_EDGE          FALLING
158 TRIGGER_THRESHOLD     0
159 MIN_TRIG_WIDTH        4
160 ZLE_THRESHOLD         2000
161 ZLE_POLARITY          POSITIVE
162 ZLE_PRE_THRESHOLD     10
163 ZLE_POST_THRESHOLD   10
164
165 [4]
166 ENABLE_INPUT          YES
167 CHANNEL_TRIGGER       ACQUISITION_ONLY
168 DC_OFFSET             139
169 TRIGGER_EDGE          FALLING
170 TRIGGER_THRESHOLD     -10
```

---

```
171 MIN_TRIG_WIDTH      75
172 ZLE_THRESHOLD      2000
173 ZLE_POLARITY       POSITIVE
174 ZLE_PRE_THRESHOLD  10
175 ZLE_POST_THRESHOLD 10
176
177 [5]
178 ENABLE_INPUT        NO
179 CHANNEL_TRIGGER     DISABLED
180 DC_OFFSET           0
181 TRIGGER_EDGE        FALLING
182 TRIGGER_THRESHOLD   0
183 MIN_TRIG_WIDTH      4
184 ZLE_THRESHOLD      2000
185 ZLE_POLARITY       POSITIVE
186 ZLE_PRE_THRESHOLD  10
187 ZLE_POST_THRESHOLD 10
188
189 [6]
190 ENABLE_INPUT        NO
191 CHANNEL_TRIGGER     DISABLED
192 DC_OFFSET           33
193 TRIGGER_EDGE        FALLING
194 TRIGGER_THRESHOLD   0
195 MIN_TRIG_WIDTH      4
196 ZLE_THRESHOLD      2000
197 ZLE_POLARITY       POSITIVE
198 ZLE_PRE_THRESHOLD  10
199 ZLE_POST_THRESHOLD 10
200
201 [7]
202 ENABLE_INPUT        NO
203 CHANNEL_TRIGGER     DISABLED
204 DC_OFFSET           45
205 TRIGGER_EDGE        FALLING
206 TRIGGER_THRESHOLD   0
207 MIN_TRIG_WIDTH      4
208 ZLE_THRESHOLD      2000
209 ZLE_POLARITY       POSITIVE
210 ZLE_PRE_THRESHOLD  10
211 ZLE_POST_THRESHOLD 10
```

# Appendix D

## Data Analysis Code - ToFSpec

The following is the data analysis code, ToFSpec. It consists of a single main script and several functions that each perform specific common tasks. I started trying to adapt existing software (Anspec) to suit our analysis needs but quickly abandoned the attempt, opting instead to do a total rewrite. I have used the same option menu styles as Anspec in places in an effort to ease the transition between the two when it is necessary to use both regularly.

```
1 % Main File
2 %
3 % Use to view and analyze events from .dat files produced by Best
4 % Controller and a CAEN DT5720 digitizer.
5 % Also works with .dat files from CAEN V1720 and byuCAENcorder4 controller.
6 %
7 % Portions of code used from several sources (JEE, Anspec)
8 %
9 % Created 2 Oct 2015
10 % By Alec Raymond (alec.raymond@yahoo.com)
11 % Last Updated 30 Mar 2016
12 % By Alec Raymond (alec.raymond@yahoo.com)
13
14 % To-Do List
15 % add an option to view/change metadata
16 % implement metadata generally/eliminate all magic numbers - nearly complete
17 % investigate loading method (much faster via matlab open than "load"
18 %   function in a script)
19 %   partial loading - matfile function
20 % generalize references to reduced/combined data - currently have to
```

```
21 % "combine" data even if you only have a single channel
22 % do not call getEvent after final event of incomplete files - may need
23 % fine tuning to work when headerEnabled = 0
24 % make interpolation in combineWaveforms more procedural - unnecessary for
25 % summed waveforms
26 % find better way to calculate time-of-flight (enforce start to recoil
27 % only)
28
29 clear;
30
31 % metadata defaults
32 metadata.headersEnabled = 1; % 1 = true; 0 = false
33 metadata.channelsEnabled = 5; % number of channels enabled
34 metadata.samplesPerEventPerChannel = 4096; % in samples (4 ns per sample)
35 metadata.maxEventsPerFile = 10000; % set using digitizer controller
36 metadata.noiseThreshold = 15; % for checking for valid pulses
37 metadata.captureThreshold = 15; % for checking for valid captures
38 metadata.numEventsPerDataReducedFile = 100000; % set size of dataReduced
   output files
39 metadata.numEventsPerPulsesFoundFile = 100000; % set size of pulsesFound
   output files
40 metadata.smoothSpan = 1; % span for moving average for smoothing waveform
41 metadata.highThreshold = 10; % threshold for pulse finder
42 metadata.lowThreshold = 8; % threshold for pulse finder
43 metadata.widthHighThreshold = 5; % how long a pulse must stay above threshold
44 metadata.widthLowThreshold = 20; % how long a tail must stay below threshold
45 metadata.triggerLocation = 2867; % approx channel of trigger
46 metadata.triggerWidth = 75; % width of trigger in channels
47 metadata.triggerPorch = 100; % for finding approx capture start location -
   choose by plotting raw events
48 metadata.zeroDelayChannel = 1; % the channel containing the zero-delay
   waveform (start counting from 1 rather than 0)
49 metadata.longestDelayChannel = 4; % the channel containing the longest-delay
   waveform (all other delays must be contiguously increasing between these
   two channels)
50 metadata.pulseFrontPorch = 3; % number of channels to save in front of pulse
   start channel
51 metadata.pulseBackPorch = 3; % number of channels to save after pulse end
   channel
52 metadata.stopCableDelayLength = 75.2; % in channels (this is difference
   between stop and start cable lengths, not total stop cable length)
53 metadata.captureStartChannel = 10700; % determine using stopPulseStart
   histogram
54 metadata.earlyAreaCutoff = 15; % channels to include in early area after start
   of pulse
55 metadata.bins = 1000; % number of bins for histograms
56 metadata.distance = 1; % distance from source to stop detector in m
57 metadata.minEnergy = .1; % min energy of interest in MeV
58 metadata.maxEnergy = 10; % max energy of interest in MeV
59
60 choice=' ';
61 while choice==' '
62     fprintf(1, '\nWhat do you want to do?\n');
63     fprintf(1, ' 0: Quit\n');
```

```
64     fprintf(1, ' 1: Plot events from .dat file\n');
65     fprintf(1, ' 2: Remove invalid events\n');
66     fprintf(1, ' 3: Plot valid events\n');
67     fprintf(1, ' 4: Combine waveforms\n');
68     fprintf(1, ' 5: Plot combined events\n');
69     fprintf(1, ' 6: Find pulses\n');
70     fprintf(1, ' 6a: Find pulses TEST\n');
71     fprintf(1, ' 7: Analyze found pulses\n');
72     choice=input('Select Item> ', 's');
73     switch choice
74     %
75         case '0' % Quit
76             break;
77     %
78         case '1' % Event-by-event inspection of .dat file
79             [dir, files] = getFiles();
80             plotRawEvents(dir, files, metadata);
81             choice=' ';
82     %
83         case '2' % Remove invalid events - saves out .mat file
84             [dir, files] = getFiles();
85             saveDir = getSaveDirectory();
86             reduceData(dir, files, saveDir, metadata);
87             choice=' ';
88     %
89         case '3' % Plot valid events from .mat file
90             [dir, files] = getFiles();
91             plotValidEvents(dir, files, metadata);
92             choice=' ';
93     %
94         case '4' % Combine delayed waveforms from .mat file
95             [dir, files] = getFiles();
96             saveDir = getSaveDirectory();
97             combineWaveforms(dir, files, saveDir, metadata);
98             choice=' ';
99     %
100        case '5' % Plot combined events from .mat file
101            [dir, files] = getFiles();
102            plotCombinedEvents(dir, files);
103            choice=' ';
104        %
105        case '6' % Find all pulses in all waveforms from .mat file
106            [dir, files] = getFiles();
107            saveDir = getSaveDirectory();
108            findPulses(dir, files, saveDir, metadata);
109            choice=' ';
110        %
111        case '6a' % Find all pulses in all waveforms from .mat file
112            [dir, files] = getFiles();
113            saveDir = getSaveDirectory();
114            findPulsesTest(dir, files, saveDir, metadata);
115            choice=' ';
116        %
117        case '7' % Analyze found pulses from .mat file
```

```

118         [dir, files] = getFiles();
119         analyzePulses(dir, files, metadata);
120         choice=' ';
121 %-----
122         otherwise %Invalid input
123             choice=' ';
124     end
125 end

```

This function is used to plot unprocessed events directly from the .dat files from the digitizer.

```

1 function plotRawEvents(dir, files, metadata)
2
3 % Use to plot raw waveforms in .dat files starting at a user defined event
4
5 channelsEnabled = metadata.channelsEnabled;
6 maxEventsPerFile = metadata.maxEventsPerFile;
7
8 fileCount = length(files);
9
10 % get starting point from user
11 jstart = input('\nWhich file do you want to start with? ');
12 istart = input('Which event? ');
13 fprintf('Any key to continue\nCntrl-c to quit\n');
14
15 % loop through the files
16 for j = jstart : fileCount
17
18     dataFileName = [dir, '/', char(files(j))];
19     fid = fopen(dataFileName, 'r');
20
21     % loop through the events
22     for i = istart : maxEventsPerFile
23
24         % skip to selected event
25         if j == jstart && i == istart
26             for k = 1 : istart - 1
27                 getEvent(fid, metadata);
28             end
29         end
30
31         waveforms = getEvent(fid, metadata);
32
33         close all
34         % plot event
35         for k = 1:channelsEnabled
36             subplot(channelsEnabled, 1, k)
37             % hold on
38             plot(waveforms.waveforms(:, k))
39             % legend('0', '1', '2', '3', '4', '5', '6', '7')
40             % xlim([2000 2030])

```

```

41         ylabelStr = ['Waveform ' num2str(k)];
42         ylabel(ylabelStr)
43         if k == 1
44             titleStr = ['Event ' num2str(i)];
45             title(titleStr)
46         end
47     %         hold off
48     end
49     pause
50 end
51 fclose(fid);
52 end

```

This function is used to select the .dat files to be processed.

```

1 function [dir, files]=getFiles()
2
3 % Use to navigate to and select files for processing.
4
5 dir = uigetdir('', 'Select folder that contains raw data files');
6 files = cellstr(uigetfile([dir, '\*.dat'], 'Select file', 'MultiSelect', 'on'));

```

The function is used to choose where to save output files to.

```

1 function dir=getSaveDirectory()
2
3 % Use to choose where to save output files.
4
5 dir = uigetdir('', 'Select folder to save output files to');

```

This function reads events from .dat files.

```

1 function [waveforms, header]=getEvent(fid, metadata)
2
3 % Returns waveforms and header (if enabled)
4
5 headersEnabled = metadata.headersEnabled;
6 channelsEnabled = metadata.channelsEnabled;
7 samplesPerEventPerChannel = metadata.samplesPerEventPerChannel;
8
9 % read headers if enabled
10 if headersEnabled == 1
11     header = fread(fid, 4, 'ubit32');
12 end
13
14 % read waveforms

```

```

15 for i = 1:channelsEnabled
16     % find end of file and set waveforms to empty so reduceData stops
17     % looking for events
18     if feof(fid)
19         waveforms = [];
20         break
21     end
22     waveforms.waveforms(:,i) = fread(fid, samplesPerEventPerChannel, 'uint16')
23     ;
24 end

```

This function sets the baseline of a waveform to zero.

```

1 function waveform = adjustWaveform(waveform)
2
3 % identify baseline
4 waveformBase = mode(waveform);
5
6 % set baseline to zero
7 waveform = waveform - waveformBase;

```

This function checks for valid pulses from both detectors and throws away events that do not qualify.

```

1 function reduceData(dir, files, saveDir, metadata)
2
3 % Use to remove invalid events from the data set.
4 %
5 % An invalid event is an event that does not have any peaks above the noise
6 % level in one or more of its waveforms.
7 %
8 % The .mat file that this script saves out contains all information from
9 % the original data set except the following:
10 %
11 %   headers
12 %   events with channels that contain no waveform
13 %   ability to easily reprocess into a different format
14 %
15 % After checking output files to ensure the reduction process was set up
16 % correctly, the original data can be deleted with no loss of information
17 % except as noted above
18 %
19 % The .mat file uses ~80% less memory than the original .dat files
20 % (depending greatly on proportion of invalid events - typically 1/3 are valid
21 % )
22 % Threw an error using "save" 3 Nov 2015 on 58th .dat file
23 % trying 50 .dat file batches

```



```
24 % shouldn't have run out of memory anywhere, still unresolved;
25 % error has not occurred since then
26
27
28 % set constants
29 noiseThreshold = metadata.noiseThreshold; % channels (CAEN spreads 2 V over
    4096 channels on y axis)
30 captureThreshold = metadata.captureThreshold;
31 maxEventsPerFile = metadata.maxEventsPerFile; % set using Best Controller
32 channelsEnabled = metadata.channelsEnabled; % number of channels enabled
33 numEventsPerDataReducedFile = metadata.numEventsPerDataReducedFile;
34 triggerLocation = metadata.triggerLocation; % approx channel of trigger
35 triggerWidth = metadata.triggerWidth; % width of trigger in channels
36 triggerPorch = metadata.triggerPorch;
37
38
39 % initialize variables
40 eventValidCount = 0;
41 totalEventValidCount = 0;
42 numDataReduced = 0;
43
44 fileCount = length(files);
45 % loop through the files
46 for j = 1 : fileCount
47
48     filename = [dir, '\', char(files(j))];
49     fid = fopen(filename, 'r');
50     fprintf('\nNumber of .dat files to reduce: %u\nCurrent file: %u - %s\
        nSearching...\n', fileCount, j, filename);
51
52     % loop through the events
53     for i = 1 : maxEventsPerFile
54
55         waveforms = getEvent(fid, metadata);
56         % stop getting events after last event in file
57         if isempty(waveforms)
58             break
59         end
60
61         % use if start and stop waveforms are not added
62         eventInvalid = 0;
63
64         % check for peak above noise level in each waveform -
65         for k = 1:channelsEnabled
66             waveform = waveforms.waveforms(:,k);
67             waveform = adjustWaveform(waveform);
68             if max(abs(waveform)) <= noiseThreshold
69                 eventInvalid = 1;
70                 break
71             end
72         end
73
74         % continue to next event if invalid
75         if eventInvalid == 1
```

```
76 %         continue
77 %     end
78
79 % use if start and stop waveforms are summed into a single waveform
80 waveform = waveforms.waveforms(:,1);
81 waveform = adjustWaveform(waveform);
82 % check for start pulse peak above noise level in first waveform
83 if max(waveform(1:triggerLocation-triggerWidth-triggerPorch)) <=
    noiseThreshold
84     continue
85 end
86 % check for capture pulse peak below noise level in first waveform
87 % (sometimes the trigger occurs unexpectedly)
88 if min(waveform) >= -captureThreshold
89     continue
90 end
91 % check for recoil pulse peak below noise level in first waveform
92 if min(waveform(1:triggerLocation-triggerWidth-triggerPorch)) >= -
    noiseThreshold
93     continue
94 end
95
96 % save event if valid
97 eventValidCount = eventValidCount + 1;
98 totalEventValidCount = totalEventValidCount + 1;
99
100 for k = 1:channelsEnabled
101     waveform = waveforms.waveforms(:,k);
102     waveform = adjustWaveform(waveform);
103     dataReduced(eventValidCount).Waveforms(:,k) = waveform;
104 end
105
106 % save out dataReduced and reset it (for smaller dataReduced file
107 % sizes)
108 if length(dataReduced) == numEventsPerDataReducedFile
109     saveStr = strcat(saveDir, '\B_Reduced', int2str(numDataReduced),
110         datestr(datetime('now'), '_dd-mmm-yyyy_HH-MM-SS'));
111     save(saveStr, 'dataReduced')
112     numDataReduced = numDataReduced + 1;
113     eventValidCount = 0;
114     clear dataReduced
115 end
116 fclose(fid);
117 fprintf('Valid events found: %u', totalEventValidCount);
118 end
119
120 % check to see whether any valid events were found; otherwise, dataReduced
121 % will not even be created
122
123 if eventValidCount > 0
124     fprintf('\nSaving...\n');
125     saveStr = strcat(saveDir, '\B_Reduced', int2str(numDataReduced), datestr(
        datetime('now'), '_dd-mmm-yyyy_HH-MM-SS'));
```

```

126     save(saveStr, 'dataReduced')
127     fprintf('\nComplete!\nNumber of files reduced: %u\nTotal valid events
           found: %u\n', fileCount, totalEventValidCount)
128 else
129     fprintf('\nComplete!\nNumber of files reduced: %u\nNo valid events found\n
           ', fileCount)
130 end

```

This function plots validated events for your viewing pleasure (and to make sure it's working correctly...).

```

1 function plotValidEvents(dir, files, metadata)
2
3 % Use to plot waveforms in .mat files.
4
5 channelsEnabled = metadata.channelsEnabled; % number of channels enabled
6
7 fileCount = length(files);
8
9 % loop through the files
10 for j = 1 : fileCount
11
12     matFileName = [dir, '\', char(files(j))];
13     fprintf('\nLoading .mat file...\n');
14     load(matFileName);
15     fprintf(' Any key: continue\n cntl-c: quit\n');
16
17     % loop through the events
18     for i = 1 : length(dataReduced)
19
20         % plot event
21         for k = 1:channelsEnabled
22             subplot(channelsEnabled,1,k)
23             plot(dataReduced(i).Waveforms(:,k))
24             ylabelStr = ['Waveform ' num2str(k)];
25             ylabel(ylabelStr)
26             if k == 1
27                 titleStr = ['File ' num2str(j) ' - Event ' num2str(i)];
28                 title(titleStr)
29             end
30         end
31         pause
32     end
33     fclose(fid);
34 end

```

This function combines sequentially delayed waveforms into a single waveform. If necessary, it is capable of interpolating any additional waveforms (if for example, you had one channel for a

start detector and three channels for a stop detector, it would triple the length of the start waveform and interpolate the new values while combining the three channels for the stop detector).

```

1 function combineWaveforms(dir, files, saveDir, metadata)
2
3 % Use to combine delayed waveforms into a single waveform.
4 %
5 % Assumes the zero delay channel is in the zeroDelayChannel-th channel in
6 % each event from the Reduced .mat file and the all subsequent delayed
7 % channels are contiguous in regularly increasing delay length
8 %
9 % After checking output files to ensure the combining process was set up
10 % correctly, the reduced data can be deleted with no loss of information
11
12 samplesPerEventPerChannel = metadata.samplesPerEventPerChannel; % in samples
    (4 ns per sample)
13 zeroDelayChannel = metadata.zeroDelayChannel; % start counting channels from 1
    (different from CAEN, which starts from 0)
14 longestDelayChannel = metadata.longestDelayChannel; % delayed channels should
    be contiguous and in order of increasing delay
15 numChannelsToCombine = longestDelayChannel - zeroDelayChannel + 1; % number of
    channels to combine
16
17 numDataCombined = 0;
18
19 fileCount = length(files);
20
21 % loop through the files
22 for j = 1 : fileCount
23
24     matFileName = [dir, '\', char(files(j))];
25     fprintf('\nLoading .mat file...\n');
26     load(matFileName);
27     fprintf('\nCombining waveforms...\n');
28
29     waitbarstr = ['File ' int2str(j) ' of ' int2str(fileCount) ' - Combining
        waveforms...'];
30     h = waitbar(0, waitbarstr);
31
32     % loop through the events
33     numEvents = length(dataReduced);
34     for i = 1 : numEvents
35         waitbar(i/numEvents)
36         % combine waveforms
37         waveformIndex = 1;
38         for n = 1:numChannelsToCombine:samplesPerEventPerChannel*
            numChannelsToCombine - numChannelsToCombine + 1
39             for m = 1:numChannelsToCombine
40 %                 combinedWaveform(n+numChannelsToCombine-m) =
41 %                 dataReduced(i).Waveforms(waveformIndex,m+1); % use if the
42 %                 first channel is not to be combined

```

```

43         combinedWaveform(n+numChannelsToCombine-m) = dataReduced(i)
           .Waveforms (waveformIndex,m);
44     end
45     waveformIndex = waveformIndex + 1;
46 end
47 %     % interpolate uncombined waveforms to match size
48 %     x = 1:numChannelsToCombine:samplesPerEventPerChannel*
numChannelsToCombine;
49 %     v = dataReduced(i).Waveforms(:,3); % this must be set manually
50 %     xq = 1:1:samplesPerEventPerChannel*numChannelsToCombine;
51 %     interpolatedWaveform = interp1(x,v,xq);
52 %     % fill in end of interpolated waveform
53 %     for k = 0:numChannelsToCombine - 2
54 %         interpolatedWaveform(end - k) = interpolatedWaveform(end -
numChannelsToCombine + 1);
55 %     end
56 %     % save interpolated and combined waveforms
57 %     dataCombined(i).Waveforms(:,1) = interpolatedWaveform;
58 %     dataCombined(i).Waveforms(:,2) = combinedWaveform;
59 %     % save combined waveforms
60 %     dataCombined(i).Waveforms(:,1) = combinedWaveform;
61 %     dataCombined(i).Waveforms(:,2) = interpolatedWaveform;
62 end
63 close(h)
64
65 % save out dataCombined
66 saveStr = strcat(saveDir, '\C_Combined',int2str(numDataCombined),datestr(
    datetime('now'), '_dd-mmm-yyyy_HH-MM-SS'));
67 save(saveStr, 'dataCombined')
68 numDataCombined = numDataCombined + 1;
69 end

```

This function plots combined waveforms for additional viewing pleasure.

```

1 function plotCombinedEvents(dir, files)
2
3 % Use to plot waveforms in .mat files.
4
5 fileCount = length(files);
6
7 % loop through the files
8 for j = 1 : fileCount
9
10     matFileName = [dir, '\', char(files(j))];
11     fprintf('\nLoading .mat file...\n');
12     load(matFileName);
13     fprintf(' Any key: continue\n cntl-c: quit\n');
14
15     % loop through the events
16 %     for i =
[149,240,364,504,529,732,906,928,936,975,987,1114,1275,1531,1634,1685,1844,1847,1883,19

```

```

17     for i = 1 : length(dataCombined)
18
19         % plot event
20         numWaveforms = size(dataCombined(1).Waveforms,2);
21         for k = 1:numWaveforms
22             subplot(numWaveforms,1,k)
23             plot(dataCombined(i).Waveforms(:,k))
24             ylabelStr = ['Waveform ' num2str(k)];
25             ylabel(ylabelStr)
26             if k == 1
27                 titleStr = ['File ' num2str(j) ' - Event ' num2str(i)];
28                 title(titleStr)
29             end
30         end
31         pause
32     end
33     fclose(fid);
34 end

```

The function identifies all pulses in every waveforms, saving their start locations and the values of each index in the pulse.

```

1 function findPulses(dir, files, saveDir, metadata)
2
3 % Use to pick out all pulses from all waveforms. Saves pulse start
4 % locations and values
5 %
6 % Currently only works on combined data
7 %
8 % Not yet sufficiently satisfying to justify deleting the combined data
9 % after finding pulses
10 %
11 % Very generous in determining what constitutes a pulse - rather have too
12 % much than leave good pulses out
13 %
14 % Saves out all pulses with the following structure:
15 %     event
16 %         pulse
17 %             pulseStart
18 %             pulseValues
19 %         pulse
20 %             ...
21 %         ...
22 %     event
23 %         ...
24 %     ...
25
26 highThreshold = metadata.highThreshold;
27 lowThreshold = metadata.lowThreshold;

```

```

28 numEventsPerPulsesFoundFile = metadata.numEventsPerPulsesFoundFile;
29 widthHighThreshold = metadata.widthHighThreshold;
30 widthLowThreshold = metadata.widthLowThreshold;
31 pulseFrontPorch = metadata.pulseFrontPorch;
32 pulseBackPorch = metadata.pulseBackPorch;
33 channelsEnabled = metadata.channelsEnabled;
34 triggerLocation = metadata.triggerLocation;
35 triggerWidth = metadata.triggerWidth;
36 triggerPorch = metadata.triggerPorch;
37
38 numFile = 0;
39 eventProcessed = 1;
40 fileCount = length(files);
41
42 % loop through the files
43 for j = 1 : fileCount
44     matFileName = [dir, '\\', char(files(j))];
45     fprintf('\nLoading .mat file...\n');
46     load(matFileName);
47     fprintf('\nFinding pulses...\n');
48     waitbarstr = ['File ' int2str(j) ' of ' int2str(fileCount) ' - Finding
49                 pulses...'];
49     h = waitbar(0, waitbarstr);
50     % loop through the events
51     numEvents = length(dataCombined);
52     for i = 1 : numEvents
53         waitbar(i/numEvents)
54         eventToProcess = dataCombined(:,i);
55         waveform = eventToProcess.Waveforms(:,1);
56         % smoothWaveform = eventToProcess.Waveforms(:,2);
57         pulseNum = 0;
58         % %
59         % %
60         % %
61         % %
62         % %
63         %
64         %
65         state = 'init';
66         n = 1;
67         while n <= length(waveform)
68             %
69             for n = 1 : length(waveform)
70                 switch state;
71                     case 'init' % in case waveform starts mid pulse
72                         if abs(waveform(n)) < highThreshold
73                             state = 'findStart';
74                         end
75                     case 'findStart'
76                         if abs(waveform(n)) >= highThreshold
77                             % do not search after trigger location (pulses in
78                                 capture tail)
79                             if n >= (triggerLocation - triggerWidth)*
80                                 channelsEnabled - triggerPorch
81                                 break

```

```

79         end
80         pulseStart = n;
81         duration = 0;
82         state = 'findWidth';
83     end
84     case 'findWidth'
85         if abs(waveform(n)) < highThreshold
86             state = 'findStart';
87         else
88             duration = duration + 1;
89             if duration >= widthHighThreshold
90                 state = 'findEnd';
91             end
92         end
93     case 'findEnd'
94         if waveform(pulseStart) > 0 && waveform(n) < lowThreshold
95             pulseEnd = n;
96             duration = 0;
97             state = 'staysLow';
98         elseif waveform(pulseStart) < 0 && waveform(n) > -
99             lowThreshold
100             pulseEnd = n;
101             duration = 0;
102             state = 'staysLow';
103         end
104     case 'staysLow'
105         if waveform(pulseStart) > 0 && waveform(n) < lowThreshold
106             duration = duration + 1;
107         elseif waveform(pulseStart) < 0 && waveform(n) > -
108             lowThreshold
109             duration = duration + 1;
110         else
111             state = 'findEnd';
112         end
113         if duration >= widthLowThreshold
114             state = 'savePulse';
115         end
116     case 'savePulse'
117         pulseNum = pulseNum + 1;
118         pulseStart = pulseStart - pulseFrontPorch;
119         % if porch pushes pulse start index beyond 1
120         if pulseStart < 1
121             pulseStart = 1;
122         end
123         pulseEnd = pulseEnd + pulseBackPorch;
124         % if porch pushes pulse end index beyond end of event
125         if pulseEnd > length(waveform)
126             pulseEnd = length(waveform);
127         end
128         event(eventProcessed).pulses(pulseNum).pulseStart =
129             pulseStart;
130         event(eventProcessed).pulses(pulseNum).pulseValues =
131             waveform(pulseStart:pulseEnd);
132         n = pulseEnd;

```



```

129         state = 'findStart';
130     %         % DEBUG PLOTS
131     %         subplot(2,1,1)
132     %         plot(1:length(waveform), waveform, 'b', pulseStart, waveform
    (pulseStart), 'gx', pulseEnd, waveform(pulseEnd), 'rx')
133     %         title(['Event ' num2str(i)])
134     %         subplot(2,1,2)
135     %         plot(pulseStart:pulseEnd, waveform(pulseStart:pulseEnd))
136     % %         subplot(4,1,3)
137     % %         plot(1:length(smoothWaveform), smoothWaveform, 'b',
    pulseStart, smoothWaveform(pulseStart), 'gx', pulseEnd, smoothWaveform(
    pulseEnd), 'rx')
138     % %         subplot(4,1,4)
139     % %         plot(pulseStart:pulseEnd, smoothWaveform(pulseStart:
    pulseEnd))
140     %         pause
141     %         % END DEBUG PLOTS
142     end
143     n = n + 1;
144 end
145 % save out pulses and clear event
146 if length(event) == numEventsPerPulsesFoundFile
147     saveStr = strcat(saveDir, '\D_PulsesFound', int2str(numFile), datestr
    (datetime('now'), '_dd-mmm-yyyy_HH-MM-SS'));
148     save(saveStr, 'event')
149     numFile = numFile + 1;
150     eventProcessed = 0;
151     clear event
152 end
153 eventProcessed = eventProcessed + 1;
154 end
155 close(h)
156 end
157 % save out pulses
158 fprintf('\nSaving...\n');
159 saveStr = strcat(saveDir, '\D_PulsesFound', int2str(numFile), datestr(datetime('
    now'), '_dd-mmm-yyyy_HH-MM-SS'));
160 save(saveStr, 'event')

```

This function calculates several pulse metrics and plots them in many ways. It is also responsible for area to energy conversion.

```

1 function analyzePulses(dir, files, metadata)
2
3 % Use to analyze pulses found, event by event
4 %
5 % Some trash eliminated using parameters determined by viewing plots:
6 % THESE PARAMETERS MUST BE CHANGED WITH ANY CHANGE IN HARDWARE SETUP
7
8 captureStartChannel = metadata.captureStartChannel;

```

```

9  earlyAreaCutoff = metadata.earlyAreaCutoff;
10 stopCableDelayLength = metadata.stopCableDelayLength;
11 d = metadata.distance;
12 bins = metadata.bins;
13 minEnergy = metadata.minEnergy;
14 maxEnergy = metadata.maxEnergy;
15
16 timeFactor = 10^-9; % to convert nanoseconds to seconds
17 m = 1.674927*10^-27; % mass of a neutron in kg
18 c = 2.9979*10^8; % speed of light in m/s
19 energyFactor = 1.6022*10^-13; % to convert Joules to MeV
20
21
22 choiceLocal=' ';
23 while choiceLocal==' '
24     fprintf(1, '\n Finished analyzing, what do you want to do?\n');
25     fprintf(1, ' 0: Return to main menu\n');
26     fprintf(1, ' 1: Calculate pulse characteristics (resets pulse types)\n');
27     fprintf(1, ' 2: Close all open figures\n');
28     fprintf(1, ' 3: Change number of bins (plots must be remade)\n');
29     fprintf(1, ' 4: Make vectors for plots\n');
30     fprintf(1, ' 5: Choose regions and change pulse types\n');
31     fprintf(1, ' 6: Change pulse types (hard-coded parameters)\n');
32     fprintf(1, ' 7: Make sorted plots\n');
33     choiceLocal=input('Select Item> ', 's');
34
35     switch choiceLocal
36     %-----
37         case '0' % Quit
38             break;
39     %-----
40         case '1' % Calculate pulse characteristics
41             fileCount = length(files);
42
43             pulseNum = 0;
44
45             % loop through the files
46             for i = 1 : fileCount
47                 matFileName = [dir, '\', char(files(i))];
48                 fprintf(['\nLoading ' char(files(i)) '...\n']);
49                 load(matFileName);
50                 stopCableDelayLength = input('Set stopCableDelayLength: '); %
                    use when combining multiple runs with varying delay
                    lengths
51                 fprintf('\nAnalyzing pulses...\n');
52                 % loop through the events
53                 for j = 1 : size(event, 2)
54                     % loop through the pulses
55                     clear validStartChannel % used for calculating time-of-
                    flight
56                     for k = 1 : size(event(j).pulses, 2)
57                         pulseValues = event(j).pulses(k).pulseValues;
58                         pulseStart = event(j).pulses(k).pulseStart;
59                         % check - pulses occurring after capture start

```

```
60     if pulseStart > captureStartChannel
61         continue
62     end
63     pulseNum = pulseNum + 1;
64     % identify and assign pulse detector
65     [~, peakIndex] = max(abs(pulseValues));
66     if pulseValues(peakIndex) > 0 % from start detector
67         pulse(pulseNum).type = 'start';
68     else % from start detector
69         pulse(pulseNum).type = 'stop';
70         pulseValues = pulseValues * -1; % invert negative
71         pulse
72     end
73     % calculate pulse characteristics
74     pulse(pulseNum).start = pulseStart;
75     pulse(pulseNum).pulseValues = pulseValues;
76     pulse(pulseNum).peak = max(pulseValues);
77     pulse(pulseNum).area = sum(pulseValues);
78     if earlyAreaCutoff > length(pulseValues)
79         pulse(pulseNum).earlyAreaRatio = 1;
80     else
81         pulse(pulseNum).earlyAreaRatio = sum(pulseValues
82             (1:earlyAreaCutoff))/pulse(pulseNum).area;
83     end
84     pulse(pulseNum).width = size(pulseValues,1);
85     % detector specific operations
86     if strcmp(pulse(pulseNum).type,'start') == 1 % start
87         detector
88         % check for and skip trash pulses
89         if pulse(pulseNum).area <= 0 % includes noise,
90             leading edge dips, and very occasional start
91             in capture tail pileups
92             pulseNum = pulseNum - 1; % will overwrite
93             trashed pulse info
94             continue
95         end
96         if pulse(pulseNum).earlyAreaRatio < 0 || pulse(
97             pulseNum).earlyAreaRatio > 1 % includes noise
98             and very occasional pileup pulses
99             pulseNum = pulseNum - 1;
100             continue
101         end
102         if pulse(pulseNum).peak < 300 && pulse(pulseNum)
103             .area > 5000
104             pulseNum = pulseNum - 1;
105             continue
106         end
107     else % stop detector
108         % check for and skip trash pulses
109         if pulse(pulseNum).area <= 0 % includes both noise
110             (-30000 < area < 0) and double gamma pulses (
111             area <= -30000)
112             pulseNum = pulseNum - 1;
113             continue
114         end
115     end
116 end
```

```

103         end
104         if pulse(pulseNum).earlyAreaRatio < 0 || pulse(
            pulseNum).earlyAreaRatio > 1 % includes noise
            and very occasional pileup pulses
105             pulseNum = pulseNum - 1;
106             continue
107         end
108     end
109     pulse(pulseNum).ToF = NaN; % to match vector size
110     pulse(pulseNum).energy = NaN; % to match vector size
111     % calculate time of flight and energy
112     if exist('validStartChannel','var')
113         pulse(pulseNum).ToF = pulseStart -
            validStartChannel - stopCableDelayLength;
114         % check for pulses that occur between start
115         % channel and end of stop cable delay length
116         if pulse(pulseNum).ToF < 0
117             pulseNum = pulseNum - 1;
118             continue
119         end
120         t = timeFactor*pulse(pulseNum).ToF;
121         v = d/t;
122         % check for coincidental pulses (real ones
123         % don't move faster than c...)
124         if v > c
125             pulseNum = pulseNum - 1;
126             continue
127         end
128         gamma = real(1/sqrt(1-(v/c)^2));
129         pulse(pulseNum).energy = m*c^2*(gamma-1)/
            energyFactor;
130     end
131     % enforce energy ROI
132     if pulse(pulseNum).energy < minEnergy
133         pulseNum = pulseNum - 1;
134         continue
135     end
136     if strcmp(pulse(pulseNum).type,'stop') == 1
137         if pulse(pulseNum).energy > maxEnergy
138             pulse(pulseNum).type = 'stopgamma';
139         elseif isnan(pulse(pulseNum).energy)
140             pulse(pulseNum).type = 'stopgamma';
141         end
142     end
143     validStartChannel = pulse(pulseNum).start; % assigned
            after using start channel from previous valid
            pulse
144         end
145     end
146     end
147     choiceLocal=' ';
148 %-----
149     case '2' % Close all open figures
150         close all

```

```
151     choiceLocal=' ';
152 %
153     case '3' % Change number of bins
154         answer = inputdlg('Enter number of bins:');
155         bins = str2double(answer{1});
156         choiceLocal=' ';
157 %
158     case '4' % Make vectors for plots
159         startPulseNum = 0;
160         stopPulseNum = 0;
161         stopGammaPulseNum = 0;
162         stopRecoilPulseNum = 0;
163         stopCapture1PulseNum = 0;
164         stopCapture2PulseNum = 0;
165         for m = 1:pulseNum
166             if strcmp(pulse(m).type,'start') == 1
167                 startPulseNum = startPulseNum + 1;
168                 startStart(startPulseNum) = pulse(m).start;
169                 startPeak(startPulseNum) = pulse(m).peak;
170                 startArea(startPulseNum) = pulse(m).area;
171                 startEarlyAreaRatio(startPulseNum) = pulse(m)
172                     .earlyAreaRatio;
173                 startWidth(startPulseNum) = pulse(m).width;
174                 startToF(startPulseNum) = pulse(m).ToF;
175                 startEnergy(startPulseNum) = pulse(m).energy;
176             elseif strcmp(pulse(m).type,'stopgamma') == 1
177                 stopGammaPulseNum = stopGammaPulseNum + 1;
178                 stopGammaStart(stopGammaPulseNum) = pulse(m).start;
179                 stopGammaPeak(stopGammaPulseNum) = pulse(m).peak;
180                 stopGammaArea(stopGammaPulseNum) = pulse(m).area;
181                 stopGammaEarlyAreaRatio(stopGammaPulseNum) = pulse(m)
182                     .earlyAreaRatio;
183                 stopGammaWidth(stopGammaPulseNum) = pulse(m).width;
184                 stopGammaToF(stopGammaPulseNum) = pulse(m).ToF;
185                 stopGammaEnergy(stopGammaPulseNum) = pulse(m).energy;
186             elseif strcmp(pulse(m).type,'stoprecoil') == 1
187                 stopRecoilPulseNum = stopRecoilPulseNum + 1;
188                 stopRecoilStart(stopRecoilPulseNum) = pulse(m).start;
189                 stopRecoilPeak(stopRecoilPulseNum) = pulse(m).peak;
190                 stopRecoilArea(stopRecoilPulseNum) = pulse(m).area;
191                 stopRecoilEarlyAreaRatio(stopRecoilPulseNum) = pulse(m)
192                     .earlyAreaRatio;
193                 stopRecoilWidth(stopRecoilPulseNum) = pulse(m).width;
194                 stopRecoilToF(stopRecoilPulseNum) = pulse(m).ToF;
195                 stopRecoilEnergy(stopRecoilPulseNum) = pulse(m).energy;
196             elseif strcmp(pulse(m).type,'stopcapture1') == 1
197                 stopCapture1PulseNum = stopCapture1PulseNum + 1;
198                 stopCapture1Start(stopCapture1PulseNum) = pulse(m).start;
199                 stopCapture1Peak(stopCapture1PulseNum) = pulse(m).peak;
200                 stopCapture1Area(stopCapture1PulseNum) = pulse(m).area;
                stopCapture1EarlyAreaRatio(stopCapture1PulseNum) = pulse(m)
                    .earlyAreaRatio;
                stopCapture1Width(stopCapture1PulseNum) = pulse(m).width;
                stopCapture1ToF(stopCapture1PulseNum) = pulse(m).ToF;
```

```

201         stopCapture1Energy(stopCapture1PulseNum) = pulse(m).energy
202         ;
203     elseif strcmp(pulse(m).type, 'stopcapture2') == 1
204         stopCapture2PulseNum = stopCapture2PulseNum + 1;
205         stopCapture2Start(stopCapture2PulseNum) = pulse(m).start;
206         stopCapture2Peak(stopCapture2PulseNum) = pulse(m).peak;
207         stopCapture2Area(stopCapture2PulseNum) = pulse(m).area;
208         stopCapture2EarlyAreaRatio(stopCapture2PulseNum) = pulse(m)
209             .earlyAreaRatio;
210         stopCapture2Width(stopCapture2PulseNum) = pulse(m).width;
211         stopCapture2ToF(stopCapture2PulseNum) = pulse(m).ToF;
212         stopCapture2Energy(stopCapture2PulseNum) = pulse(m).energy
213         ;
214     else
215         stopPulseNum = stopPulseNum + 1;
216         stopStart(stopPulseNum) = pulse(m).start;
217         stopPeak(stopPulseNum) = pulse(m).peak;
218         stopArea(stopPulseNum) = pulse(m).area;
219         stopEarlyAreaRatio(stopPulseNum) = pulse(m).earlyAreaRatio
220         ;
221         stopWidth(stopPulseNum) = pulse(m).width;
222         stopToF(stopPulseNum) = pulse(m).ToF;
223         stopEnergy(stopPulseNum) = pulse(m).energy;
224     end
225 end
226 choiceLocal=' ';
227 %
228 -----
229 case '5' % Choose regions and reassign pulse type to region name
230
231     % peak vs area plot - recoil selection
232     clear h vertices xvertices yvertices
233     figure
234     scatter(stopArea, stopEarlyAreaRatio, 'k.')
235     title('Early Area Ratio vs Area - Select Recoil')
236     h = impoly;
237     vertices = getPosition(h);
238     xvertices = vertices(:,1);
239     yvertices = vertices(:,2);
240     for m = 1:pulseNum
241         if strcmp(pulse(m).type, 'stop') == 1 % unsorted pulse from
242             stop detector (preserves gamma type assignment from
243             earlier energy check)
244             if strcmp(pulse(m).type, 'start') == 0 % not start detector (
245                 will overwrite previous type assignment)
246                 if inpolygon(pulse(m).area, pulse(m).earlyAreaRatio,
247                     xvertices, yvertices) == 1
248                     pulse(m).type = 'stoprecoil';
249                 else
250                     pulse(m).type = 'stopgamma';
251                 end
252             end
253         end
254     end
255 end
256 close
257 % early area ratio vs area plot - capture 1 selection

```



```

296         end
297     else
298         if pulse(m).area < 15000
299             pulse(m).type = 'stopcapture1';
300         else
301             pulse(m).type = 'stopcapture2';
302         end
303     end
304 end
305 end
306 choiceLocal=' '; % remake vectors
307 %


---


308 case '7' % Remake plots
309
310 %     figure
311 %     hold on
312 %     scatter(stopGammaArea,stopGammaPeak,'k.')
313 %     scatter(stopRecoilArea,stopRecoilPeak,'k.')
314 %     scatter(stopCapture1Area,stopCapture1Peak,'k.')
315 %     scatter(stopCapture2Area,stopCapture2Peak,'k.')
316 %     hold off
317 %     title('Peak vs Area')
318 %     xlabel('Area')
319 %     ylabel('Peak')
320 %
321 %     figure
322 %     hold on
323 %     scatter(stopGammaArea,stopGammaPeak,'g.')
324 %     scatter(stopRecoilArea,stopRecoilPeak,'r.')
325 %     scatter(stopCapture1Area,stopCapture1Peak,'c.')
326 %     scatter(stopCapture2Area,stopCapture2Peak,'b.')
327 %     hold off
328 %     title('Peak vs Area')
329 %     xlabel('Area')
330 %     ylabel('Peak')
331 %     legend('Gamma','Recoil','Capture 1','Capture 2')
332
333 figure
334 hold on
335 scatter(stopGammaArea,stopGammaEarlyAreaRatio,'k.')
336 scatter(stopRecoilArea,stopRecoilEarlyAreaRatio,'k.')
337 scatter(stopCapture1Area,stopCapture1EarlyAreaRatio,'k.')
338 scatter(stopCapture2Area,stopCapture2EarlyAreaRatio,'k.')
339 hold off
340 title('Early Area Ratio vs Total Area')
341 xlabel('Area')
342 ylabel(['Early Area Ratio (' num2str(earlyAreaCutoff) ' channels)'
343         ])
344
345 figure
346 hold on
347 scatter(stopGammaArea,stopGammaEarlyAreaRatio,'g.')
348 scatter(stopRecoilArea,stopRecoilEarlyAreaRatio,'r.')
349 scatter(stopCapture1Area,stopCapture1EarlyAreaRatio,'c.')

```



```

349         scatter(stopCapture2Area, stopCapture2EarlyAreaRatio, 'b.')
350     hold off
351     title('Early Area Ratio vs Total Area')
352     xlabel('Area')
353     ylabel(['Early Area Ratio (' num2str(earlyAreaCutoff) ' channels)'
354            ])
355     legend('Gamma', 'Recoil', 'Capture 1', 'Capture 2')
356
357     figure
358     h = histogram([stopGammaToF stopRecoilToF], bins, 'EdgeColor', 'k', '
359                 FaceColor', 'none');
360     title('ToF Histogram')
361     xlabel('ToF (ns)')
362     ylabel('Count');
363     xlim([0 500])
364
365     figure
366     hold on
367     histogram(stopGammaToF, bins, 'EdgeColor', 'g', 'FaceColor', 'none', '
368             BinWidth', h.BinWidth)
369     histogram(stopRecoilToF, bins, 'EdgeColor', 'r', 'FaceColor', 'none', '
370             BinWidth', h.BinWidth)
371     hold off
372     title('ToF Histogram')
373     xlabel('ToF (ns)')
374     ylabel('Count');
375     legend('Gamma', 'Recoil')
376     xlim([0 500])
377
378     figure
379     hold on
380     scatter(stopGammaToF, stopGammaPeak, 'k.')
381     scatter(stopRecoilToF, stopRecoilPeak, 'k.')
382     scatter(stopCapture1ToF, stopCapture1Peak, 'k.')
383     scatter(stopCapture2ToF, stopCapture2Peak, 'k.')
384     hold off
385     title('Peak vs ToF (linear)')
386     xlabel('ToF (ns)')
387     ylabel('Peak');
388     xlim([0 500])
389
390     figure
391     hold on
392     scatter(stopGammaToF, stopGammaPeak, 'g.')
393     scatter(stopRecoilToF, stopRecoilPeak, 'r.')
394     scatter(stopCapture1ToF, stopCapture1Peak, 'c.')
395     scatter(stopCapture2ToF, stopCapture2Peak, 'b.')
396     hold off
397     title('Peak vs ToF (linear)')
398     xlabel('ToF (ns)')
399     ylabel('Peak');
400     legend('Gamma', 'Recoil', 'Capture 1', 'Capture 2')
401     xlim([0 500])

```

```
399 %           figure
400 %           scatter(stopRecoilToF,stopRecoilPeak,'r.')
401 %           title('Peak vs ToF (linear)')
402 %           xlabel('ToF (ns)')
403 %           ylabel('Peak');
404 %           legend('Recoil')
405 %           xlim([0 500])
406 %
407 %           figure
408 %           hold on
409 %           scatter(stopGammaToF,stopGammaPeak,'g.')
410 %           scatter(stopRecoilToF,stopRecoilPeak,'r.')
411 %           scatter(stopCapture1ToF,stopCapture1Peak,'c.')
412 %           scatter(stopCapture2ToF,stopCapture2Peak,'b.')
413 %           set(gca,'xscale','log','yscale','log')
414 %           hold off
415 %           title('Peak vs ToF (log)')
416 %           xlabel('ToF (ns)')
417 %           ylabel('Peak');
418 %           legend('Gamma','Recoil','Capture 1','Capture 2')
419 %           xlim([1 10^4])
420 %
421 %           figure
422 %           scatter(stopRecoilToF,stopRecoilPeak,'r.')
423 %           set(gca,'xscale','log','yscale','log')
424 %           title('Peak vs ToF (log)')
425 %           xlabel('ToF (ns)')
426 %           ylabel('Peak');
427 %           legend('Recoil')
428 %           xlim([1 10^4])
429 %
430 %           figure
431 %           scatter(stopRecoilEnergy,stopRecoilPeak,'r.')
432 %           title('Peak vs Energy (linear)')
433 %           xlabel('Energy (MeV)')
434 %           ylabel('Peak');
435 %           legend('Recoil')
436 %           xlim([minEnergy maxEnergy])
437 %
438 %           figure
439 %           scatter(stopRecoilEnergy,stopRecoilPeak,'r.')
440 %           set(gca,'xscale','log','yscale','log')
441 %           title('Peak vs Energy (log)')
442 %           xlabel('Energy (MeV)')
443 %           ylabel('Peak');
444 %           legend('Recoil')
445 %
446 %           figure
447 %           hold on
448 %           scatter(stopGammaToF,stopGammaArea,'k.')
449 %           scatter(stopRecoilToF,stopRecoilArea,'k.')
450 %           scatter(stopCapture1ToF,stopCapture1Area,'k.')
451 %           scatter(stopCapture2ToF,stopCapture2Area,'k.')
452 %           hold off
```

```

453 %         title('Area vs ToF (linear)')
454 %         xlabel('ToF (ns)')
455 %         ylabel('Area');
456 %         xlim([0 500])
457 %
458 %         figure
459 %         hold on
460 %         scatter(stopGammaToF,stopGammaArea,'g.')
461 %         scatter(stopRecoilToF,stopRecoilArea,'r.')
462 %         scatter(stopCapture1ToF,stopCapture1Area,'c.')
463 %         scatter(stopCapture2ToF,stopCapture2Area,'b.')
464 %         hold off
465 %         title('Area vs ToF (linear)')
466 %         xlabel('ToF (ns)')
467 %         ylabel('Area');
468 %         legend('Gamma','Recoil','Capture 1','Capture 2')
469 %         xlim([0 500])
470 %
471 %         figure
472 %         scatter(stopRecoilToF,stopRecoilArea,'r.')
473 %         title('Area vs ToF (linear)')
474 %         xlabel('ToF (ns)')
475 %         ylabel('Area');
476 %         legend('Recoil')
477 %         xlim([0 500])
478 %
479 %         figure
480 %         hold on
481 %         scatter(stopGammaToF,stopGammaArea,'g.')
482 %         scatter(stopRecoilToF,stopRecoilArea,'r.')
483 %         scatter(stopCapture1ToF,stopCapture1Area,'c.')
484 %         scatter(stopCapture2ToF,stopCapture2Area,'b.')
485 %         set(gca,'xscale','log','yscale','log')
486 %         hold off
487 %         title('Area vs ToF (log)')
488 %         xlabel('ToF (ns)')
489 %         ylabel('Area');
490 %         legend('Gamma','Recoil','Capture 1','Capture 2')
491 %         xlim([1 10^4])
492 %
493 %         figure
494 %         scatter(stopRecoilToF,stopRecoilArea,'r.')
495 %         set(gca,'xscale','log','yscale','log')
496 %         title('Area vs ToF (log)')
497 %         xlabel('ToF (ns)')
498 %         ylabel('Area');
499 %         legend('Recoil')
500 %         xlim([1 10^4])
501 %
502 %         figure
503 %         h = histogram(stopRecoilEnergy,bins,'EdgeColor','r','FaceColor','
504 %             none');
505 %         hold on
506 %         % fit energy function to histogram

```

```

506     xvalues = h.BinEdges(1:end-1);
507     yvalues = h.Values;
508     [xData, yData] = prepareCurveData( xvalues, yvalues );
509     % find and remove empty bins
510     emptyBins = find(yData == 0);
511     xData(emptyBins) = [];
512     yData(emptyBins) = [];
513     %
514     ft = fitype( 'a*x.^(1/2).*exp(-x/1.42)', 'independent', 'x', '
dependent', 'y' ); % MCNP manual - Maxwellian Spectrum
515     ft = fitype( 'a*exp(-x/1.025).*sinh((2.926*x).^(1/2))', '
independent', 'x', 'dependent', 'y' ); % MCNP manual - Watt
Spectrum
516     opts = fitoptions( 'Method', 'NonlinearLeastSquares' );
517     opts.Display = 'Off';
518     opts.StartPoint = 10;
519     [fitresultEnergyHist, ~] = fit( xData, yData, ft, opts );
520     h = plot(fitresultEnergyHist, 'b-', xData, yData, 'b. ');
521     set(h, 'LineWidth', 2)
522     title('Energy Histogram (linear)')
523     xlabel('Energy (MeV)')
524     ylabel('Count');
525     legend('Recoil', 'Data for fit', 'Watt Spectrum')
526     %
527     xlim([minEnergy maxEnergy])
528     hold off
529
530     figure
531     histogram(stopRecoilEnergy, bins, 'EdgeColor', 'r', 'FaceColor', 'none'
);
532
533     hold on
534     h = plot(fitresultEnergyHist, 'b-', xData, yData, 'b. ');
535     set(h, 'LineWidth', 2)
536     set(gca, 'xscale', 'log', 'yscale', 'log')
537     title('Energy Histogram (log)')
538     xlabel('Energy (MeV)')
539     ylabel('Count');
540     legend('Recoil', 'Data for fit', 'Watt Spectrum')
541     xlim([minEnergy maxEnergy])
542     hold off
543
544     figure
545     scatter(stopRecoilEnergy, stopRecoilArea, 'r.')
546     hold on
547     % initial fit - area to energy calibration function
548     xvalues = stopRecoilEnergy;
549     yvalues = stopRecoilArea;
550     [xData, yData] = prepareCurveData( xvalues, yvalues );
551     ft = fitype( 'a*x.^(3/2)', 'independent', 'x', 'dependent', 'y' )
; % Knoll 2010, p577
552     opts = fitoptions( 'Method', 'NonlinearLeastSquares' );
553     opts.Display = 'Off';
554     opts.StartPoint = 10;
555     [fitresultAreavsEnergy, ~] = fit( xData, yData, ft, opts );
556     h = plot(fitresultAreavsEnergy, 'b-');
557     set(h, 'LineWidth', 2)

```

```

555     title('Area vs Energy (linear)')
556     xlabel('Energy (MeV)')
557     ylabel('Area');
558     ylim([0 max(stopRecoilArea)])
559     legend('Recoil','Least-squares fit')
560 %     xlim([minEnergy maxEnergy])
561     hold off
562
563     figure
564     scatter(stopRecoilEnergy,stopRecoilArea,'r.')
565     hold on
566     h = plot(fitresultAreavsEnergy, 'b-');
567     set(h,'LineWidth',2)
568     set(gca,'xscale','log','yscale','log')
569     xlim([minEnergy maxEnergy])
570     ylim([0 max(stopRecoilArea)])
571     title('Area vs Energy (log)')
572     xlabel('Energy (MeV)')
573     ylabel('Area');
574     legend('Recoil')
575     hold off
576
577     % hand tweak initial value for proportionality constant
578     handtweak = 1;
579     while handtweak ~=0
580         fprintf(1,'\nChoose new proportionality constant?\n');
581         fprintf(1,'    Current value is %f\n',fitresultAreavsEnergy.a);
582         handtweak = input('    Enter new value (0 to keep current value
583             ): ');
584         if handtweak ~= 0
585             fitresultAreavsEnergy.a = handtweak;
586         else
587             break
588         end
589
590     figure
591     scatter(stopRecoilEnergy,stopRecoilArea,'r.')
592     hold on
593     h = plot(fitresultAreavsEnergy, 'b-');
594     set(h,'LineWidth',2)
595     title('Area vs Energy (linear)')
596     xlabel('Energy (MeV)')
597     ylabel('Area');
598     ylim([0 max(stopRecoilArea)])
599     legend('Recoil','Least-squares fit')
600     hold off
601
602     figure
603     scatter(stopRecoilEnergy,stopRecoilArea,'r.')
604     hold on
605     h = plot(fitresultAreavsEnergy, 'b-');
606     set(h,'LineWidth',2)
607     set(gca,'xscale','log','yscale','log')

```

```

608         xlim([minEnergy maxEnergy])
609         ylim([0 max(stopRecoilArea)])
610         title('Area vs Energy (log)')
611         xlabel('Energy (MeV)')
612         ylabel('Area');
613         legend('Recoil','Least-squares fit')
614         hold off
615     end
616
617     % calculate energy histogram from area histogram
618     k = fitresultAreavsEnergy.a;
619     energyFromArea = (stopRecoilArea./k).^(2/3);
620     figure
621     henergyFromArea = histogram(energyFromArea,bins,'EdgeColor','r','
        FaceColor','none');
622     hold on
623     % fit energy function to histogram
624     xvalues = henergyFromArea.BinEdges(1:end-1);
625     yvalues = henergyFromArea.Values;
626     [xData, yData] = prepareCurveData( xvalues, yvalues );
627     % find and remove empty bins
628     emptyBins = find(yData == 0);
629     xData(emptyBins) = [];
630     yData(emptyBins) = [];
631     %
632     ft = fitype( 'a*x.^(1/2).*exp(-x/1.42)', 'independent', 'x', '
dependent', 'y' ); % MCNP manual - Maxwellian Spectrum
633     ft = fitype( 'a*exp(-x/1.025).*sinh((2.926*x).^(1/2))', '
independent', 'x', 'dependent', 'y' ); % MCNP manual - Watt
        Spectrum
634     opts = fitoptions( 'Method', 'NonlinearLeastSquares' );
635     opts.Display = 'Off';
636     opts.StartPoint = 10;
637     [fitresultEnergyHist, ~] = fit( xData, yData, ft, opts );
638     h = plot(fitresultEnergyHist, 'b-', xData, yData, 'b. ');
639     set(h, 'LineWidth', 2)
640     legend('Recoil','Data for fit','Watt Spectrum')
641     %
642     xlim([minEnergy maxEnergy])
643     title('Energy Histogram from Area')
644     xlabel('Energy (MeV)')
645     ylabel('Count');
646     hold off
647
648     % correct for Hydrogen cross section - data from https://www-
        nds.iaea.org/exfor/endl.htm
649     % get H cross section data
650     scalexAxis = xlsread('D:\Code\ToFSpec\Resources\
        HnCrossSectionData.xlsx', 'A:A');
651     scaleyAxis = xlsread('D:\Code\ToFSpec\Resources\
        HnCrossSectionData.xlsx', 'B:B');
652     % find ROI
653     [~, scaleMinROIIndex] = min(abs(scalexAxis - min(
        henergyFromArea.BinEdges)));
654     [~, scaleMaxROIIndex] = min(abs(scalexAxis - max(
        henergyFromArea.BinEdges)));

```

```

653     scalexAxis = scalexAxis(scaleMinROIIndex:scaleMaxROIIndex);
654     scaleyAxis = scaleyAxis(scaleMinROIIndex:scaleMaxROIIndex);
655     % normalize to 1 for scaling
656     scaleyAxis = scaleyAxis/max(scaleyAxis);
657     % match vector size with energyFromArea histogram BinEdges
658     scalexq = linspace(henergyFromArea.BinEdges(1),
        henergyFromArea.BinEdges(end-1),bins);
659     scalevq1 = interp1(scalexAxis,scaleyAxis,scalexq,'pchip');
660     % build scaling vector
661     scaleVector = 1./scalevq1;
662     % scale BinValues
663     scaleEnergyFromArea = (henergyFromArea.Values).*scaleVector;
664     % make plots
665     figure
666     scatter(scalexAxis,scaleyAxis,'ro')
667     hold on
668     scatter(scalexq,scalevq1,'b.')
669     scatter(scalexq,scaleVector,'y.')
670     scatter(scalexq,scaleVector.*scalevq1,'g.')
671     legend('effect vector','interpolated effect vector','correction
        vector','corrected effect vector','Location','best')
672     hold off
673
674     % refit to corrected calculated energy spectrum
675     % fit energy function to histogram
676     xvalues = henergyFromArea.BinEdges(1:end-1);
677     yvalues = scaleEnergyFromArea;
678     [xData, yData] = prepareCurveData( xvalues, yvalues );
679     % find and remove empty bins
680     emptyBins = find(yData == 0);
681     xData(emptyBins) = [];
682     yData(emptyBins) = [];
683     %
684     ft = fittype( 'a*x.^(1/2).*exp(-x/1.42)', 'independent', 'x', '
dependent', 'y' ); % MCNP manual - Maxwellian Spectrum
685     ft = fittype( 'a*exp(-x/1.025).*sinh((2.926*x).^(1/2))', '
independent', 'x', 'dependent', 'y' ); % MCNP manual - Watt
        Spectrum
686     opts = fitoptions( 'Method', 'NonlinearLeastSquares' );
687     opts.Display = 'Off';
688     opts.StartPoint = 10;
689     [fitresultEnergyHist, ~] = fit( xData, yData, ft, opts );
690     figure
691     h = plot(fitresultEnergyHist, 'b-');
692     hold on
693     set(h,'LineWidth',2)
694     plot(xData, yData, 'r-');
695     legend('Watt Spectrum','Data for fit')
696     %
697     xlim([minEnergy maxEnergy])
698     title('Energy Histogram from Area - Corrected')
699     xlabel('Energy (MeV)')
700     ylabel('Count');
701     hold off
702
703     choiceLocal=' ';

```

```
702 %  
703     otherwise %Invalid input  
704         choiceLocal=' '  
705     end  
706 end
```



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