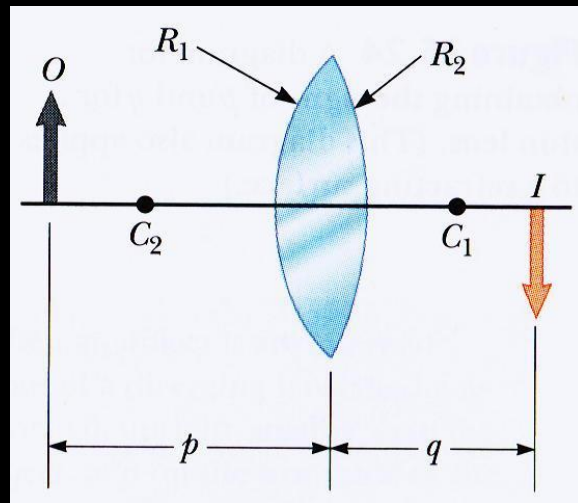
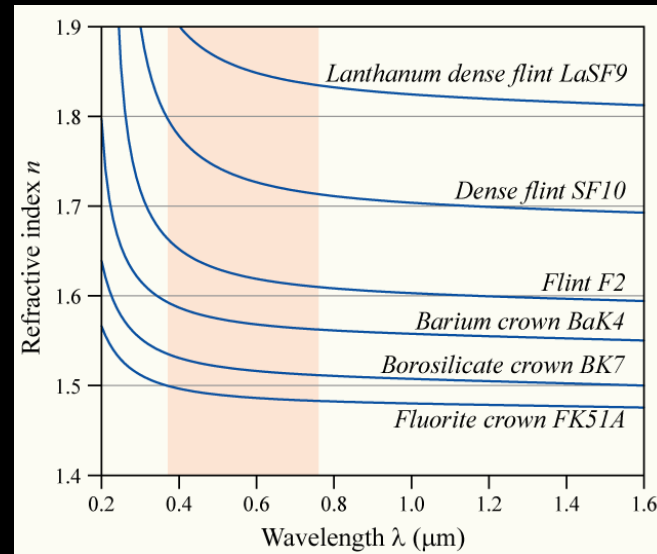
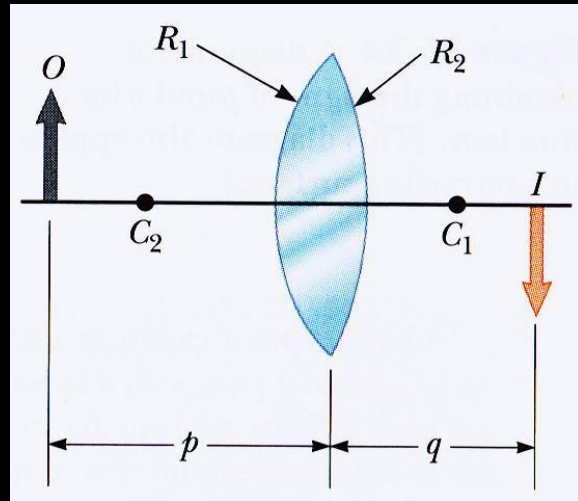


# Quiz



- Red light coming from  $O$  produces an image at  $I$ . Where will blue light coming from  $O$  form an image?
  - same place
  - closer to the lens
  - farther from lens

# Chromatic Aberration

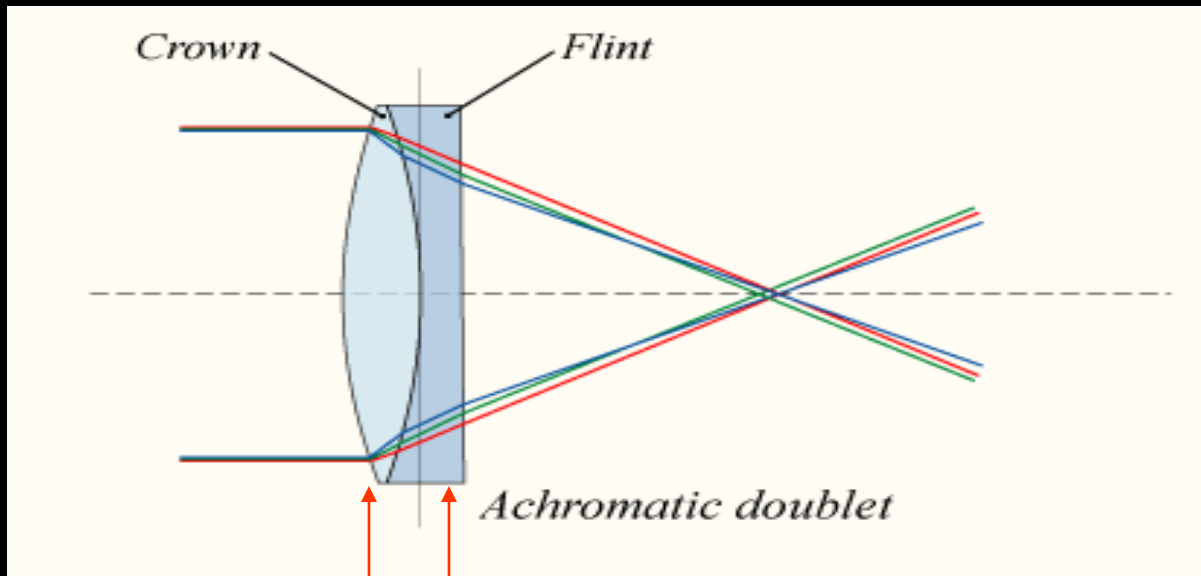
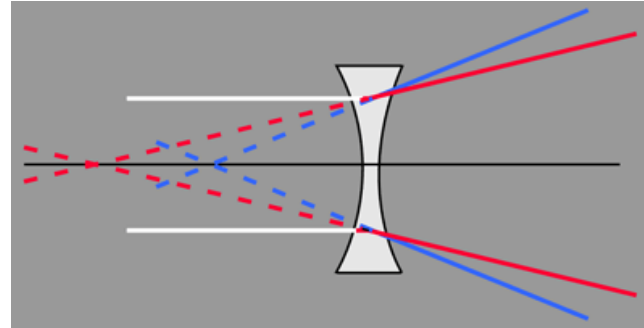
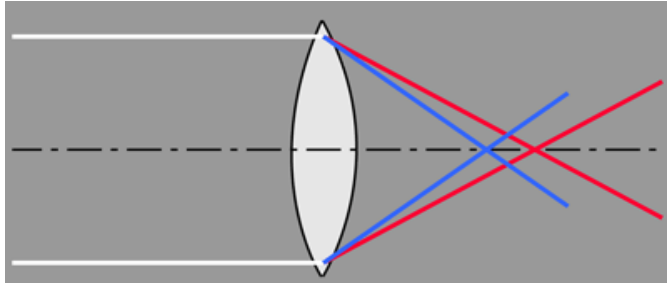


- Red light coming from  $O$  produces an image at  $I$ . Where will blue light coming from  $O$  form an image?
  - a. same place
  - b. closer to the lens
  - c. farther from lens

Lens-makers' eqn:

$$\frac{1}{f} = (n - 1) \left( \frac{1}{R_1} - \frac{1}{R_2} \right)$$

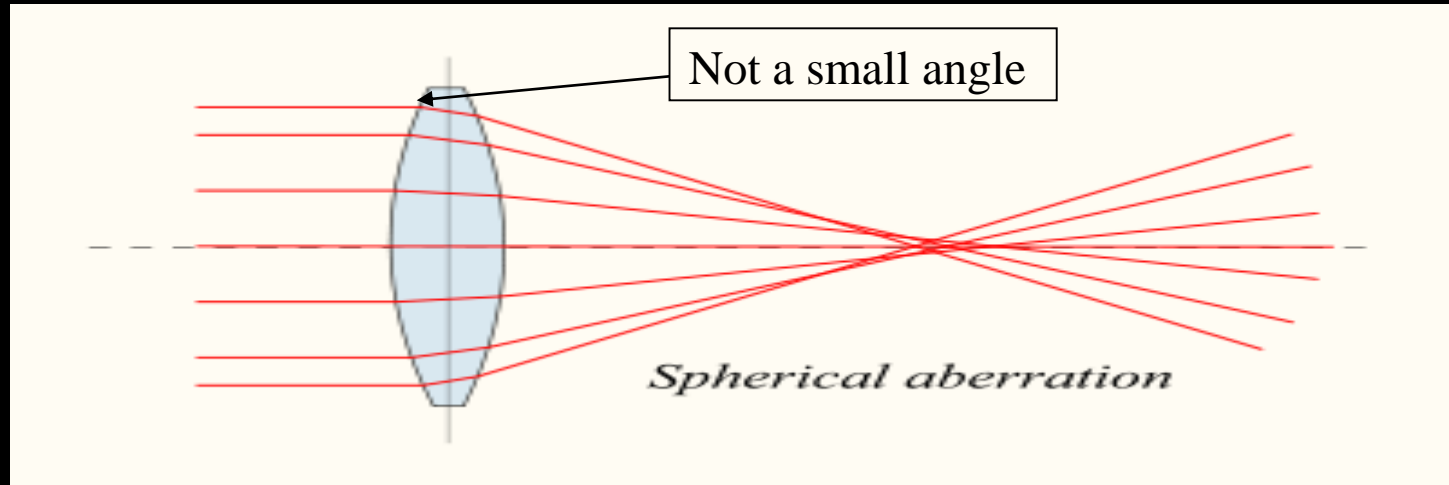
# Achromats



This lens mainly determines  $f$

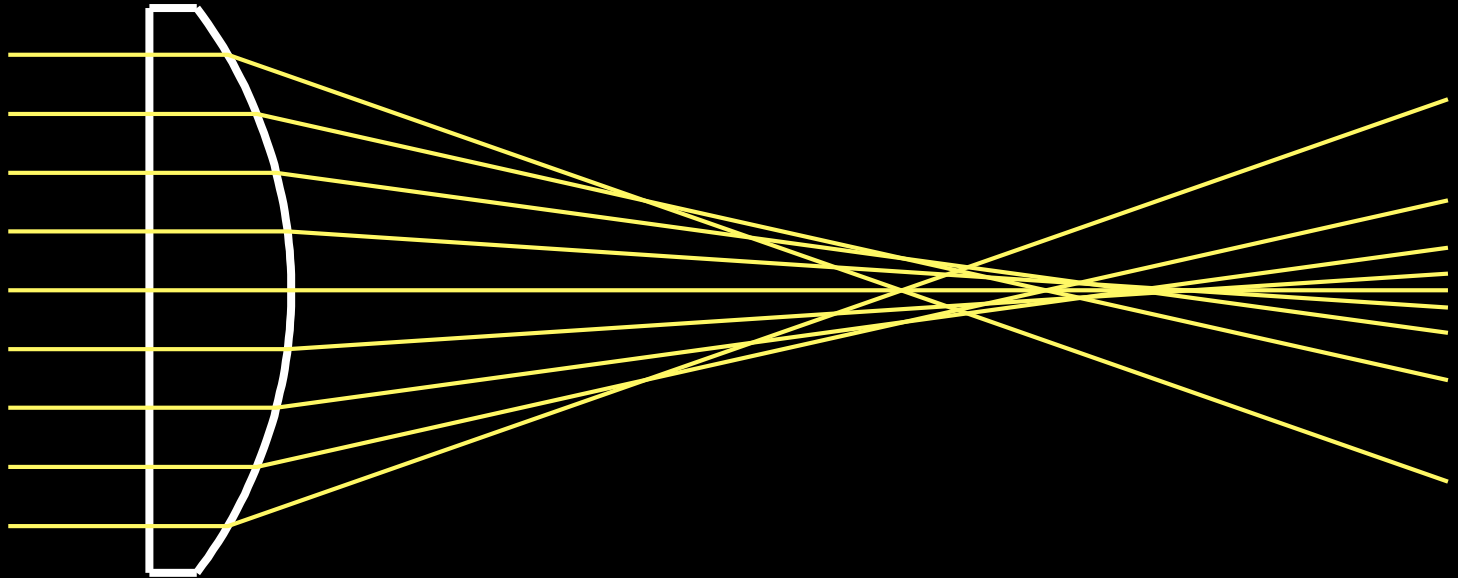
This lens mainly corrects dispersion

# Spherical Aberration



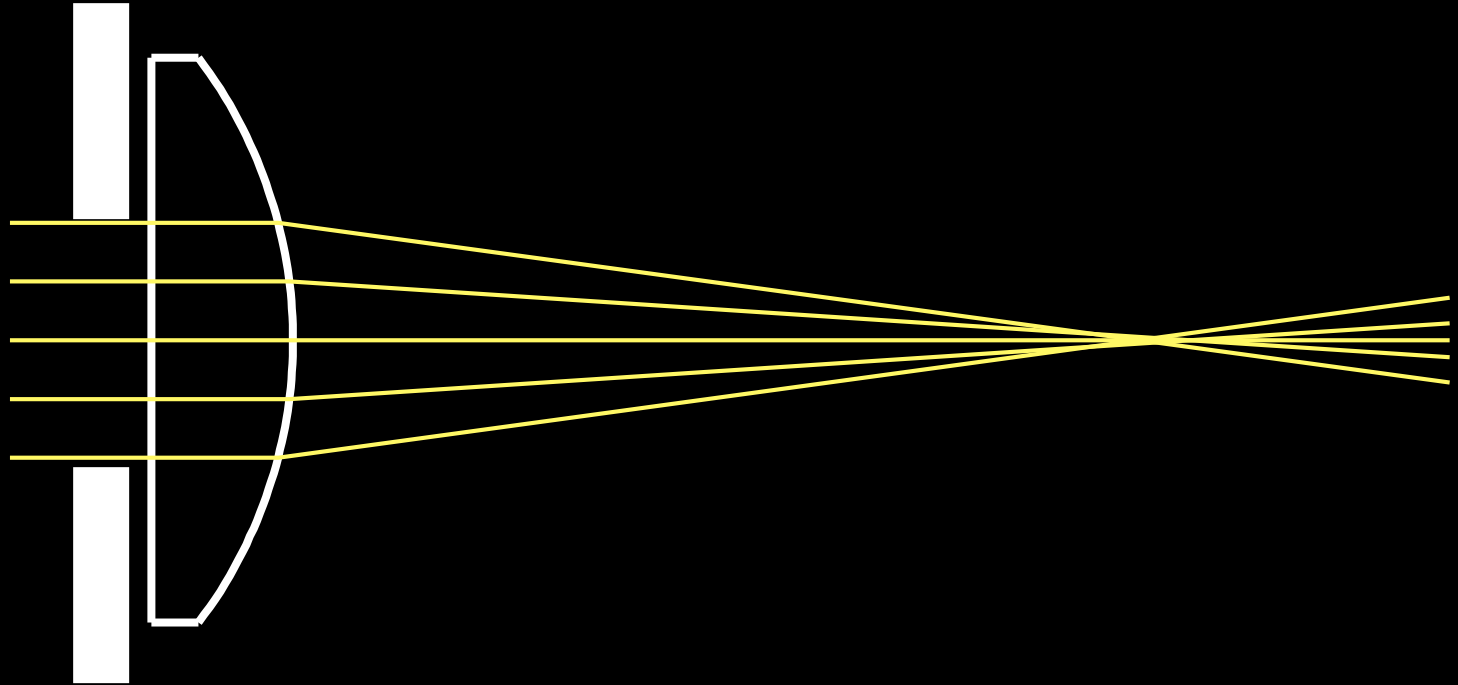
**Rays on the outside of the lens focus closer than rays on the inside of the lens**

# Spherical Aberration

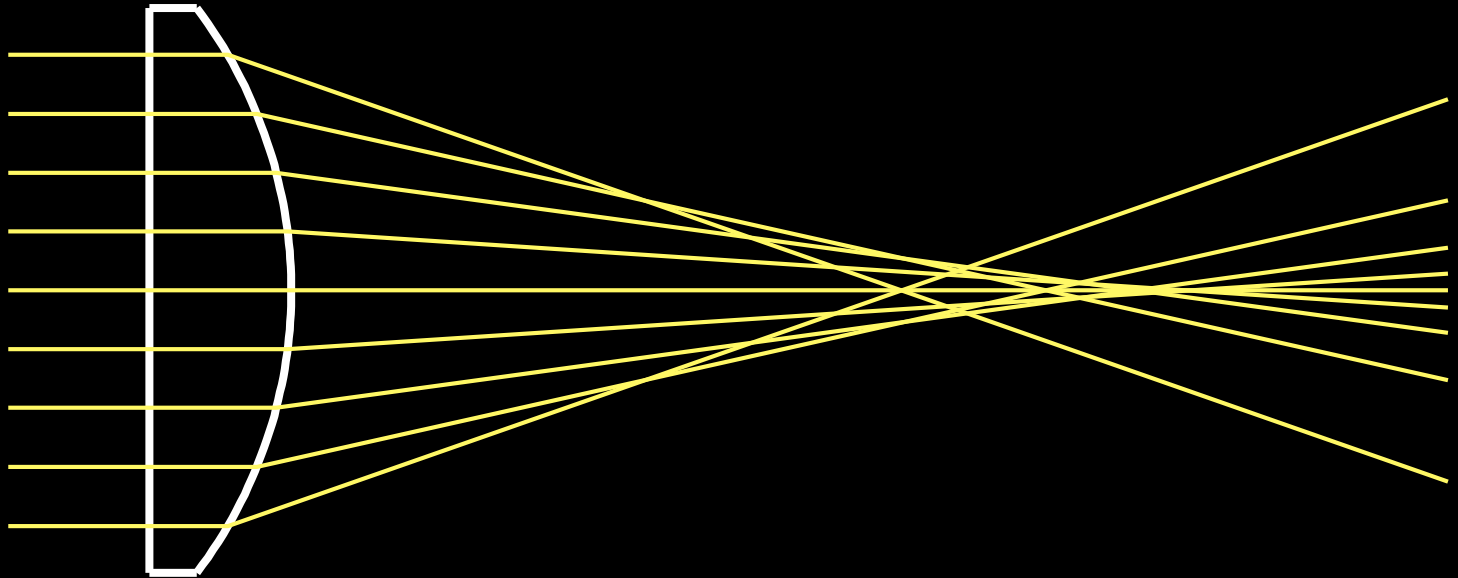


Credit: the next several slides were produced by Dr. Durfee

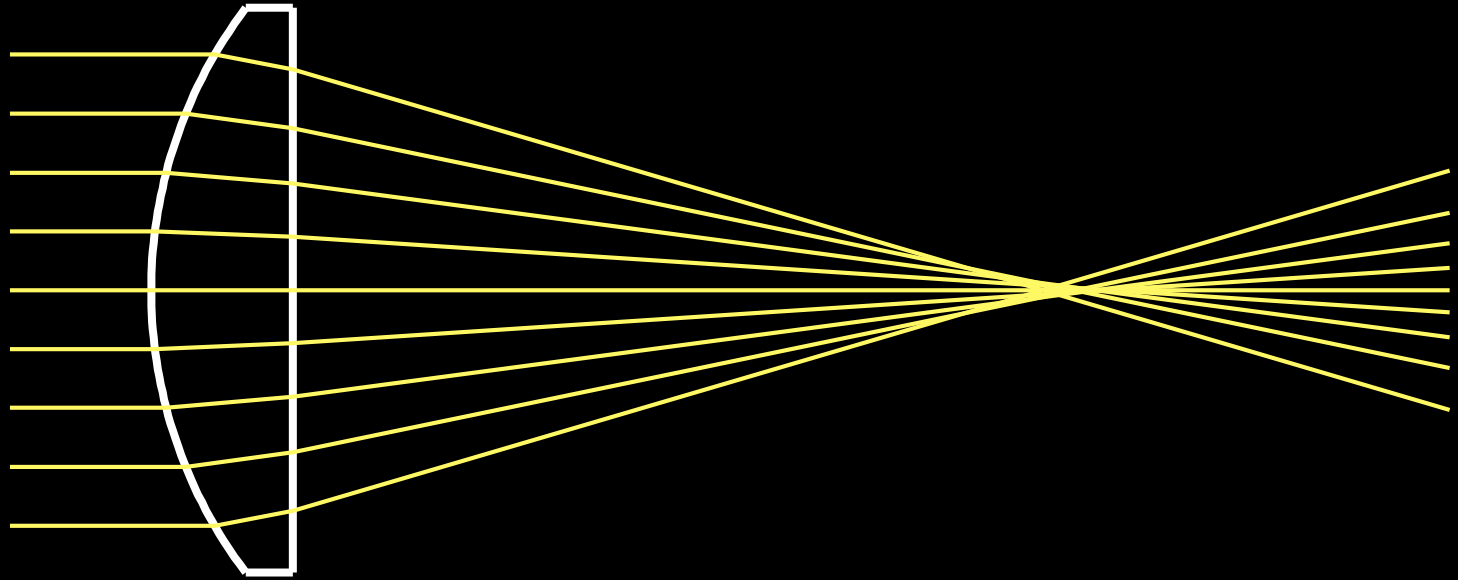
# Reducing Spherical Aberration with Aperture



# Spherical Aberration



# Reducing Spherical Aberration by Reversing Lens

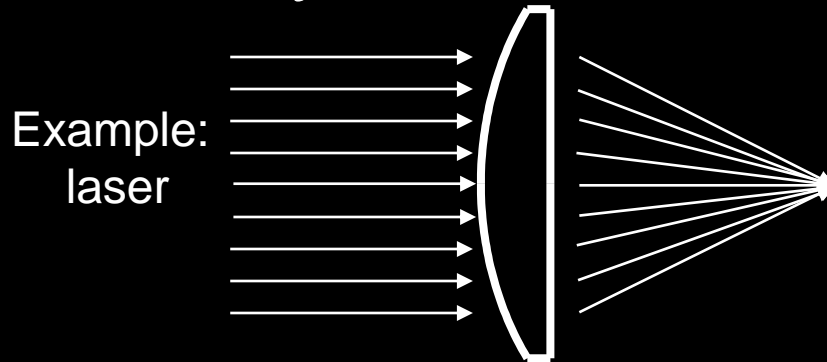


Curved side facing parallel rays

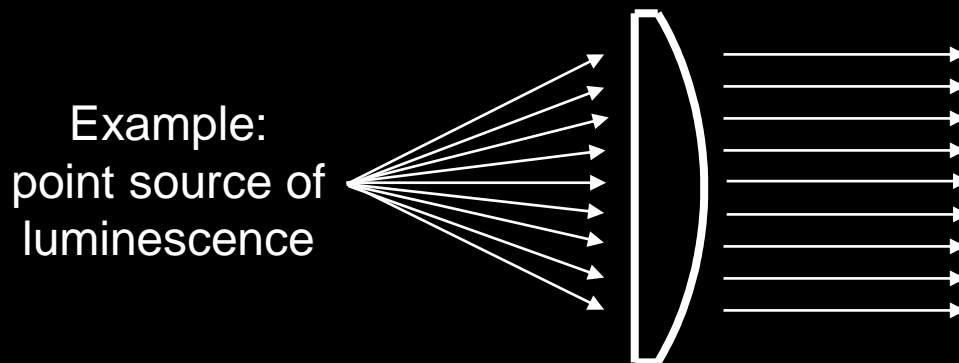


# Spherical Aberration Rules: “Flat to curved, curved to flat”

Parallel rays: curved side first



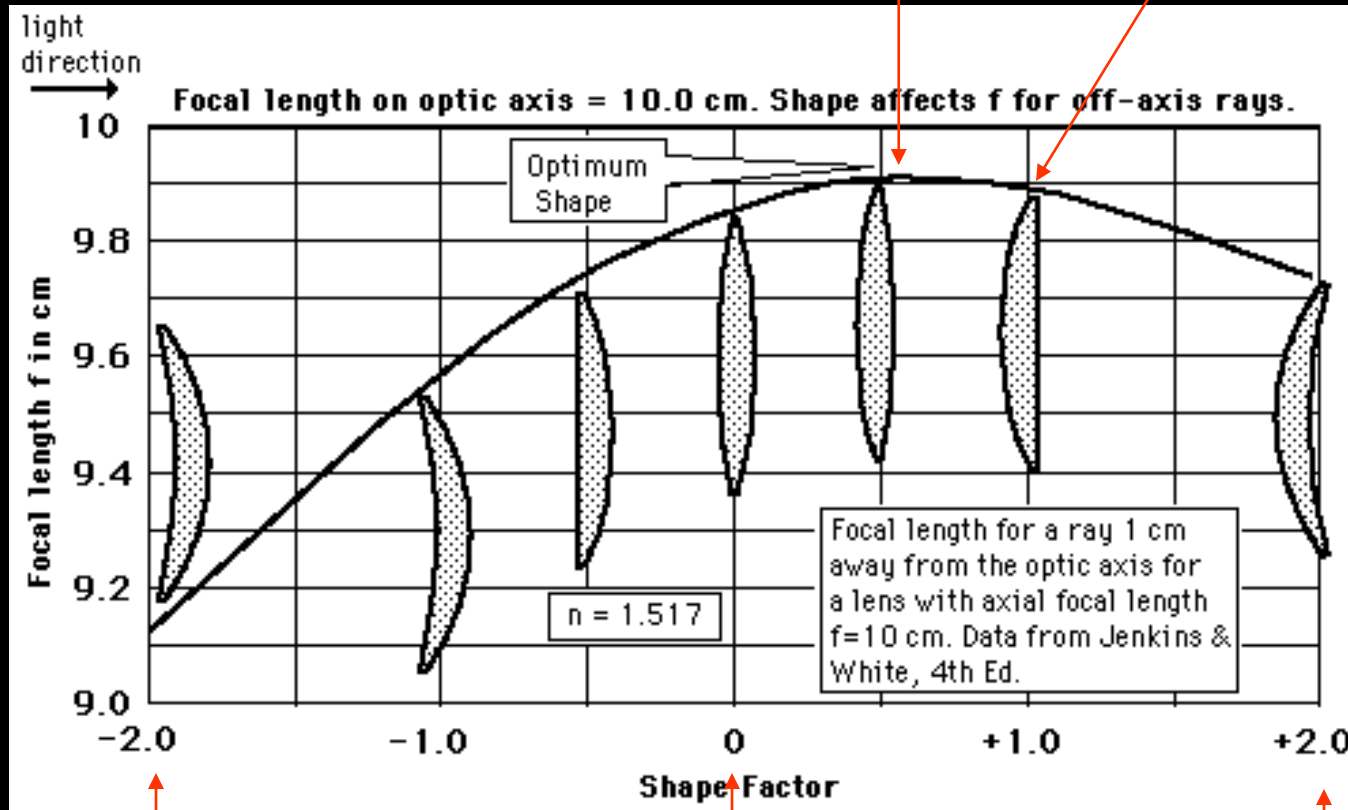
Diverging rays: flat side first



# Shape of lens

plano-convex:  
close to optimum  
when this direction

optimum



$R_2 = 1/3 R_1$

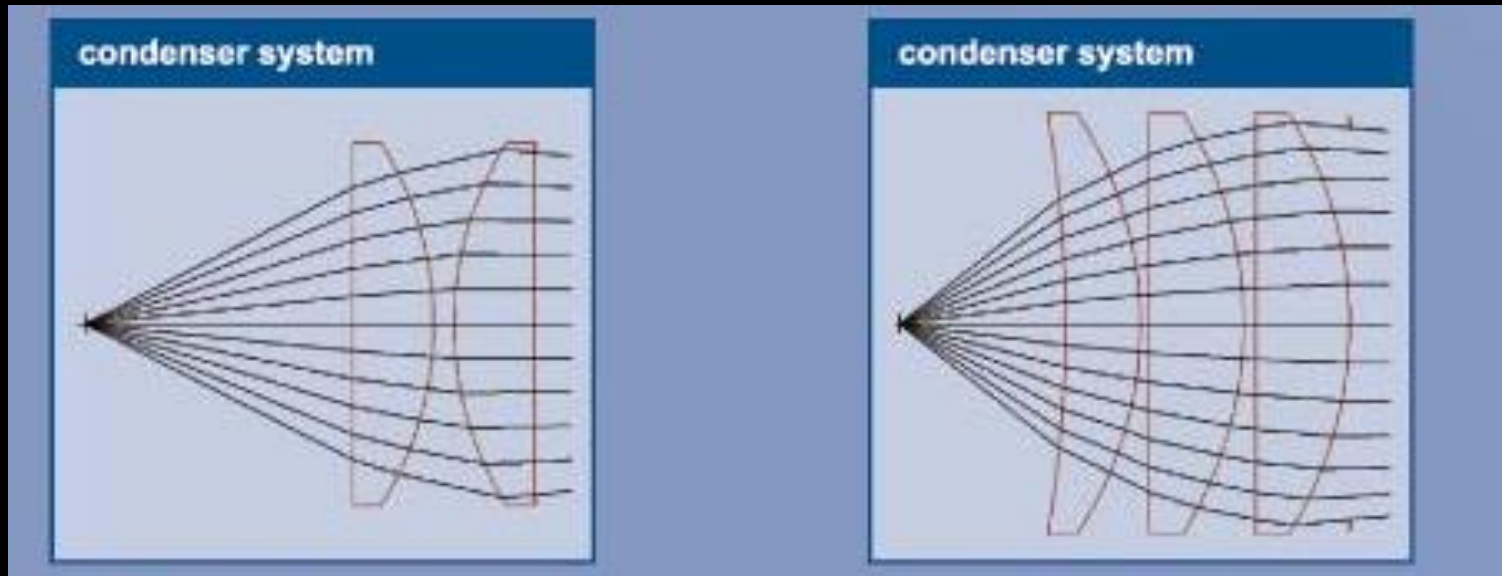
$R_2 = -R_1$

$R_2 = 3 R_1$

shape factor =  $(R_2 + R_1) / (R_2 - R_1)$

Acknowledgement: I got this from Dr. Hess, but I don't know where he got it from.

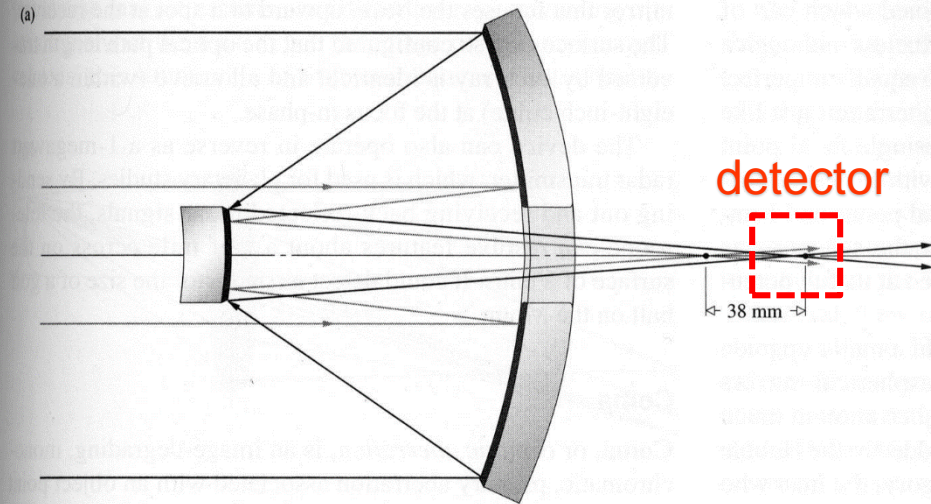
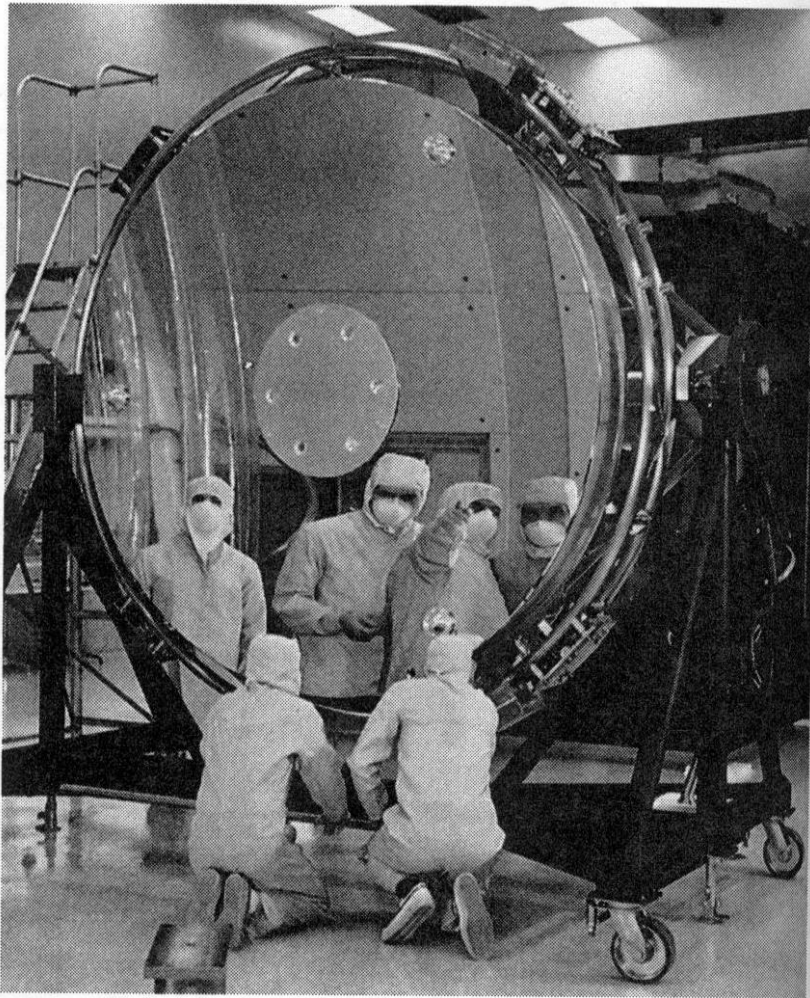
# Ray Tracing To Correct For Aberration



Acknowledgement: I got this from Dr. Hess, but I don't know where he got it from.

# The 1993 Hubble Telescope Repair

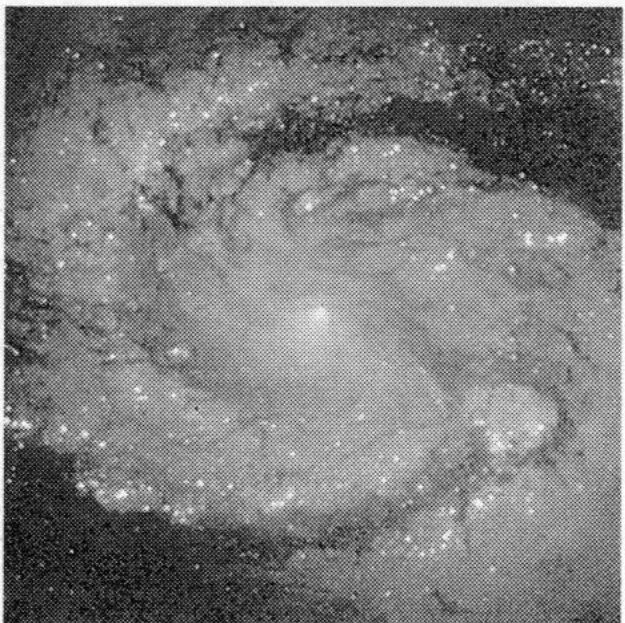
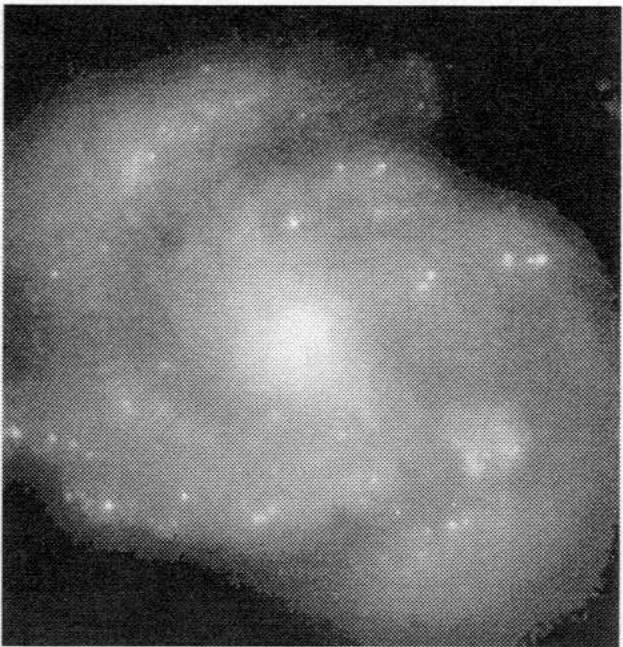
Pictures and  
story from Hecht



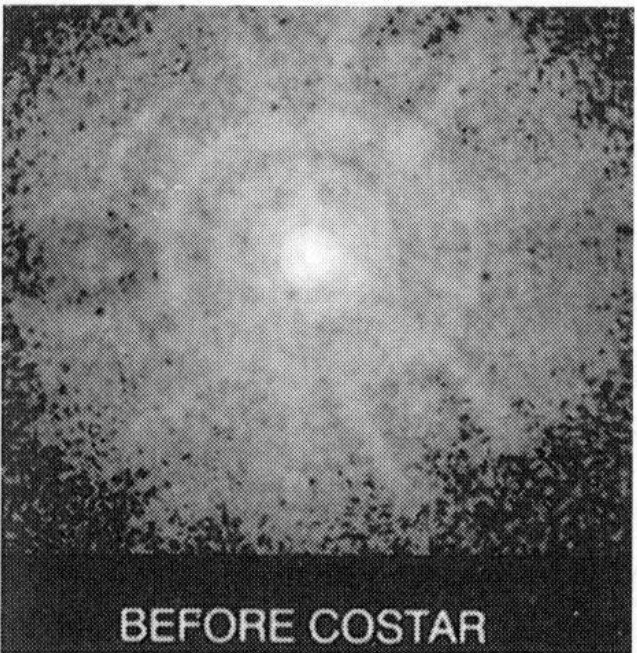
**Figure 6.19** (a) Because the primary mirror is too flat, rays from the outer edges met at a point 38 mm beyond the point where inner rays converge.

The 2.4-m-diameter hyperboloidal primary mirror of the Hubble Space Telescope. (Photo courtesy of NASA.)

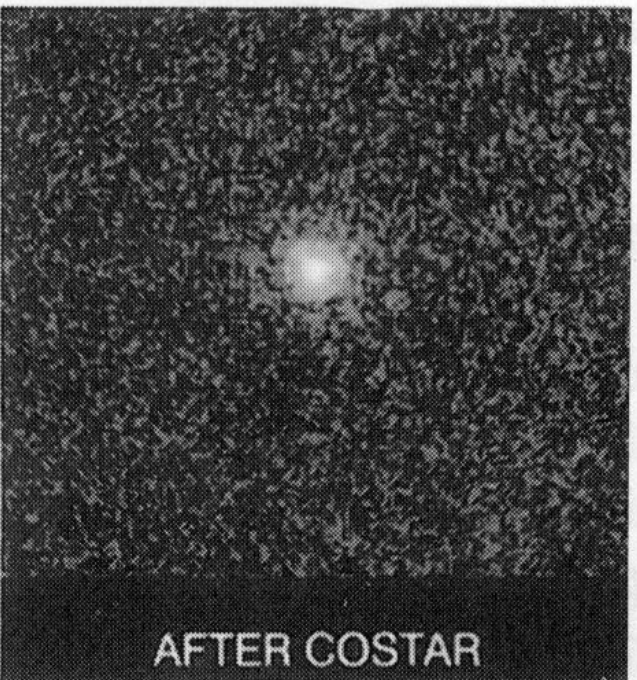




HST images of the M-100 galaxy with (before repair) and without (after repair) spherical aberration. (Photos courtesy of NASA.)



BEFORE COSTAR



AFTER COSTAR

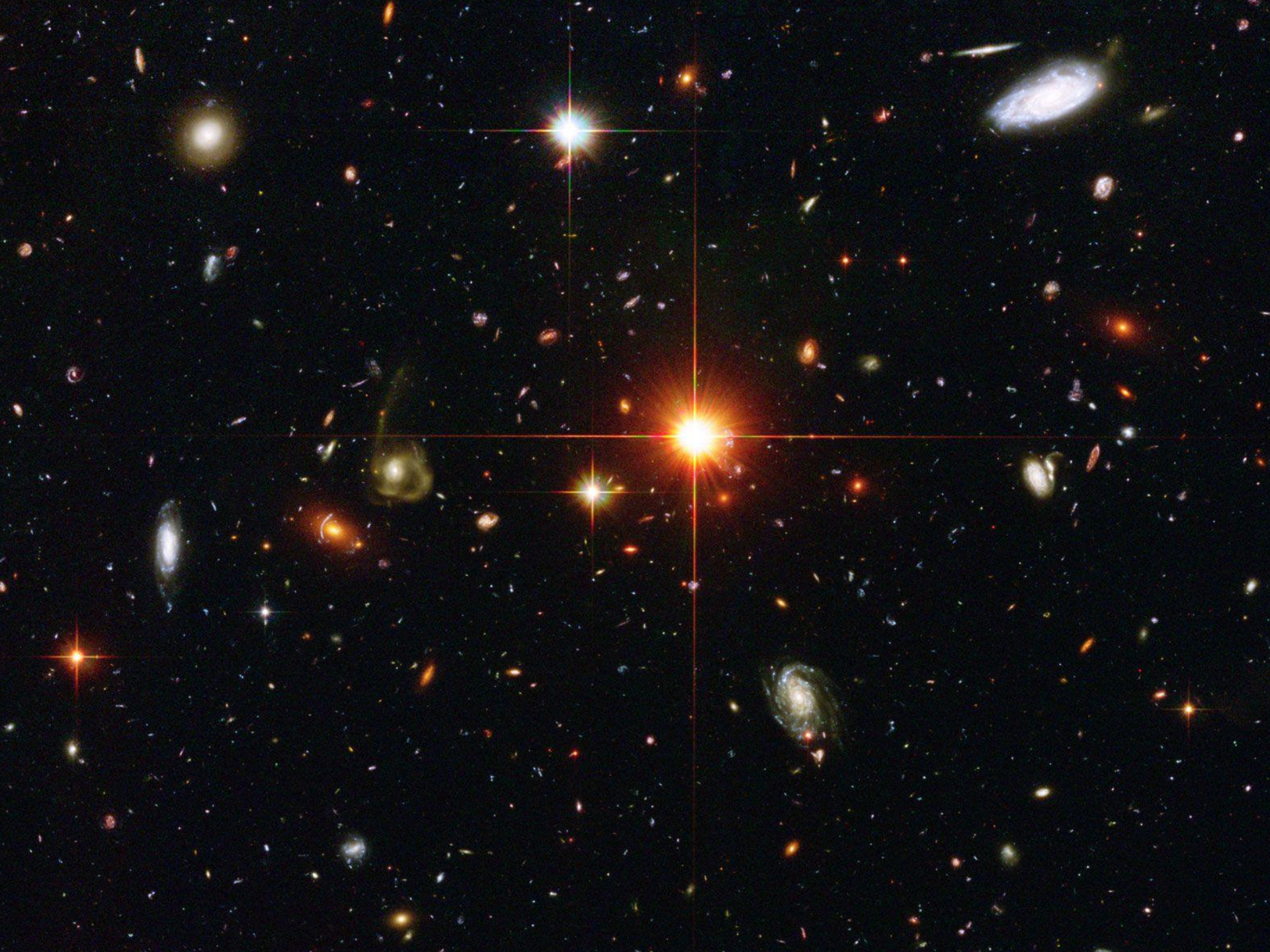














Betelgeuse

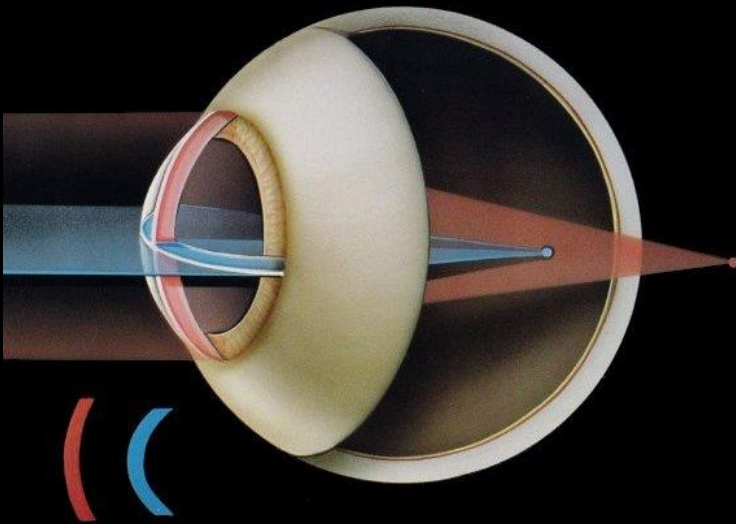


size of Jupiter's orbit

angular size: similar to  
resolving a car's  
headlights from 6,000  
miles away

# Astigmatism

- Lens shape = spherical + cylindrical: rays in different planes have different focal lengths



Viestenz et al. Zeitschrift der Deutschen Ophthalmologischen Gesellschaft, 104, 620-7 (2007).

Astigmatism	
Original	Compromise
Horizontal Focus	Vertical Focus

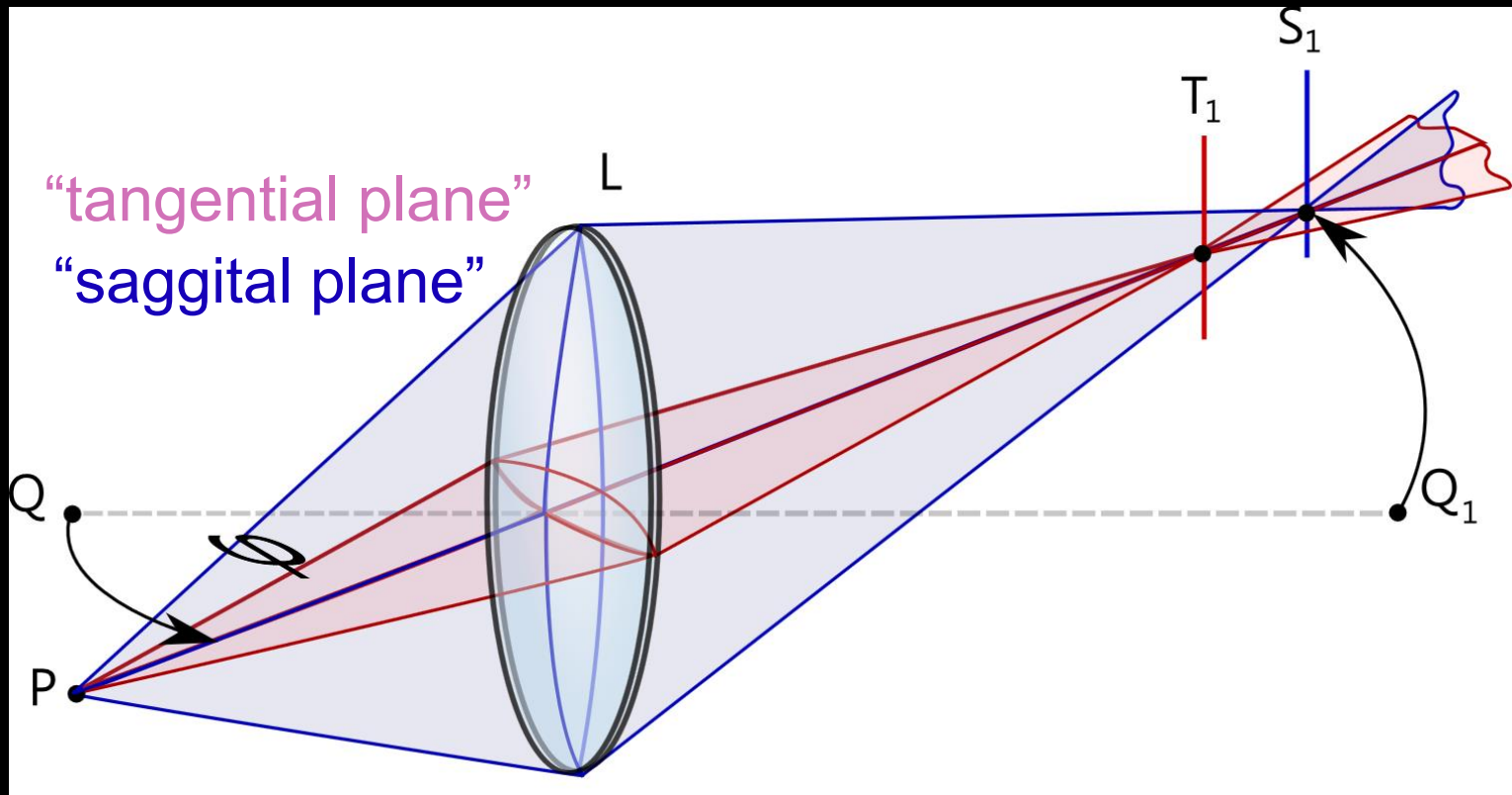
Blur from astigmatic lens at different distances

Wikipedia: "Astigmatism"

Correction: add an opposite cylindrical component in corrective lens

# Astigmatism, part 2

- Rays in different planes have different focal lengths



# Coma

from Hecht

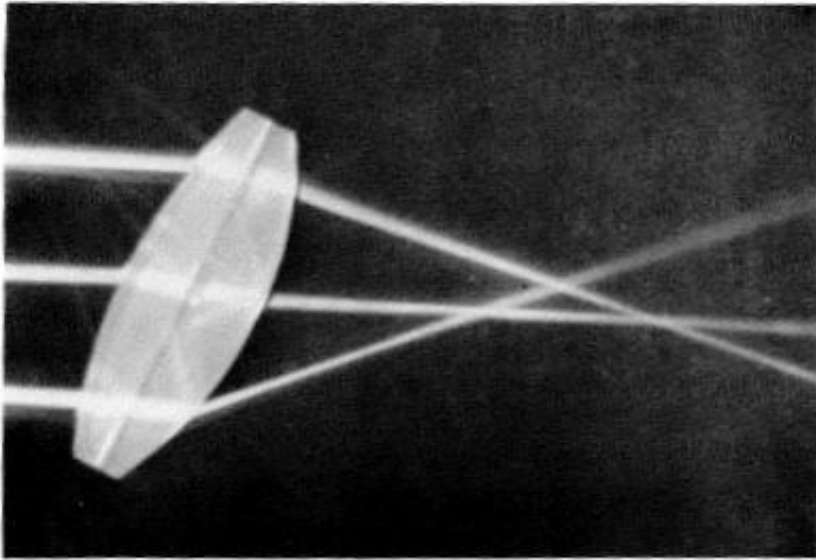
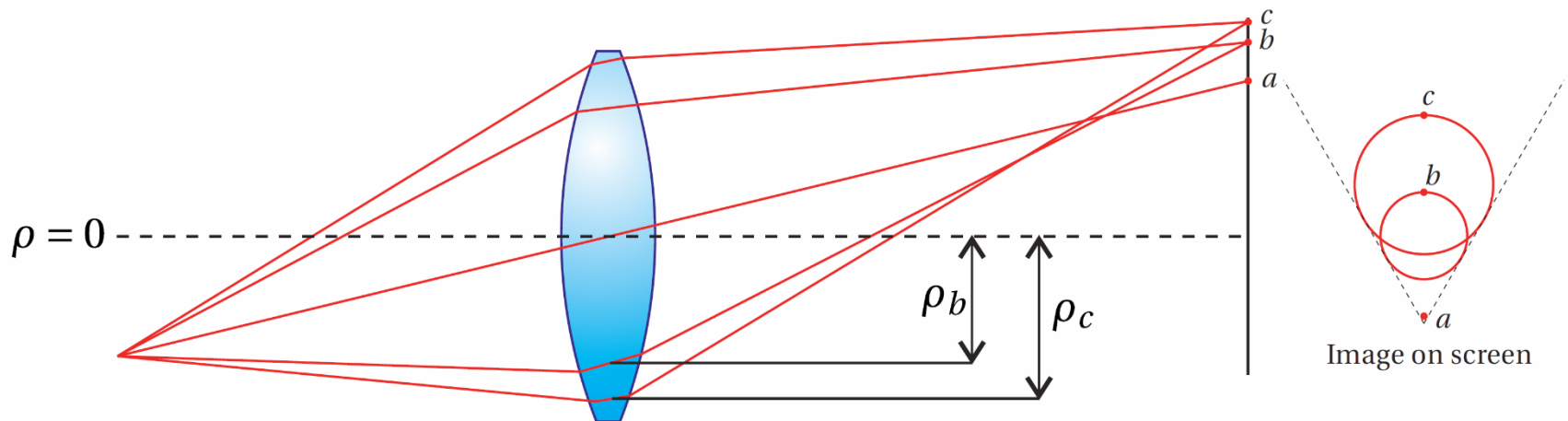


Fig. 6.18 Positive coma. [Photo by E.H.]

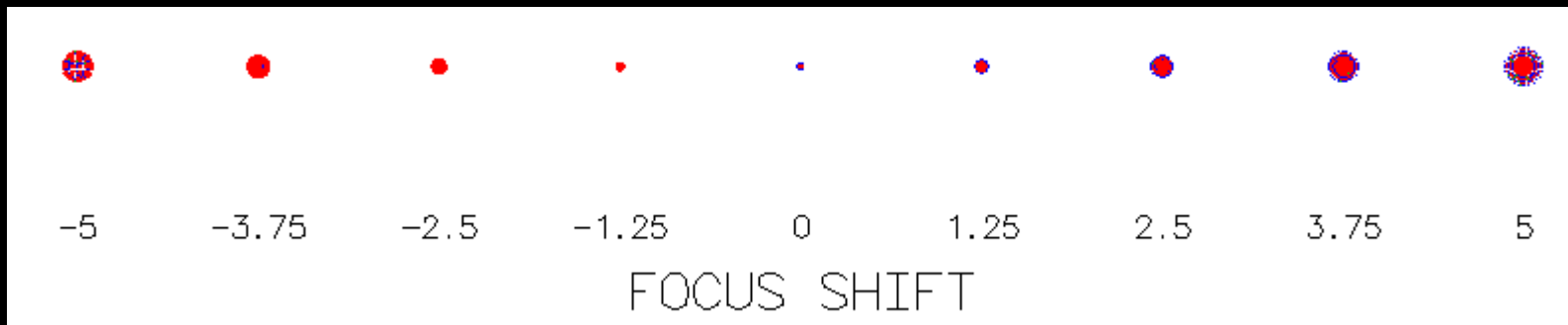
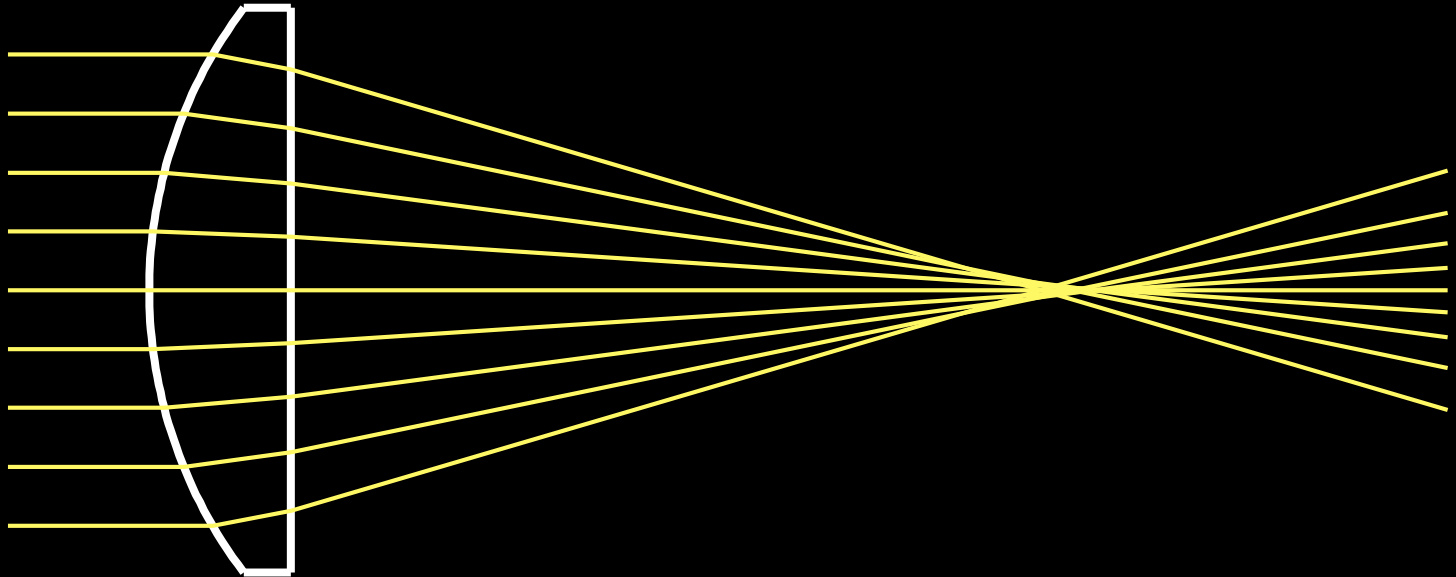
(b)

from P&W

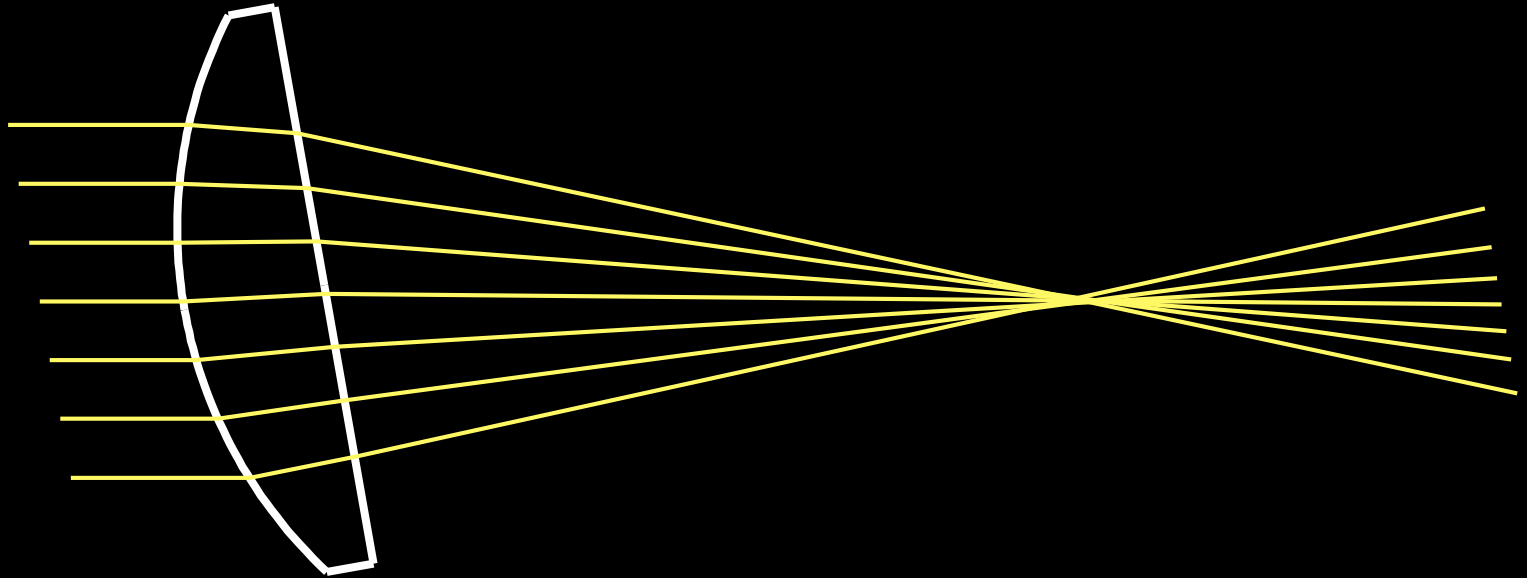


# 0 Degree Tilt

(next few slides from Dr. Durfee)



# 10 Degree Tilt



-5



-3.75



-2.5



-1.25



0



1.25



2.5



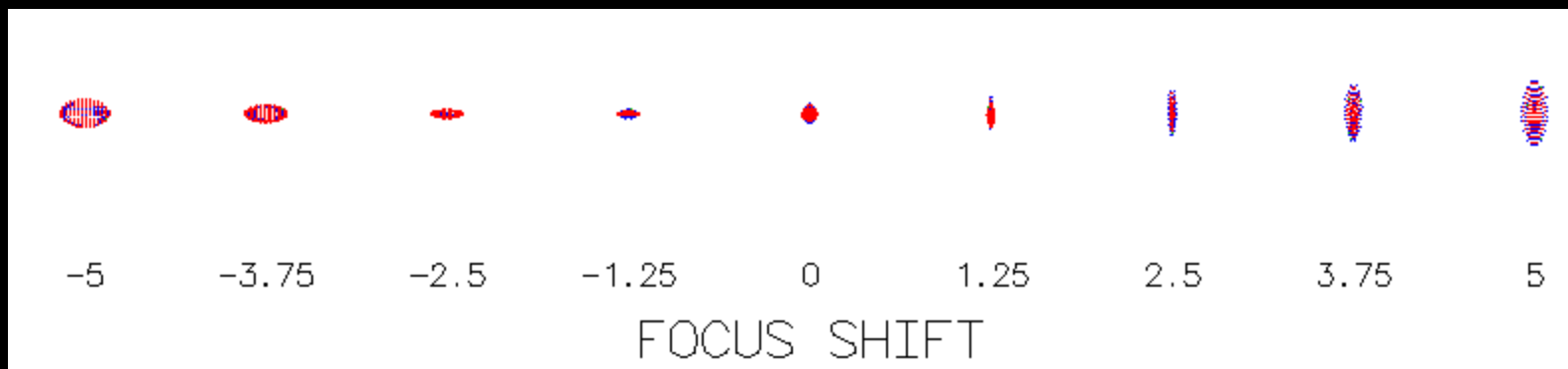
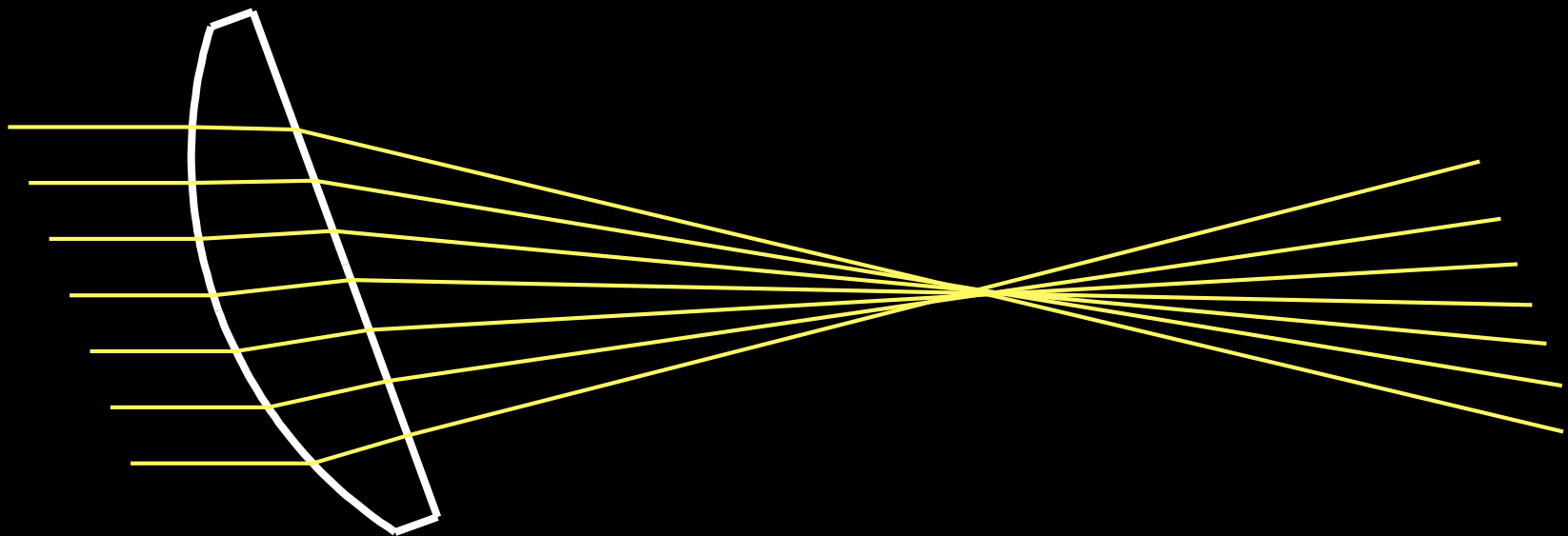
3.75



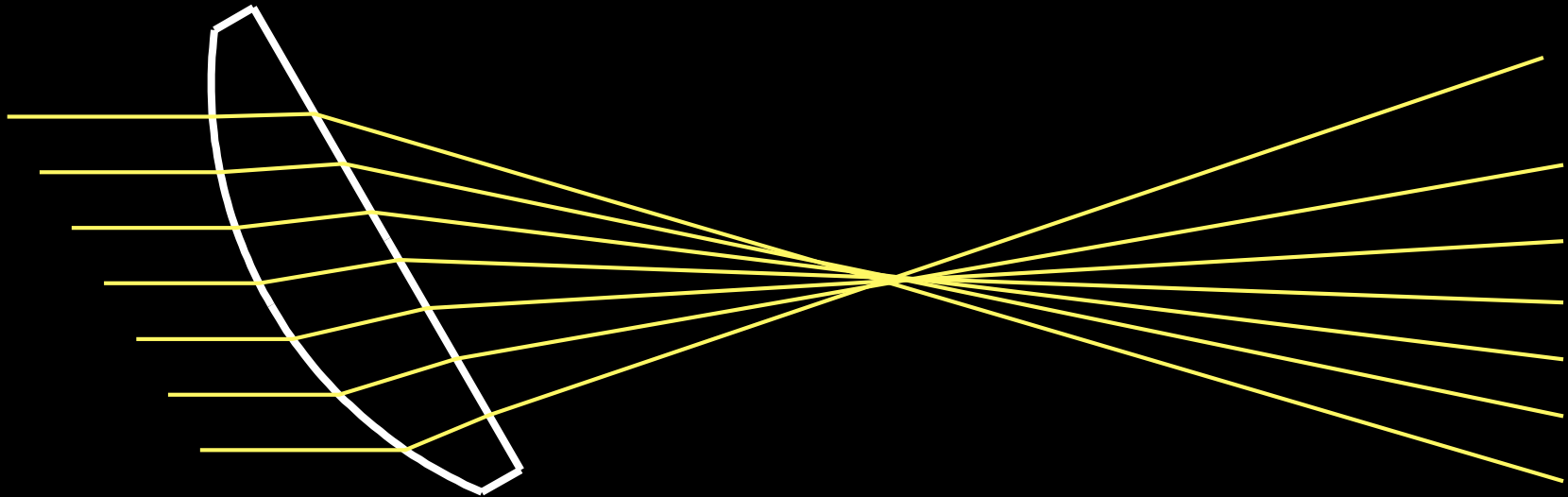
5

FOCUS SHIFT

# 20 Degree Tilt



# 30 Degree Tilt



-5



-3.75



-2.5



-1.25



0



1.25



2.5



3.75

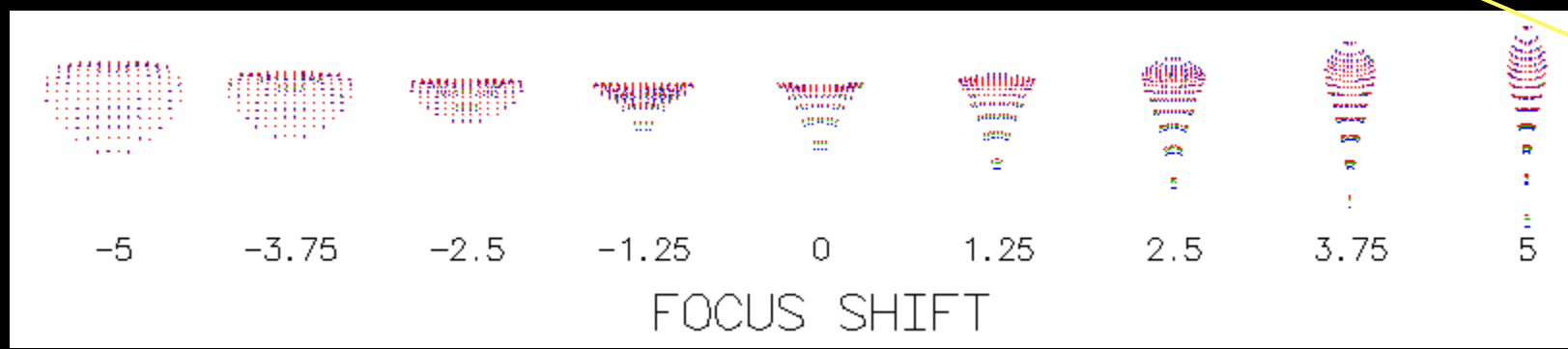
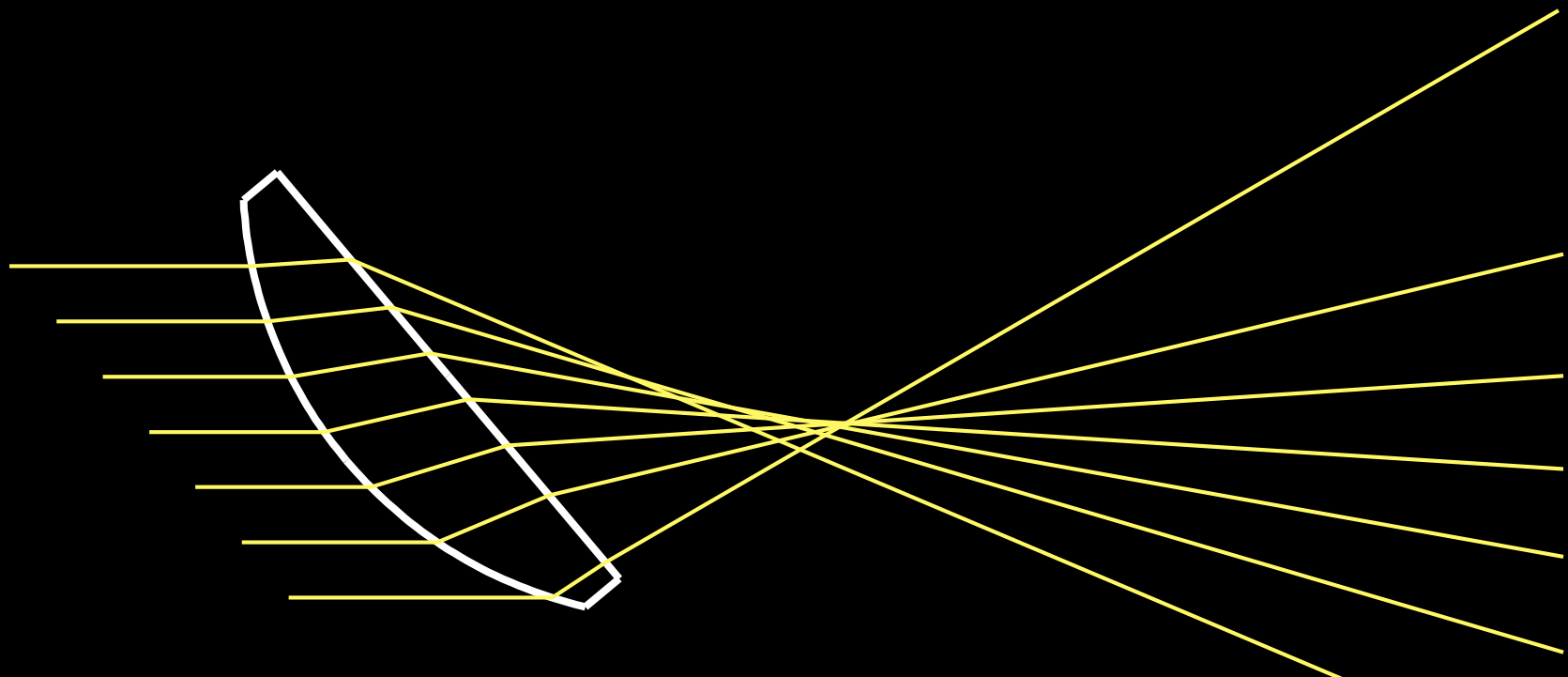


5

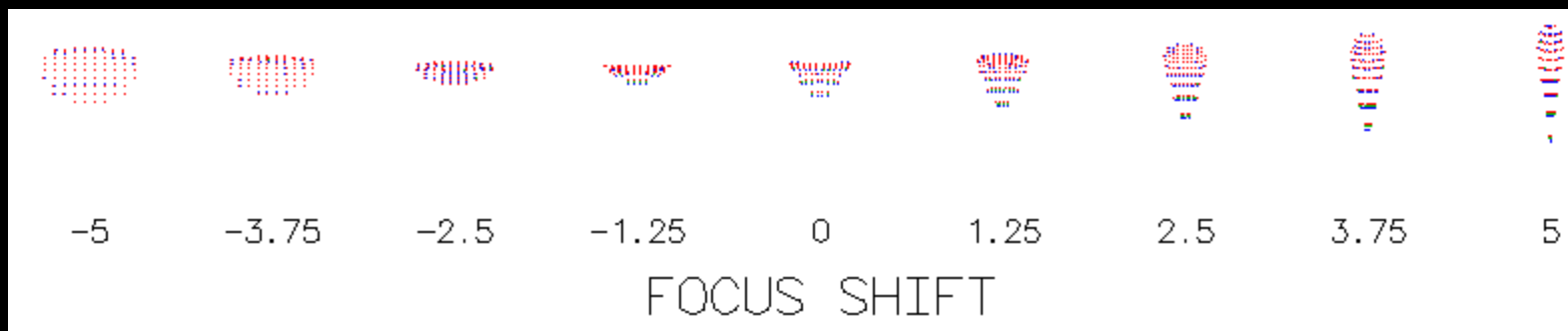
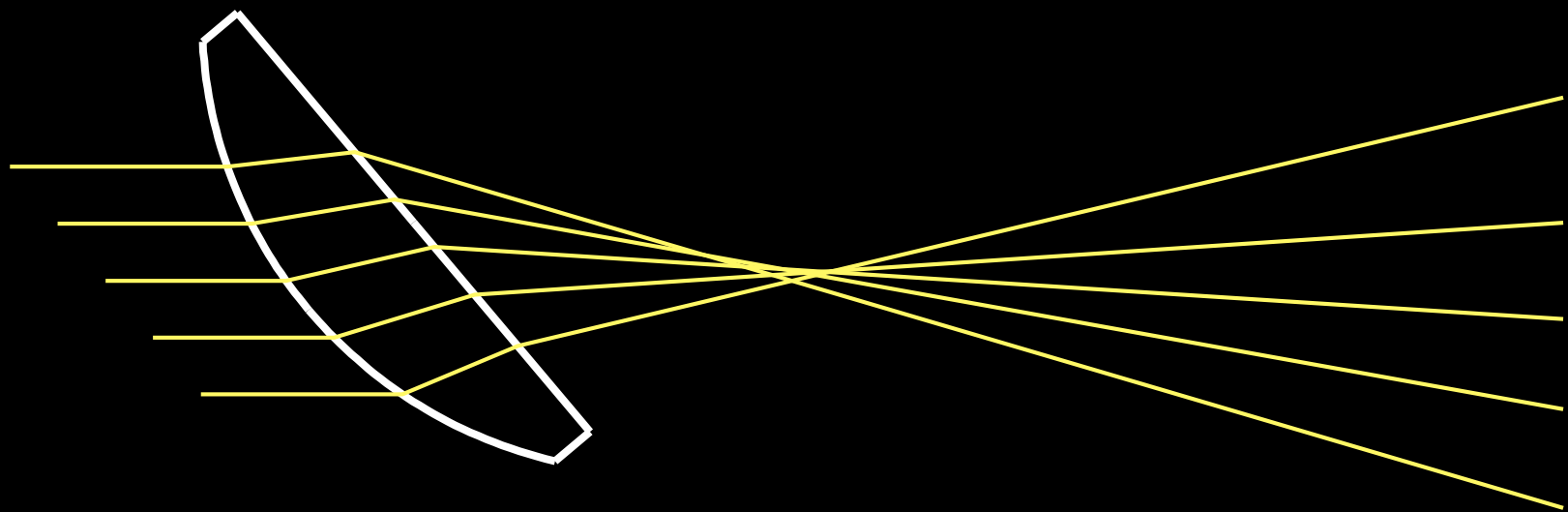
FOCUS SHIFT



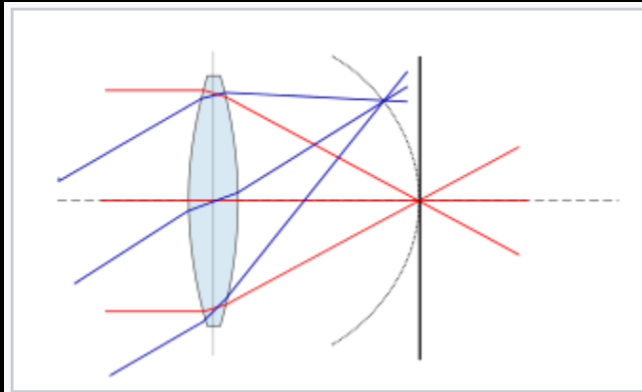
# 40 Degree Tilt



# 40 Degree Tilt + Aperture



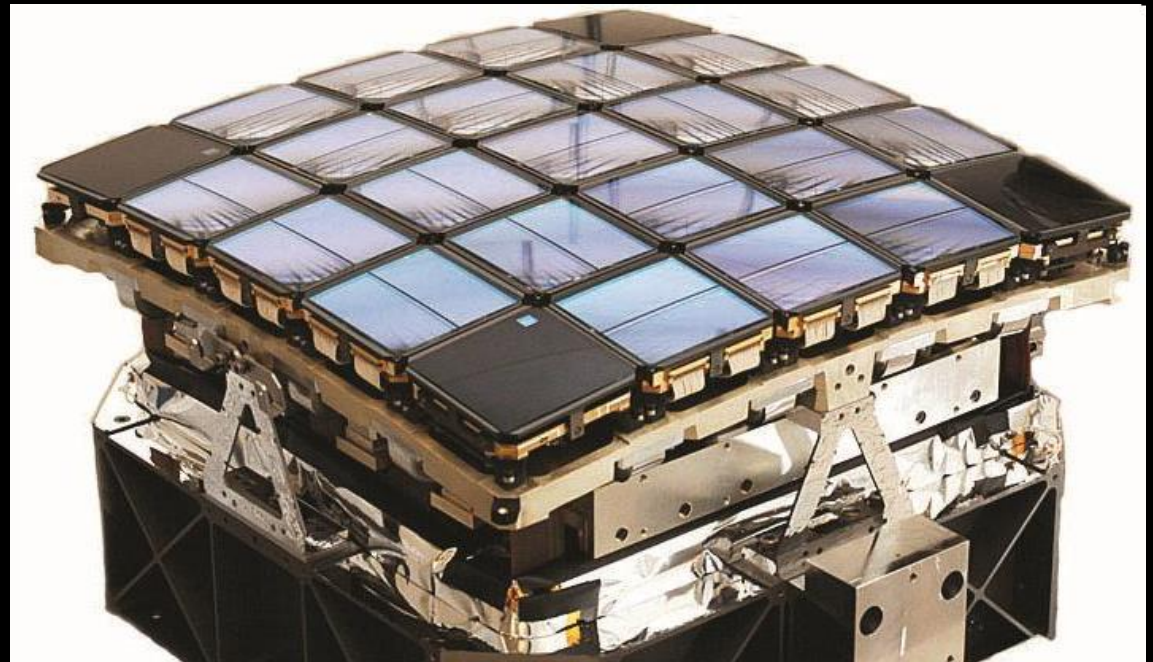
# Petzval field curvature



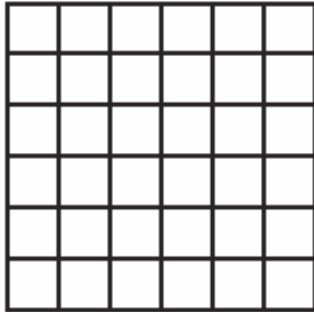
Field curvature: the image "plane" (the arc) deviates from a flat surface (the vertical line).

Usual solution: use multiple lenses to form one overall "lens" whose focal length increases with ray angle

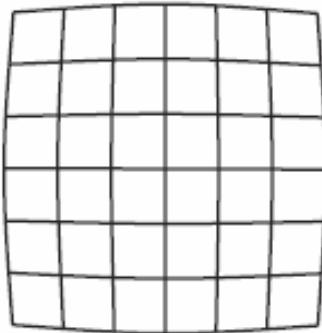
Solution 2: Curve your detector. This is detector on Kepler space telescope (searching for extra-solar planets)



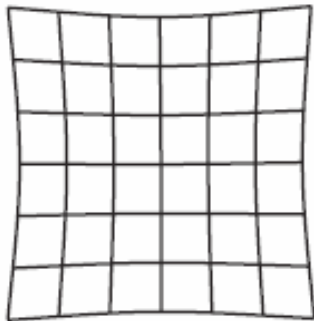
# Distortion



Undistorted



Barrel Distortion



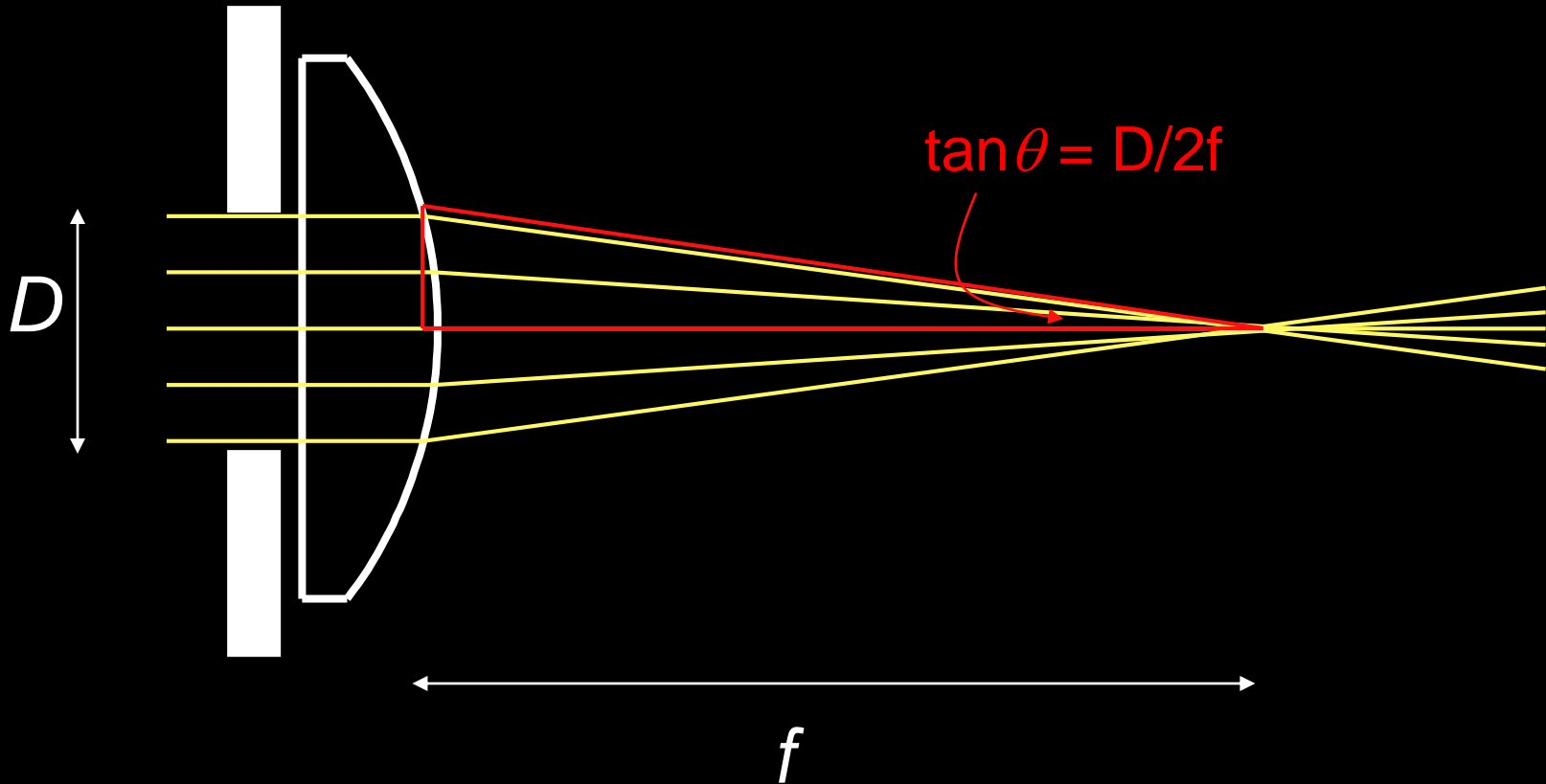
Pincushion Distortion

Far from center: magnification = less

Far from center: magnification = more

# Apertures

- f-number = "f/#" =  $f/D$



“Numerical aperture”:  $NA = \sin \theta$   
 $= D/2f$  for small angles  
( =  $1/(2f\text{-number})$  )