

Review of Example 1

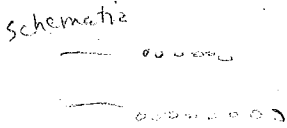
$T = 15K$

$E = 2 \cdot 10^{-23}$ — BF = .38 Prob = $\frac{.38}{1.38} = 28\%$

$E = 0$ — BF = 1 Prob = $\frac{1}{1.38} = 72\%$

continuing

(18) → what if we have 100 atoms? How many in ground state? in excited state?



for state E_i : $N_{ave} = N_{tot} \cdot \frac{BF_i}{BF_0 + BF_1 + BF_2 + \dots}$

(19) → What is average energy of the 100 atoms?

weighted average

$E_{ave} = E_0 P(0) + E_1 P(1)$

(20) → what if 2 states at $E = 2 \cdot 10^{-23} J$?

BF = $2 \times .38 = .76$
BF = 1

Prob $\frac{.76}{1.76} = 43\%$
 $\frac{1}{1.76} = 57\%$

Relationship to Phys 123 eqn $\Delta S = \int \frac{dQ}{T}$

of handout $\frac{dS}{dE} = \frac{1}{T}$
 $dS = \frac{dE}{T}$

$dS = \frac{dQ}{T}$

$dQ = T dS \rightarrow \frac{1}{T} \text{ actually} = \frac{dS}{dQ}$

1st Law $\Delta E_{int} = Q + W_{on}$

$\Delta E_{int} = \int T dS - \int P dV$

$dE = T dS - P dV$

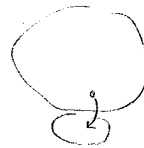
if $P dV \approx 0$, then $dE = T dS$

$\frac{1}{T} = \frac{dS}{dE}$ as in handout

We've overlooked two things which need to be discussed now

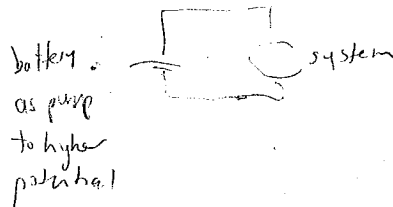
- ① The first law is incomplete. Sometimes particles being added/subtracted cause a change of energy.

Example: gravitational potential energy



particle entering system gives up gravitational potential energy, leading to more thermodynamic

Example: chemical potential energy



There are other more subtle examples, but because of similarity to the quantity used to characterize this is called the "chemical potential", μ .

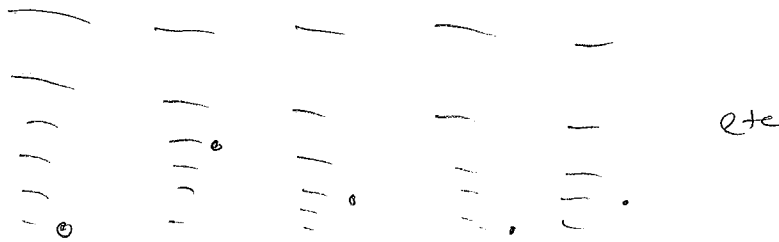
$$\mu = \text{change in energy due to a change in } N$$

First law:
$$dE_{int} = TdS + \underbrace{pdV}_{\substack{\text{still} \\ \text{consider} \\ \text{negligible}}} + \underbrace{\mu dN}_{\substack{\text{Must} \\ \text{account for}}}$$

Result: instead of BF, use "Gibbs factor"

$$P_s \sim e^{-(E_s - \mu N_s)/kT}$$

② We've assumed systems are non-interacting. I.e. their quantum states are independent
 actual



Drawn as



} shortcut notation, not really saying atom 1 has 4 electrons in ground state.

(why can't be possible?)

Pauli exclusion principle

If single atom, or if interacting atoms, then may have to worry about ↑

When do we have to worry about interacting?

one way to estimate for electrons:

$$p = \frac{h}{\lambda} \leftarrow \text{de Broglie wavelength}$$

$$KE = \frac{1}{2}mv^2 = \frac{p^2}{2m}$$

$$\rightarrow p = \sqrt{2mE}$$

$$\sqrt{2mE} = \frac{h}{\lambda}$$

$$\lambda = \frac{h}{\sqrt{2mE}}$$

KE from thermal effects: $E \sim kT$

$$\lambda = \frac{h}{\sqrt{2mkT}}$$

add. factor of $\frac{1}{\sqrt{\pi}}$ for no apparent reason

$$\lambda_Q = \frac{h}{\sqrt{2\pi mkT}} = \text{"quantum length"}$$

for electron at room T, $\lambda_Q = 4.3 \text{ nm}$

What is separation of atoms in solid? much smaller. Interactions!

day 17 pg 4

How to handle interacting systems — states are ^{states of} combined system, not isolated atoms

Two types: Fermions — obey Pauli exclusion
ex: electrons, protons, neutrons

Bosons — don't obey Pauli exclusion
ex: photons, phonons

← focus on this state. what is N_{ave} ?

("reservoir" = all other states)

$$N_{ave} = N_0 P(0) + N_1 P(1) + N_2 P(2) + \dots$$

like Erc equation

Fermions

can only have 0 or 1

$$N_{ave} = 0 \cdot P(0) + 1 \cdot P(1)$$

$$= \frac{1 \cdot e^{-(E-\mu)/kT}}{1 + e^{-(E-\mu)/kT}}$$

$$\times \frac{e^{+(E-\mu)/kT}}{e^{-(E-\mu)/kT}}$$

$$N_{ave} = \frac{1}{e^{(E-\mu)/kT} + 1}$$

Fermi-Dirac distribution often: f_{FD}

Note: this also just = probability of that state being occupied, since occupancy is either 0 or 1 (this case only!)

Bosons

can have any number

$$N_{ave} = 0 P(0) + 1 P(1) + 2 P(2) + 3 P(3) + \dots$$

$$= 0 + \frac{1 e^{-(E-\mu)/kT}}{1 + e^{-(E-\mu)/kT} + e^{-2(E-\mu)/kT} + \dots} + \frac{2 e^{-2(E-\mu)/kT}}{1 + e^{-(E-\mu)/kT} + e^{-2(E-\mu)/kT} + \dots} + \dots$$

Let $x = e^{-(E-\mu)/kT}$

$$N_{ave} = \frac{0 + x + 2x^2 + 3x^3 + \dots}{1 + x + x^2 + x^3 + \dots}$$

claim: that $= \frac{x}{1-x}$

proof = cross multiply

$$(1-x) \left(x + 2x^2 + 3x^3 + \dots \right) \stackrel{?}{=} x \left(1 + x + x^2 + x^3 + \dots \right)$$

$$\begin{array}{r} x + 2x^2 + 3x^3 + \dots \\ -x^2 - 2x^3 - 3x^4 - \dots \\ \hline \end{array}$$

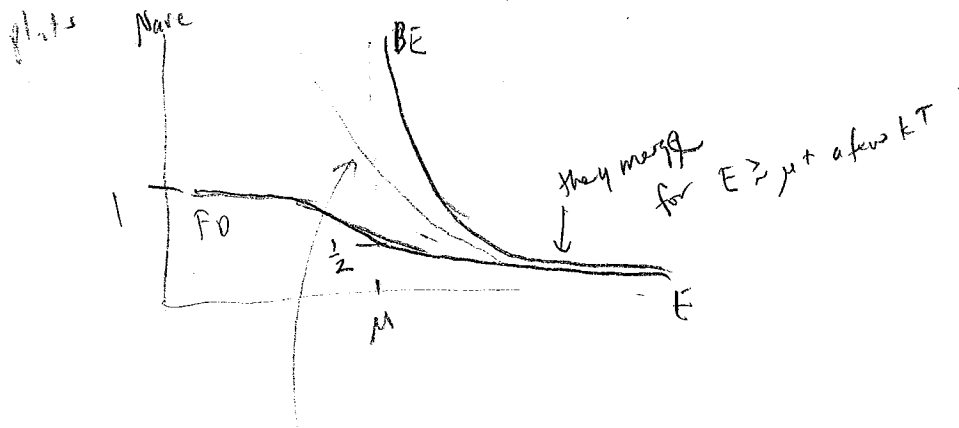
$x + x^2 + x^3 + x^4 + \dots$ ✓ yes it's equal

$$N_{ave} = \frac{e^{-(E-\mu)/kT}}{1 - e^{-(E-\mu)/kT}} \times \frac{e^{+C}}{e^{+C}}$$

$$N_{ave} = \frac{1}{e^{(E-\mu)/kT} - 1}$$

Bose-Einstein distribution often: f_{B-E}

small difference in sign \leftrightarrow big difference in behavior



Interim. Turns out that even though we neglected μ in BE discussion, you can calculate μ (hard, see Schroder pg 268)

$$N_{ave} = \frac{1}{e^{(E-\mu)/kT}}$$

$$N_{ave} = N_{tot} \frac{DF(E)}{\sum_{all\ DFs}}$$

$$= e^{-(E-\mu)/kT}$$

All 3 same when $E \gg \mu$, when $N_{ave} \ll 1$

"classical" regime

$\mu = -kT \ln \left(\frac{N_{tot}}{g} \right)$